

S. Riggi for the SKA Dish LMC team (*)
INAF - Osservatorio Astrofisico di Catania, Italy

The Square Kilometer Array (SKA): the world's largest and most sensitive radio observatory

- ❖ **2 radio antenna arrays** (Fig. 1)
 - SKA1-Mid: South Africa
 - SKA1-Low: Western Australia
- ❖ **2 construction phases foreseen**
 - Phase 1 (2018-2024): SKA1-Mid + SKA1-Low arrays
 - Phase 2 (mid 2020s): expansions of both arrays in Southern Africa and Australia
- ❖ **SKA pre-construction phase 1 ongoing**
 - Lead by 8 major consortia coordinated by the SKA Organization
 - 10 member countries, >100 research institutions and companies across 20 countries
 - Consortia completing their critical design

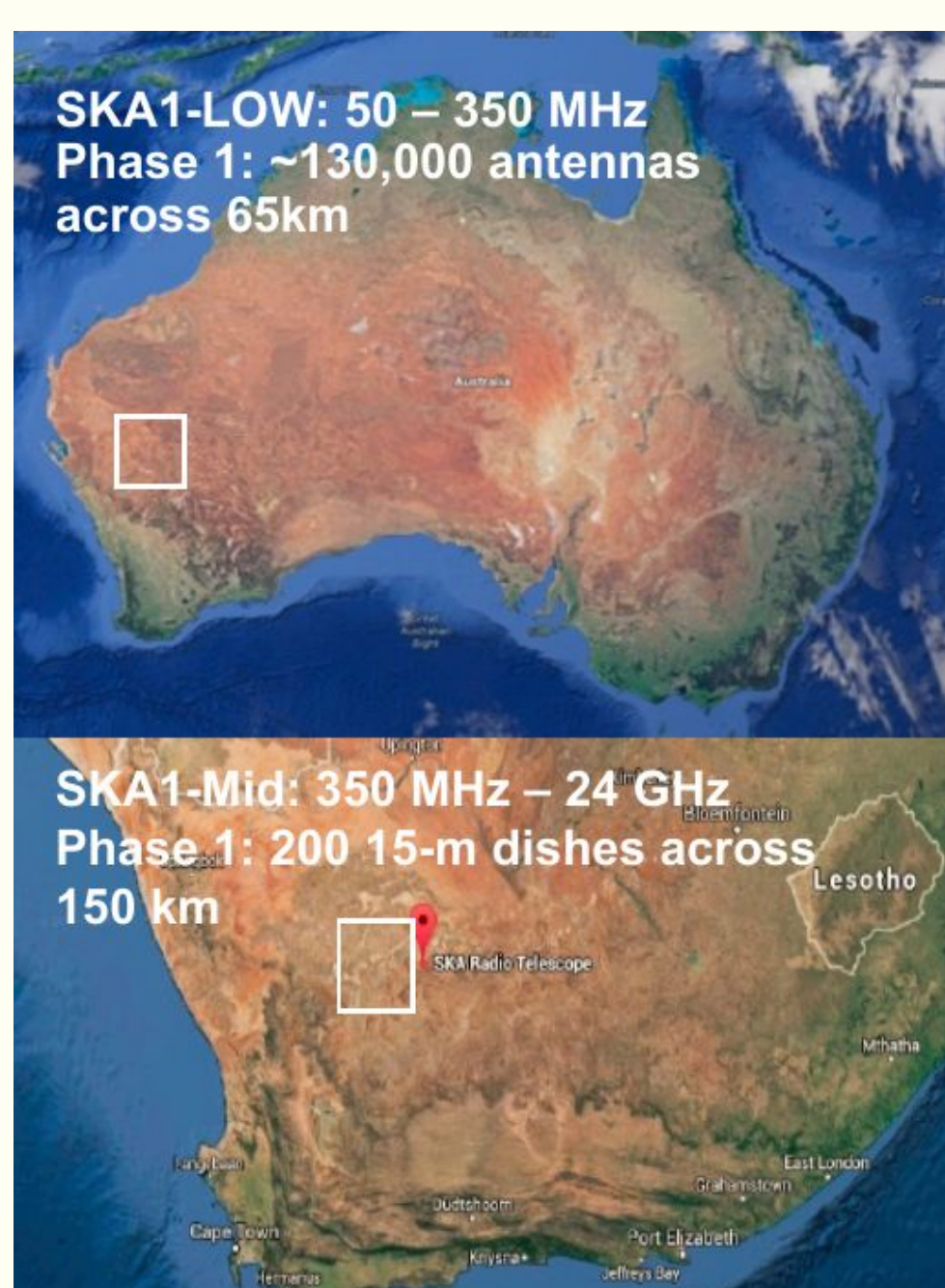


Fig. 1: SKA1-Low (top) and Mid (bottom) locations

SKA1-Mid array layout (Fig. 2)

- 130 15-m diameter dish antennas (+64 Meerkat dishes)
- Extension: 120 km
- Observing frequencies: 0.35-14 GHz
- Phase 1: band 1, 2 in each dish, band 5 in 67 dishes
- Extension to ~2500 antennas in SKA 2



Fig. 2: Pictorial view of SKA1-Mid array (left) and dish antenna (right).
Image credits: SKAO

SKA DISH OVERVIEW

Dish instrumentation (Fig. 3) mounted at 3 major locations (indexer, pedestal, yoke) and designed by:

- **Dish Structure (DS):** antenna structure and optics, feed indexer, servo systems, power distribution and safety systems.
- **Single-Pixel Feed (SPF):** feed packages (OMTs, LNAs), feed helium cooling and vacuum system and relative controllers.
- **Receiver (SPFRx):** RF digitizer system and relative controllers.

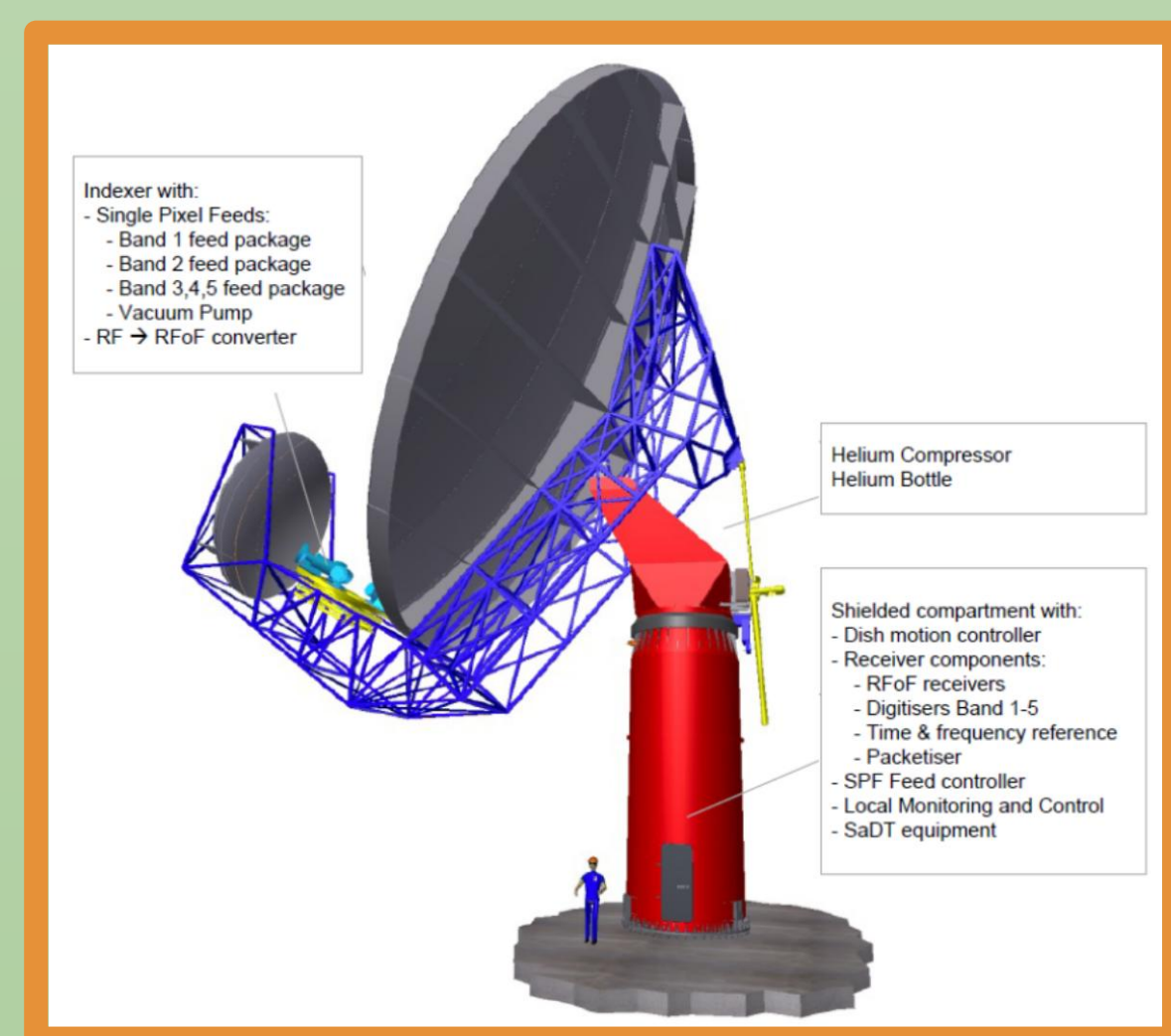


Fig. 3: Dish instrumentation

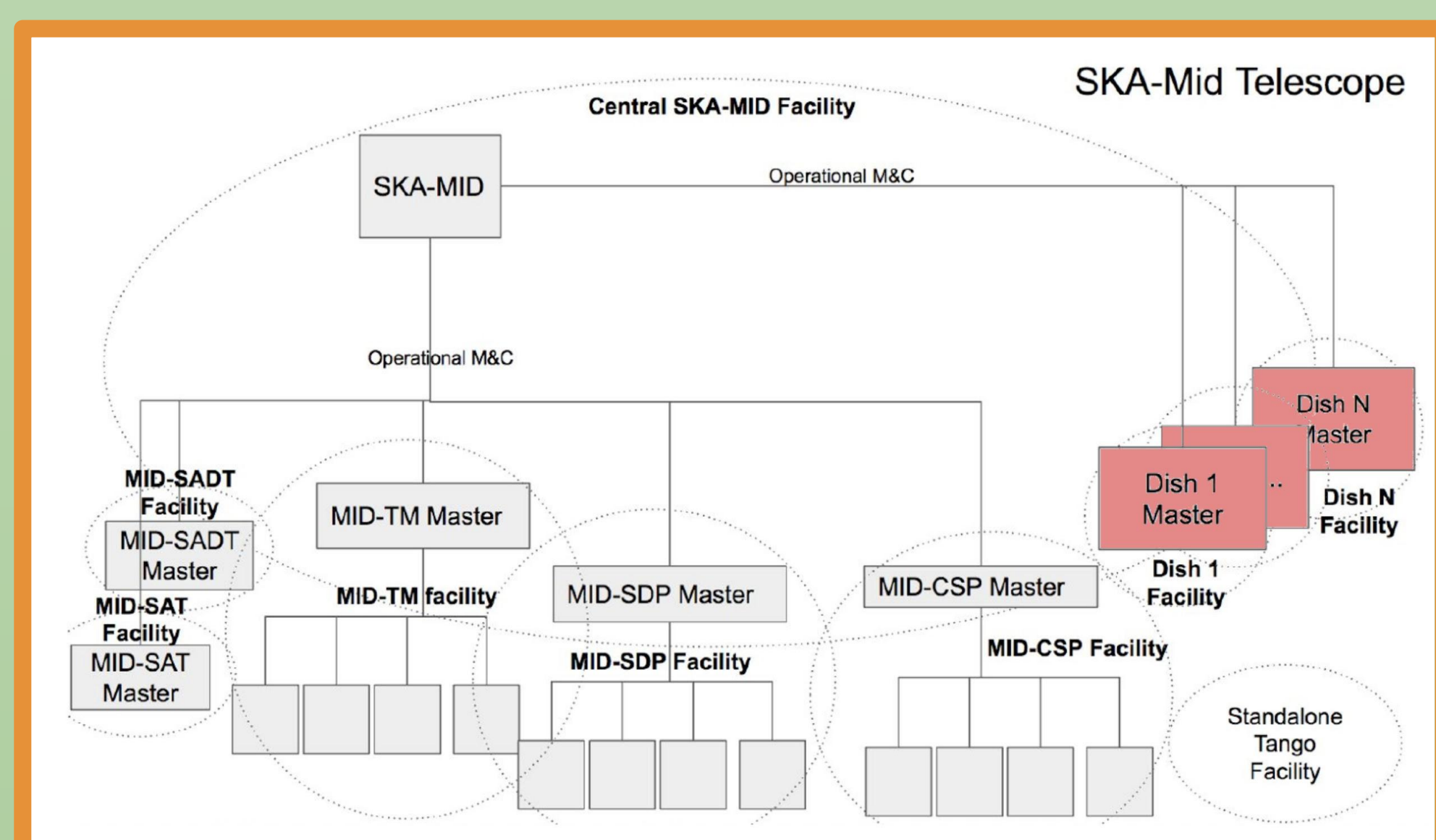


Fig. 4: High-level SKA1-Mid CS hierarchy

DISH LMC ROLE IN SKA CS

- ❖ **SKA Control System (CS) design** (Fig. 4)
 - based on Tango framework
 - organized in hierarchical Element facilities (=Tango domains)
 - Harmonized across Elements with common guidelines & patterns
- ❖ **Each Dish has a Local Monitoring and Control (LMC) system**
 - LMC provides dish master control and rolled-up monitoring to Telescope Manager (TM)
 - TM remotely coordinates the telescope scientific and M&C operations
- ❖ **Dish LMC "size" from a M&C view**
 - Monitoring points: <2000
 - M&C data flow to TM: ~200 kbps
 - Fastest M&C rate: 100 ms
 - Number of Tango Device Servers (TDS): <20

DISH LMC DESIGN AND PROTOTYPE

- ❖ **Dish LMC high-level architecture designed** (Fig. 5)
 - 1 master TDS for high-level Dish M&C
 - 1 TDS controller per Dish sub-element
 - 1 logger/archiver/alarm handler TDS per Dish instance
 - Additional TDS for LMC self M&C, power control
- ❖ **Prototype implementation started**
 - Aim: design validation + dish qualification
 - Technologies selected (Table 1)
 - Tango extensions implemented

LMC Tango extensions

- Logging to syslog targets
- Dynamic attributes from XML config files
- Tango attribute-based command state machines
- Formula attributes
- Tango device proxy utilities
- Task sequences
- Tango-Nagios interface

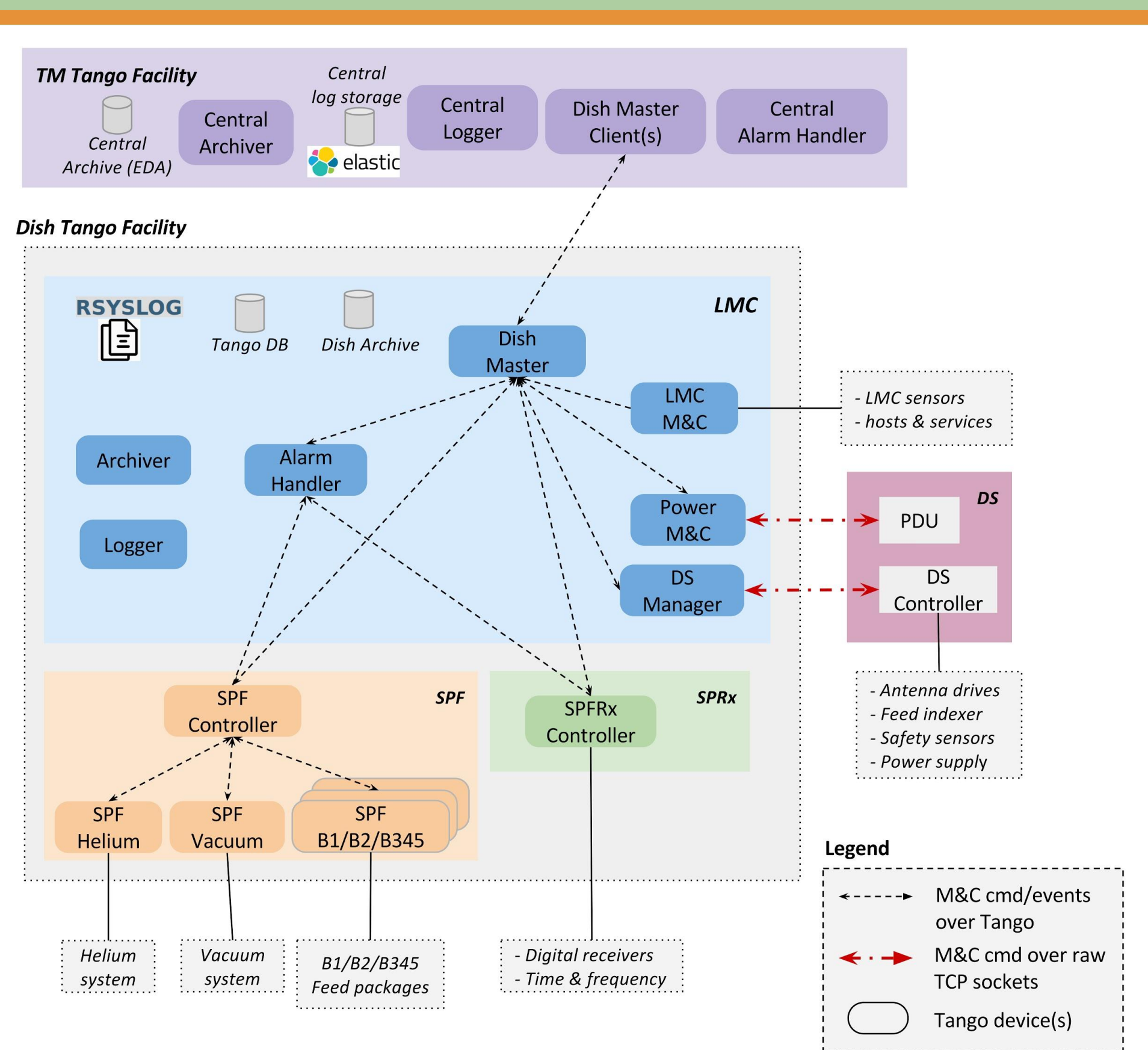


Fig. 5: Dish LMC high-level architecture

Table 1: List of Dish LMC prototype employed technologies

Type	Technology
Progr. language(s)	C++ (control system), python (configuration & testing)
M&C framework	Tango 9
Tango tools	PyTango, yat4tango, Elettra Alarm System, HDB++ Archiver, Taurus GUI
Libraries & tools	boost, jsoncpp, pugixml, Exptk, Nagios 4
Build system	cmake
Version control	git (GitLab)
Testing	Google Test, python behave + nose
Configuration management	Ansible
Continuous integration	Jenkins
Documentation	Doxygen + Sphinx/breathe
Virtualization & OS	VirtualBox, Ubuntu 14/16