

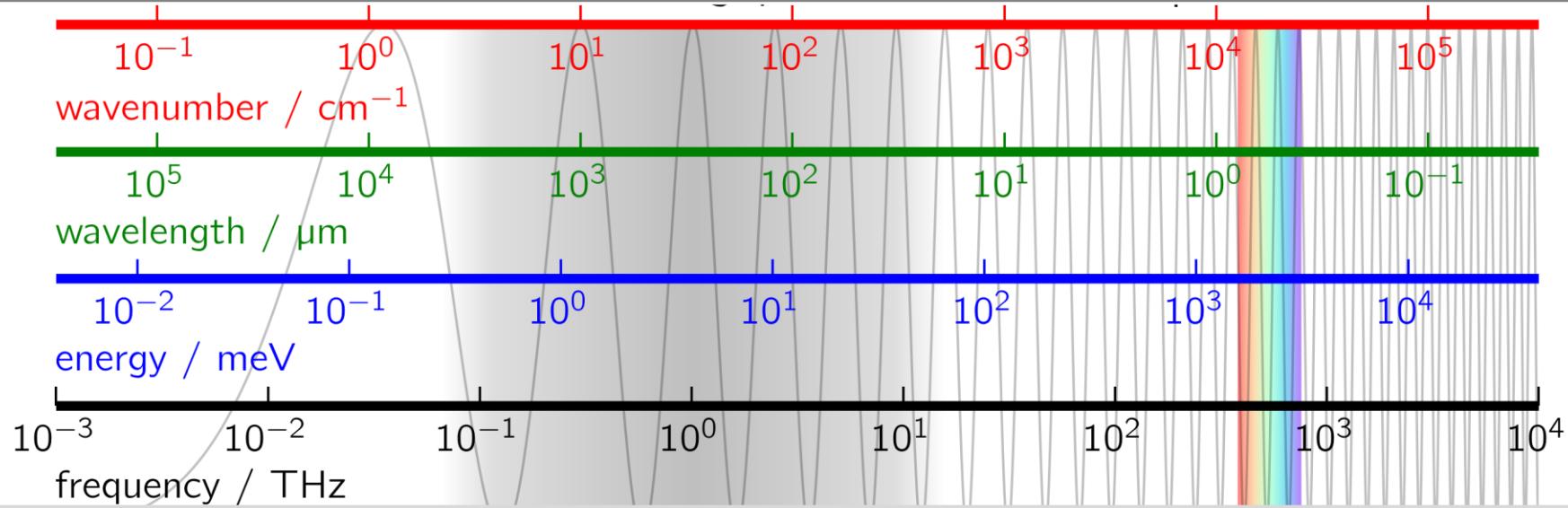
# THz Detection Techniques Overview

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**IBIC2020**

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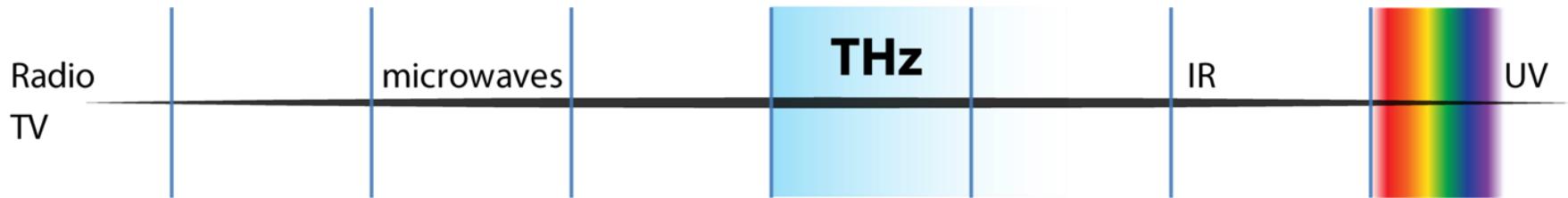


# Overview

- Introduction
- Thermal Detectors
- Direct Detection Devices
- Hetherodyne Detection
- Sampling Detection
- Summary

# Introduction – THz spectrum

f: 300 MHz    3 GHz    30 GHz    300 GHz    3 THz    30 THz    300 THz



λ: 1 m    10 cm    1 cm    1 mm    100 μm    10 μm    1 μm

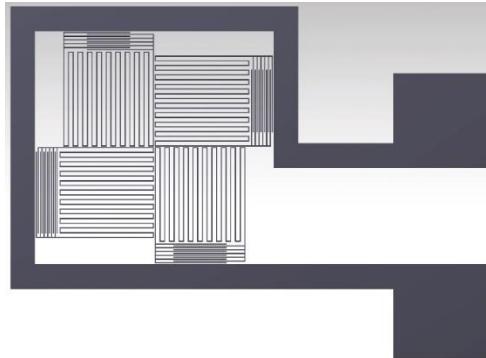
E: 1.24 μeV    12.4 μeV    124 μeV    1.24 meV    12.4 meV    124 meV    1.24 eV

- Many natural molecular dynamics happen in the THz range → molecular sensing, absorbed by atmospheric gases and water
- Non ionizing radiation → safety ambient thermal noise can be an issue ( $k_B T$  (300K) ~ THz!)
- Pass through many dielectrics, reflected by metals → imaging

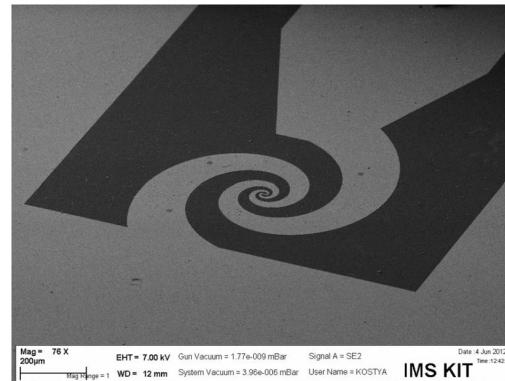
# THz Diagnostics - Overview of Sys. Req.



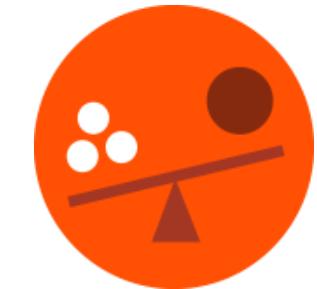
Dynamic range  
(e.g. Heterodyne  $\sim 100$  dB)



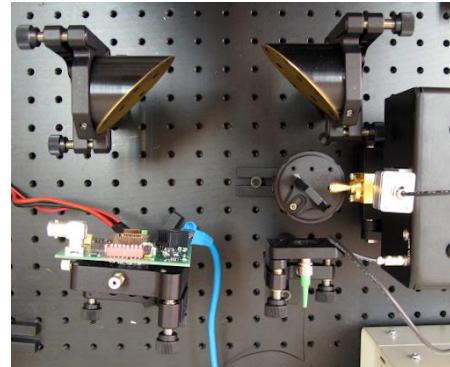
High Sensitivity  
(e.g. KID,  $\sim 10^{-19}$  W)



Speed  
(e.g. IMS YBCO, < 10 ps)

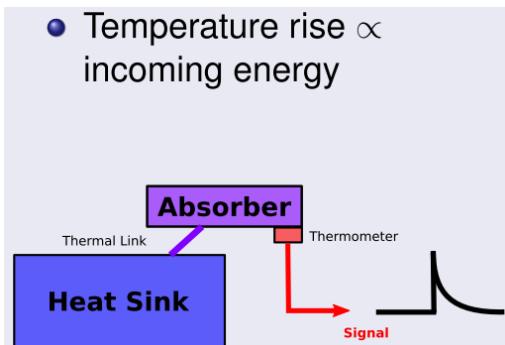


TRADE-OFFS  
No perfect device



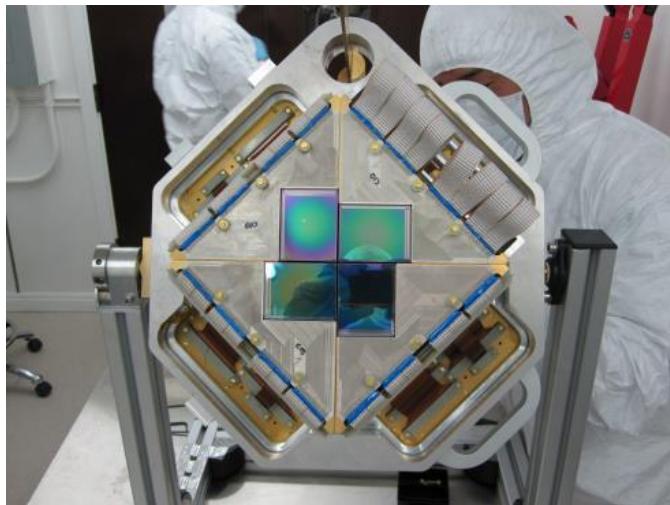
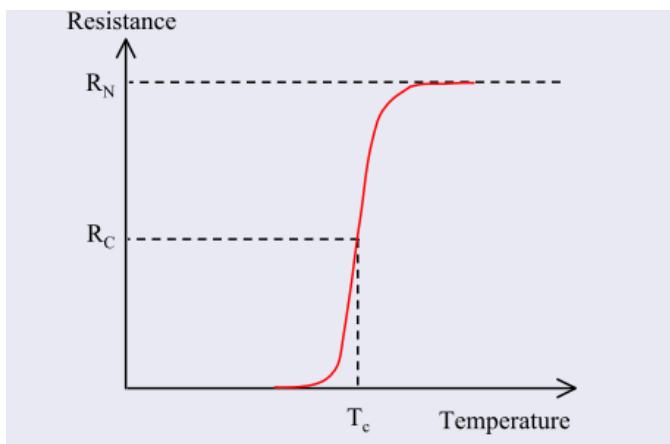
Bandwidth  
(e.g. EO sampling, 100 GHz  $\rightarrow$  37 THz)

# Thermal Detector – Bolometer



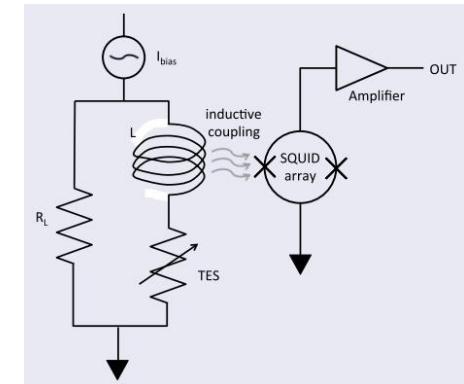
- Absorber + thermal reservoir + thermal connection
- Absorbed power  $P = G * \Delta T$ ,  $\Delta T$  given by resistive thermometer
- Faster than pyros and Golay cells ( $\sim 10 \mu s$ )
- Lowest noise room temperature detector (NEP:  $\sim 10^{-12} W/Hz^{1/2}$ )
- Widely used in large arrays in thermal cameras and in astrophysics (MAMBO2, SHARC II, SCUBA, P-Artemis...)

# Thermal Detector – Transition Edge Sensor

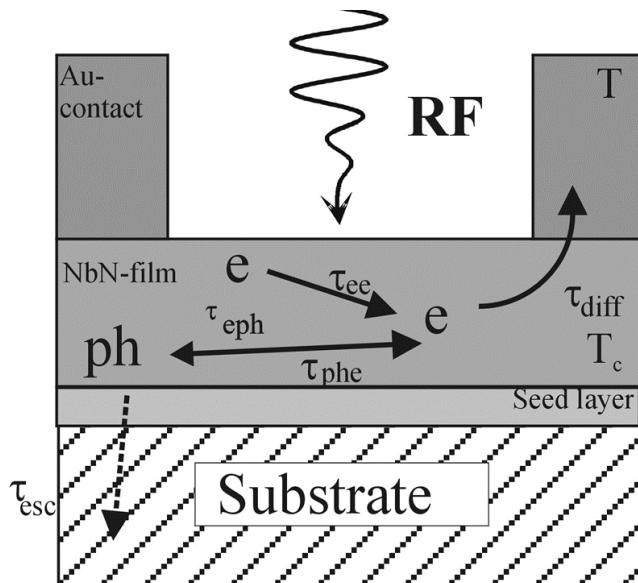


SCUBA2

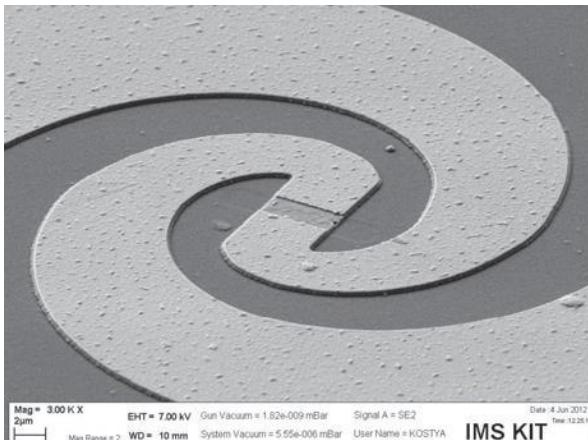
- Cryogenic version of bolometers
- Absorber is held at the superconducting transition temperature → strong T-R response → high det. eff. (~ 98 %)
- SQUID multiplexing + voltage bias →  $dR/dT < 0$  → negative electrothermal feedback → stability (low noise)
- Extremely sensitive: NEP:  $\sim 10^{-19} \text{ W/Hz}^{1/2}$  @  $10^{-15} \text{ W}$  (background load dependent)
- Slow response  $\sim \mu\text{s}$
- Expensive (50 mK!)



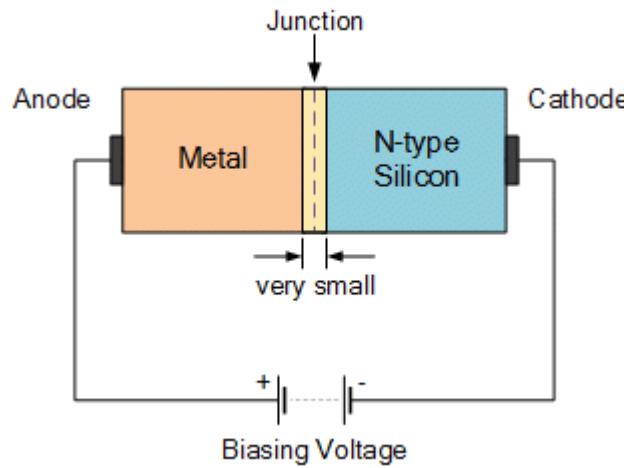
# Thermal Detector – Hot Electron bolometer



- Weak  $e^-$  - ph coupling at thermal equilibrium
- Radiation breaks the coupling → hot electrons
- Heat capacity = electrons
- Thermal conductance =  $e^-$  - ph relax. time
- Resp. time:  $\sim 10$  ps
- NEP:  $10^{-18} \text{ W/Hz}^{1/2}$



# Direct Detection – Schottky Diode



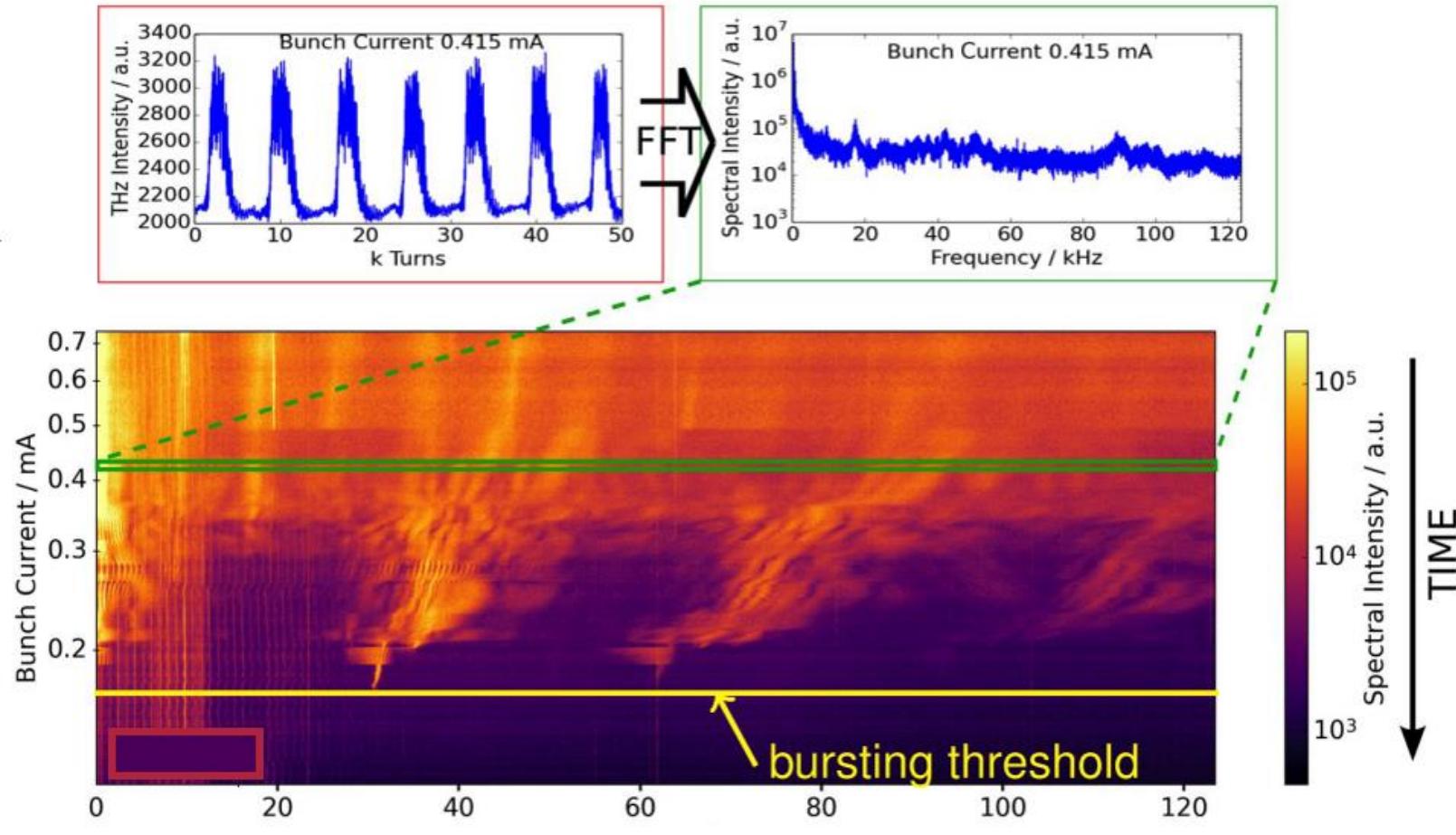
- Schottky Barrier Diodes are formed by the junction of metal with a semiconductor
- Very low forward voltage drop (but very high reverse leakage current)
- Majority carrier semiconductor device → very fast switching action (< 10 ps)
- Room temperature operation
- Can be operated bias free → low noise ( $10^{-12}$  W/Hz $^{1/2}$ )
- Ultra-Wideband: can be as wide as 50 GHz → 5 THz



<https://www.electronics-tutorials.ws/diode/schottky-diode.html>

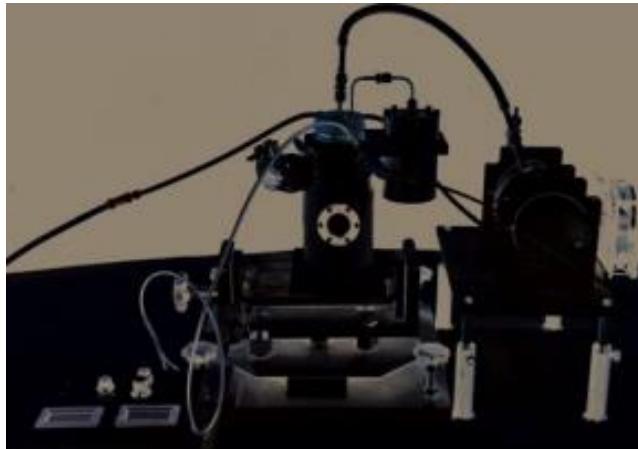
# Direct Detection – Schottky Diode

- Application: microbunching measurement @ KIT – IBPT
- Required time resolution: 2 ns

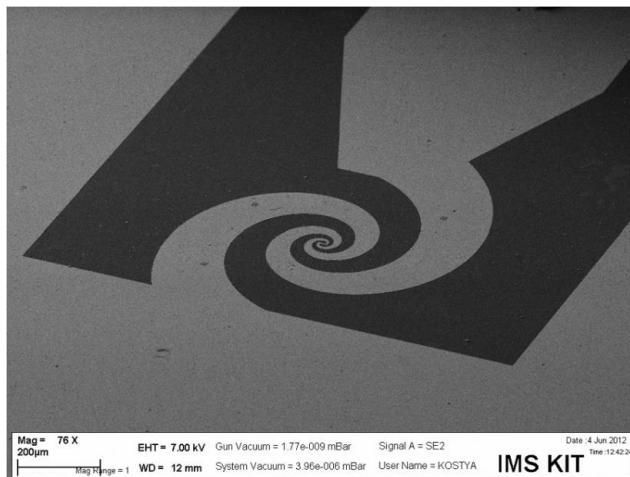


M. Brosi et al., PhysRevAccelBeams.19.110701

# Direct Detection – YBaCuO thin film detector



- 30 nm YBaCuO thin films detector
- Detection mechanism: non-bolometric, vortex assisted @ THz (Phys.RevB 85, 174511 (2012))
- High temp. superconductor → LN<sub>2</sub> cryo
- Ultrawide band 30 GHz → 2.5 THz
- Response time < 15 ps → CSR pulse real time evo

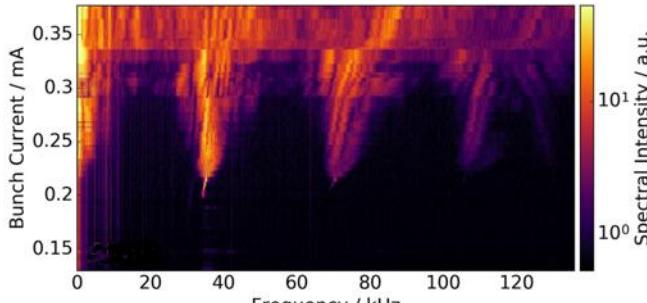
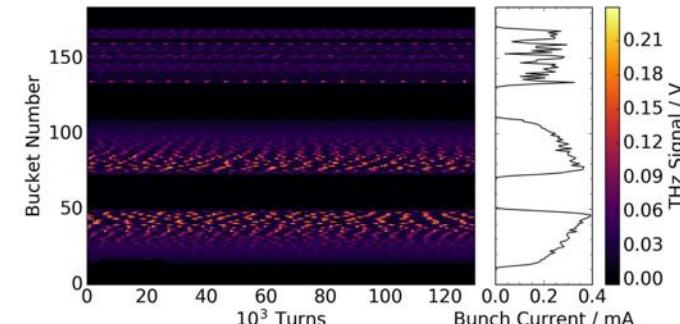
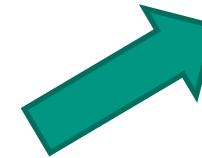


Thoma, P.; Raasch, J.; et al.; IEEE Trans. Appl. Supercond., vol.23, no.3, June 2013

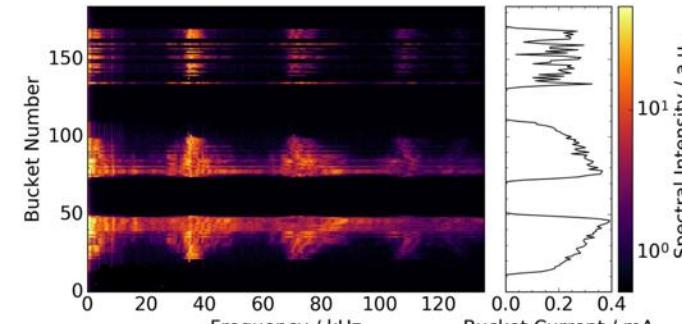
# Direct Detection – Readout system example

## KAPTURE: KArlsruhe Pulse Taking Ultra-Fast Readout Electronics

- YBCO, HEB, Schottky compatible
- Continuous acquisition
- Real time eval. Via GPUs



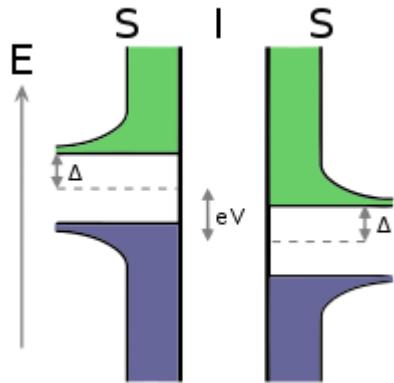
Sorting (bunch current)



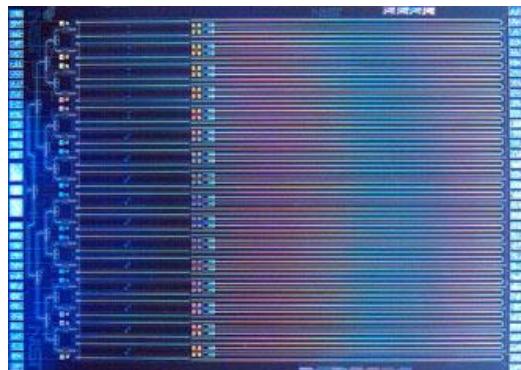
Bunch signal FFT

M. Caselle, et al. An ultra-fast data acquisition system for coherent synchrotron radiation with terahertz detectors, JInst 9 C01024

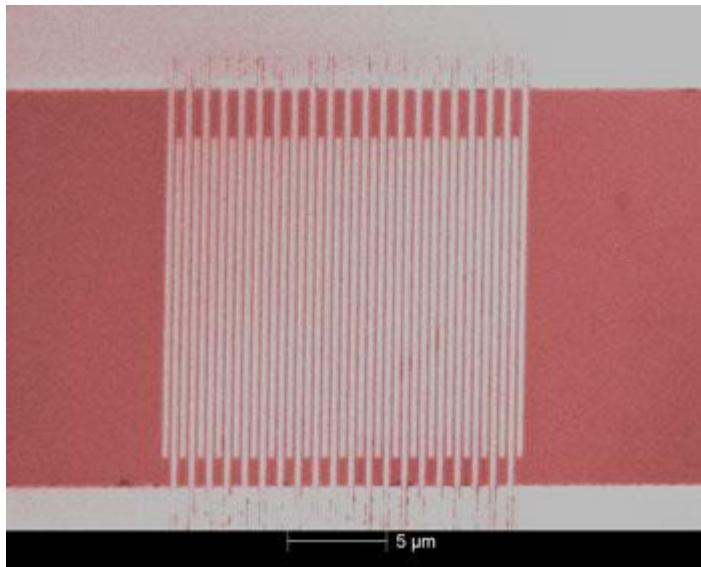
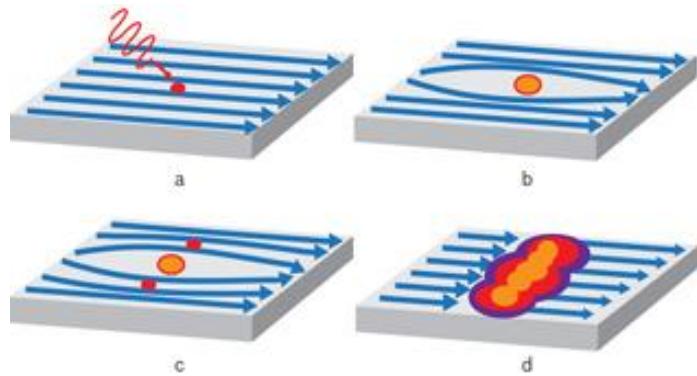
# Direct Detection – Superconducting Tunnel Junction



- Based on photon-assisted quantum tunneling of charge carriers through the insulating barrier
- Charge carriers: supercarriers (Cooper) and quasiparticles (electrons)
- The junction is voltage biased
- Incoming photons break down Cooper pairs, generating excess quasi particles → detectable tunnel current
- NEP:  $10^{-18} \text{ W/Hz}^{1/2}$  @ 1 THz Ariyoshi et al. Appl. Phys. Lett. 95, 193504
- Broadband (0.7 – 2.4 THz)



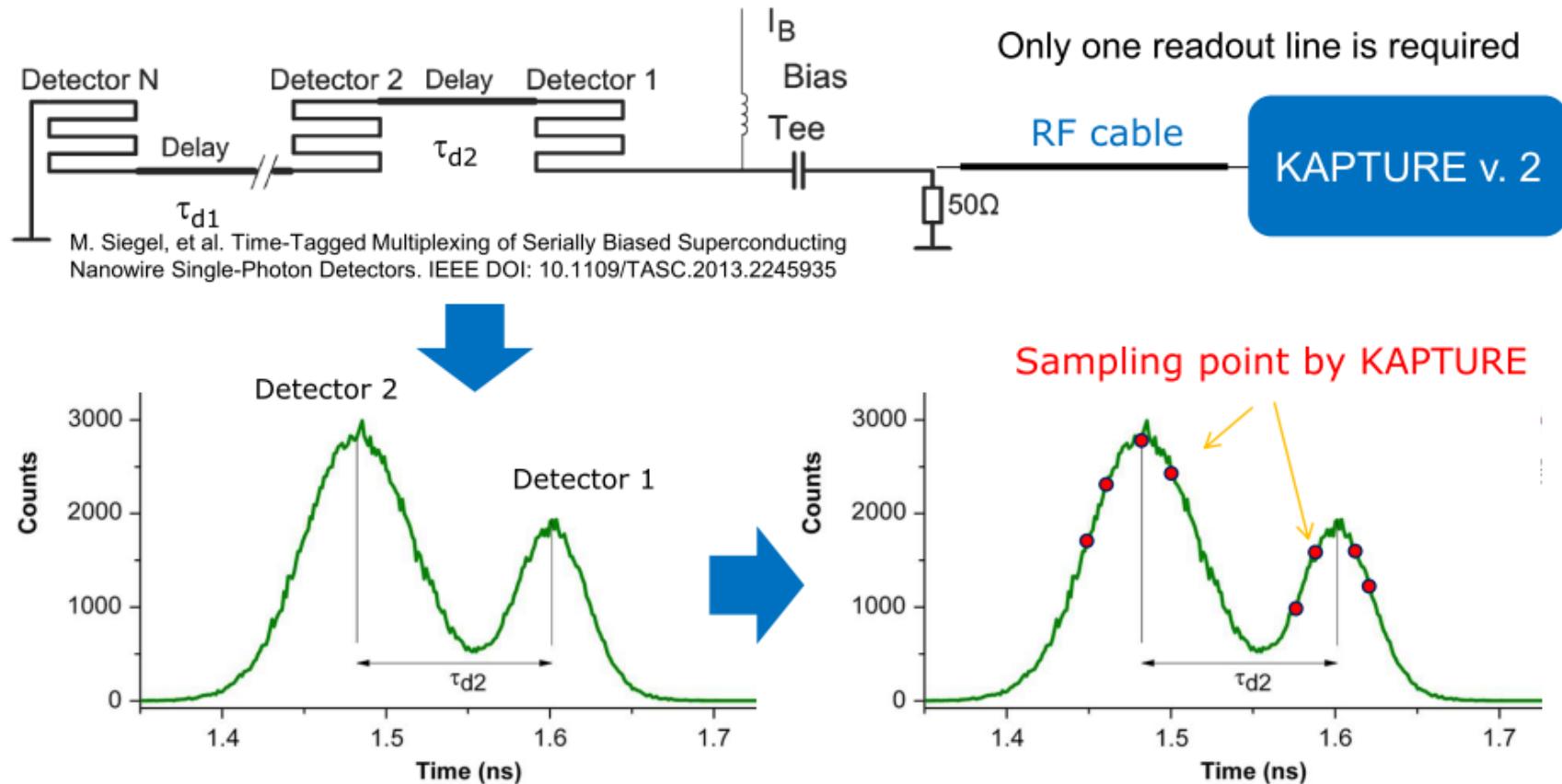
# Direct Detection – Superconducting Nanowire Single Photon Detector



- Biasing current close to superconducting critical current
- Photons incident on the nanowire generate quasi particle
- Quasi particles lower critical current locally → local resistance hotspots
- Typical nanowire dimensions: 5nm thickness, 100 nm width
- Very high detection efficiency ( $> 90\%$ )
- Extremely sensible (NEP:  $\sim 5 \cdot 10^{-21} \text{ W/Hz}^{1/2}$ )
- Resp. time  $< 20 \text{ ps}$

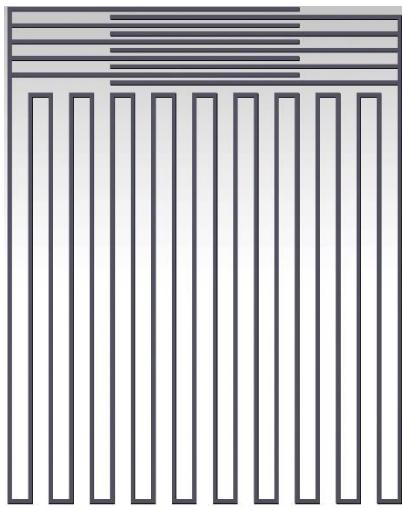
# Direct Detection – Superconducting Nanowire Single Photon Detector

Multipixel THz SNSP Detectors – Ongoing investigation @ KIT IMS & IBPT!

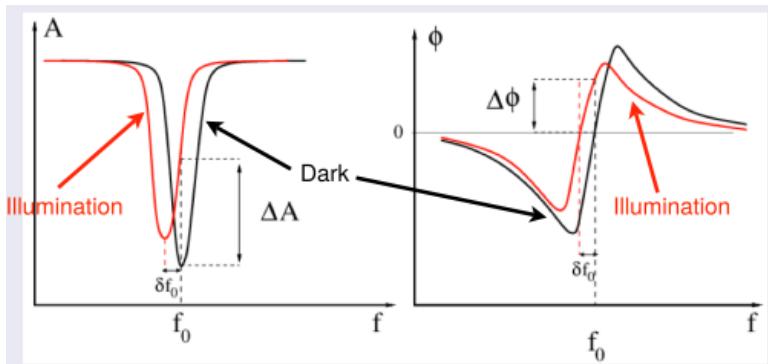


Designed to sample two ultra fast pulses with very short time distance. Time distance settable by FPGA from 25 ps to 400 ps with incremental step of 25 ps.

# Direct Detection – Kinetic Inductance Detector



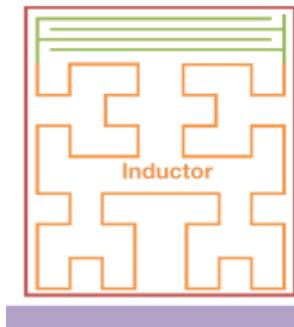
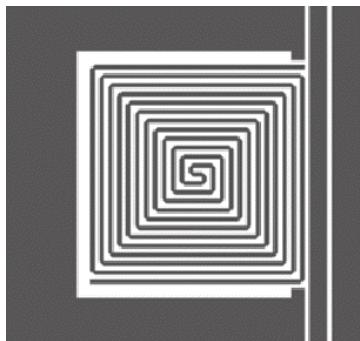
- Meandered Line: L
- Interdigital Capacitor: C



$$\sigma = \frac{ne^2\tau}{m(1+i\omega\tau)} = \frac{ne^2\tau}{m(1+\omega^2\tau^2)} - i \frac{ne^2\omega\tau^2}{m(1+\omega^2\tau^2)}$$

- Appreciable ONLY in superconductors
- $L_K \sim 1/n_C$
- Photons → Cooper pairs breakage → excess quasiparticles →  $L_K$  increases upon photon absorption
- Incoming photons are detected through a change in the resonant frequency and phase of the circuit ( $S_{21}$  parameter)

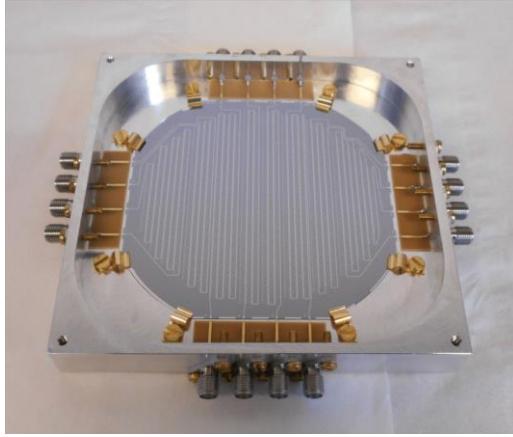
# Direct Detection – Kinetic Inductance Detector



- Many possible configurations (ex: grid, spiral, Hilbert inductor) → polarization sensitiveness, lumped elements (no antenna)
- Ease of fabrication: single layer deposition, planar geometry, MS/CPW, photolithography
- Very high number of pixels ( $>1000$ ) feasible with simple read out
- Extreme sensitivity possible (record NEP  $3.8 * 10^{-19} \text{ W/Hz}^{1/2}$ )
- Energy res.  $\sim 100 \text{ meV}$  (theoretical GR limit  $\sim 10 \text{ meV}$ )
- Rise time  $\sim 50 \text{ ps}$  (intrinsic  $\sim 10 \text{ ps}$ )
- High degree of tunability: bandwidth from RF GHz to X rays

# Direct Detection – Kinetic Inductance Detector

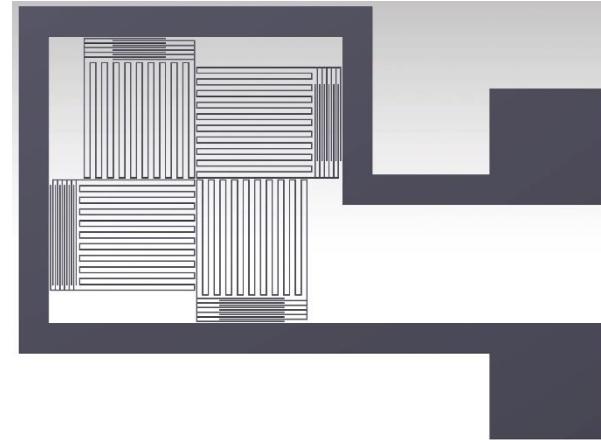
## Astrophysics Application – NIKA2



R.Adam et al., A&A 609, A115 (2018)

- Dual band camera (150 / 260 GHz) for IRAM's 30 m radiotelescope
- 2900 detectors over three monolithic arrays
- Al thin film over HRSi
- 150 mK operating temperature

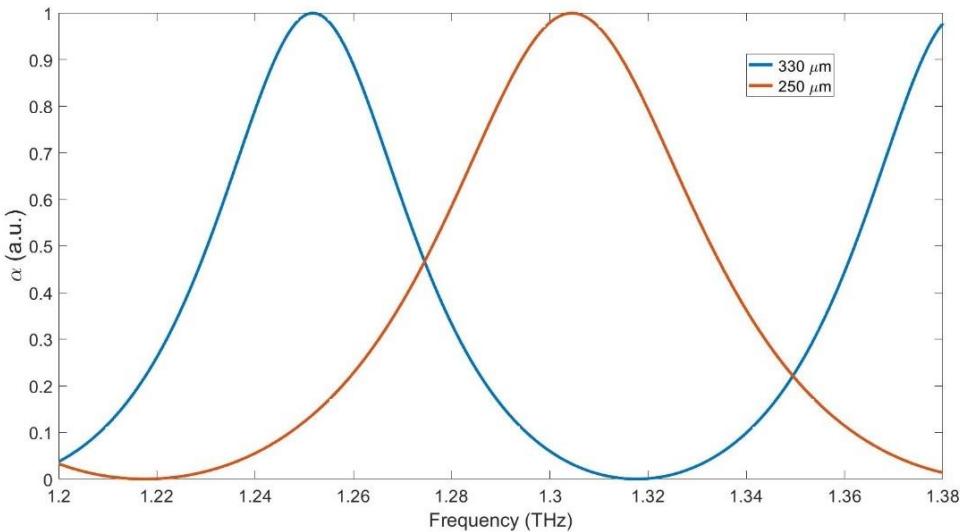
## Fusion Plasma Diagnostics – Polarimeter



F.Mazzocchi et al., Fus.EngDes 130, May 2018

- 4 pixel, polarization sensitive array
- NbN 15 nm thin film, Si/Sapphire/ Diamond substrates
- 1.3 THz detection frequency
- 4.2 K operating temperature

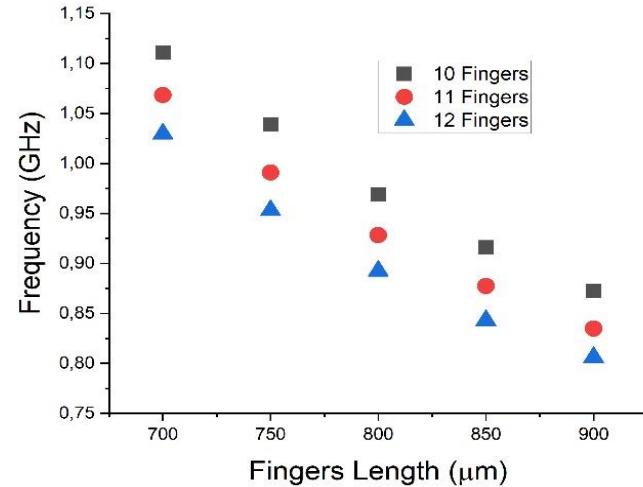
# KID – Characteristics and Tuning



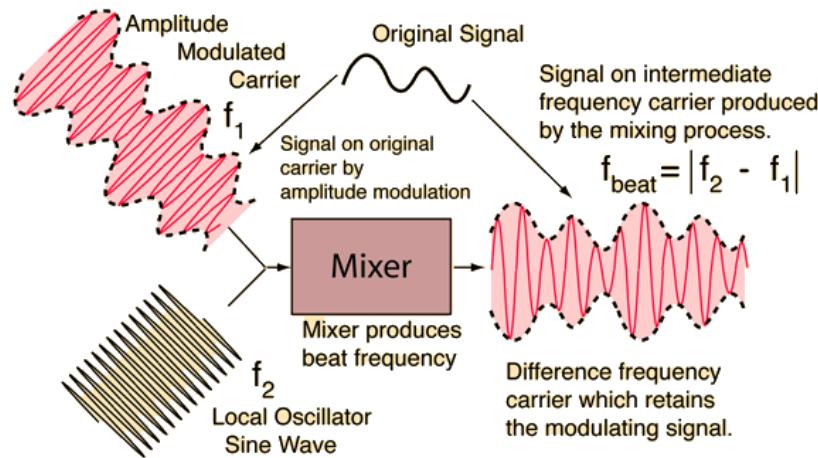
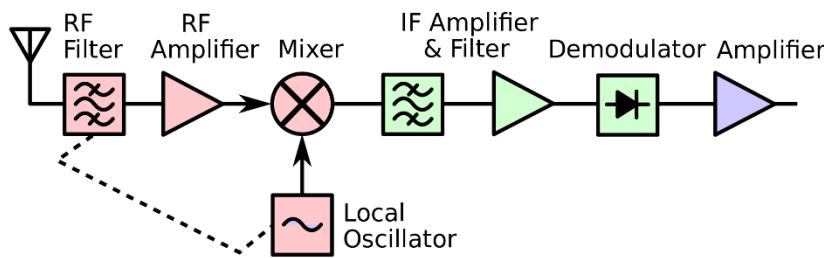
- Radiation coupling: substrate + thin film impedance
- Backshort  $\rightarrow$  substrate thick.  $\rightarrow$  spectral response
- Interdigital capacitor finger length  $\rightarrow$  resonators tuning

$$\alpha = 1 - \left| \frac{Z_0 - Z_{Eff}}{Z_0 + Z_{Eff}} \right|^2 \quad Z_{Eff} = \frac{1}{\frac{1}{Z_{KID}} + \frac{1}{Z_{Sub}}}$$

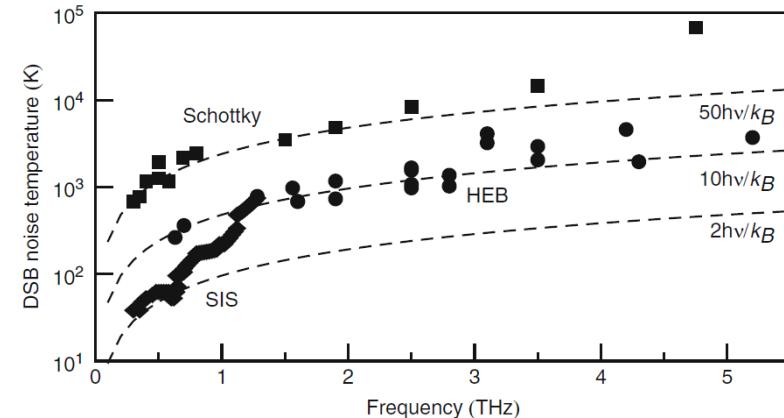
$$Z_{KID} = \frac{\rho_{NbN} S}{d_{NbN} w} \quad Z_{Sub} = j Z_{Sub} \tan(\beta l),$$



# Heterodyne Detection

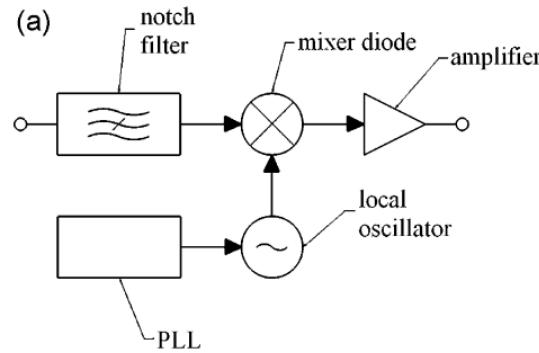


- Detection techniques based on mixing the incoming THz signal with LO signal
- Mixer: non linear component, producing the beat IF
- THz LO: Multipliers ( $\rightarrow 2$  THz), QCLs ( $> 2$  THz)
- THz Mixer: Schottky (RT), HEB, STJ (Cryo)

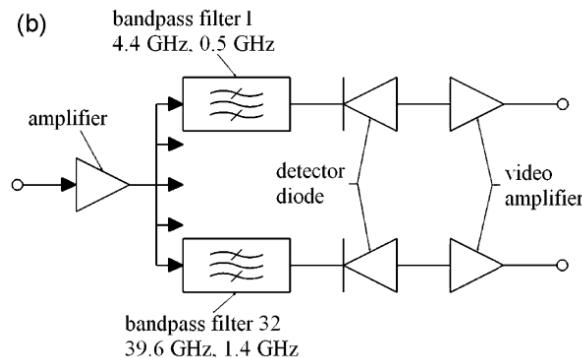


E. Bründermann, H.W. Hübers, M.F. Kimmit, Terhertz Techniques, Springer

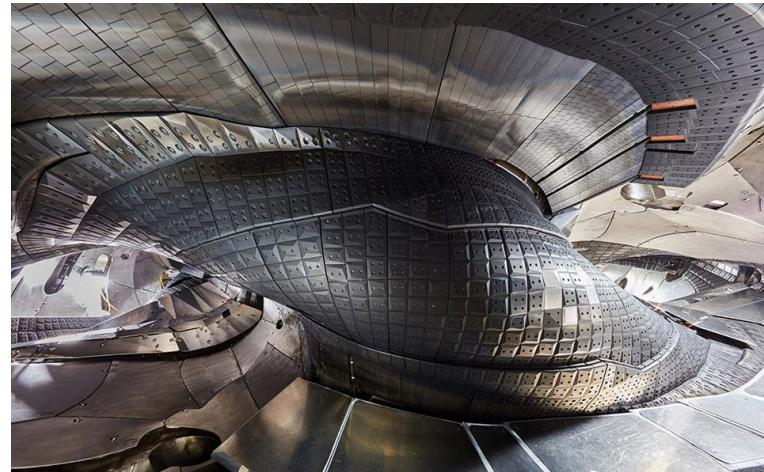
# Heterodyne Detection – W7X ECE 32ch radiometer



Front end

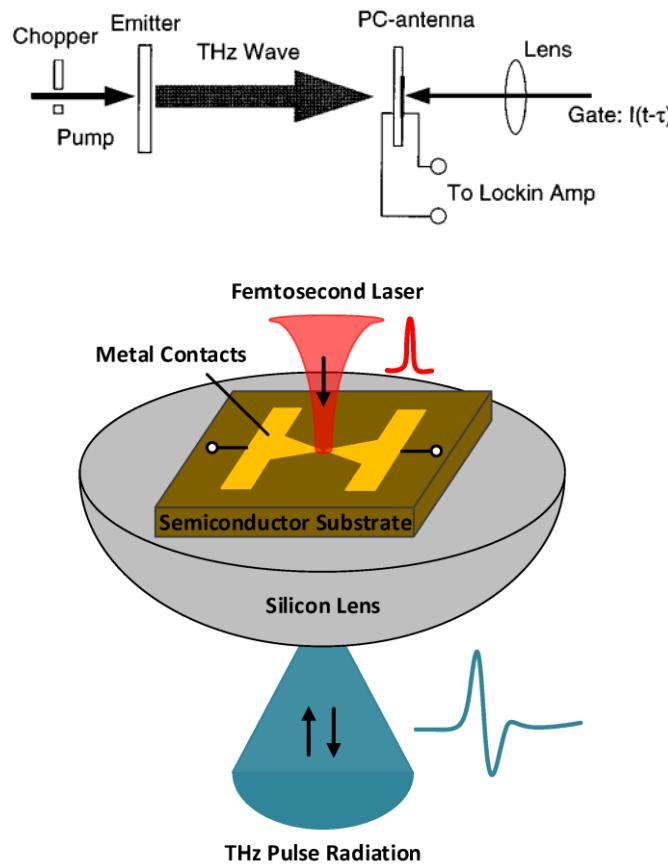


IF stage



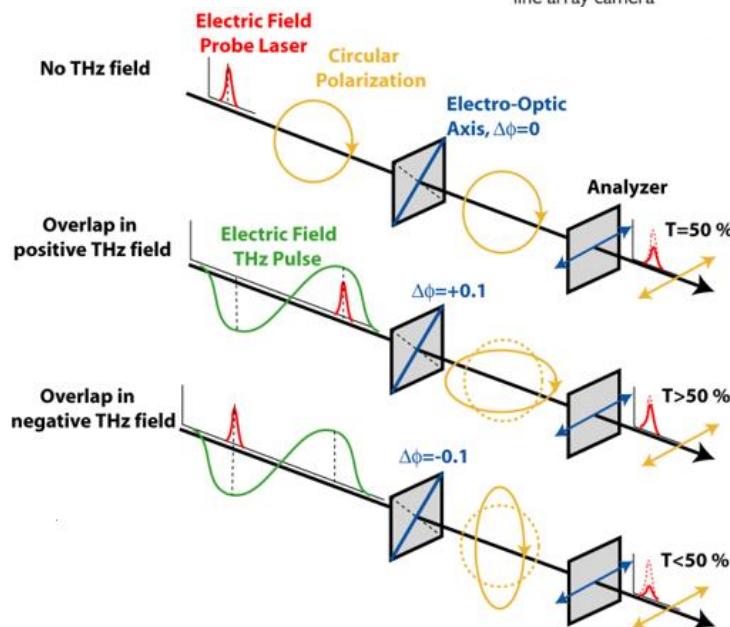
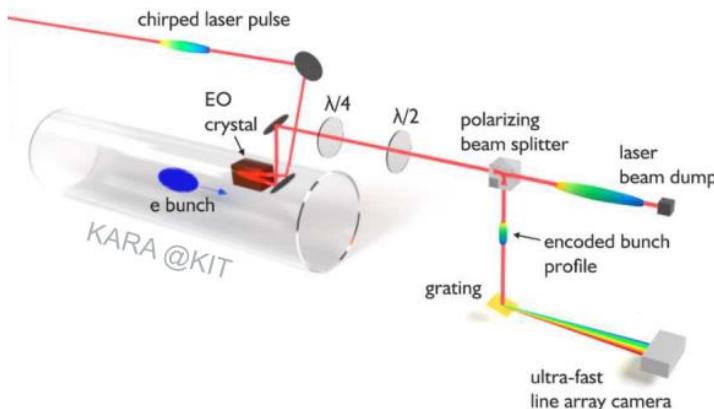
- Electron Cyclotron Emission diagnostic → plasma electron temperature profile
- Detects second harmonic of electron gyromotion fundamental mode at 70 GHz
- Notch filters at 140 GHz (ECRH) to avoid detector overload
- 126 – 160 GHz range, downconverted to 4 – 40 GHz

# Sampling detection – Auston switch



- Auston switch: gated photoconductive antenna (semiconductor bridged gap)
- Femtosecond pulse laser increases the conductivity of the antenna, excites charge carriers
- Incoming THz radiation induces a measurable photocurrent
- Response time depends on antenna structure and photocarriers lifetime ( $\sim 300$  fs for InGaAs emitter)
- Antenna parameters also limit the bandwidth to typically 6 THz

# Sampling Detection – Electro Optical

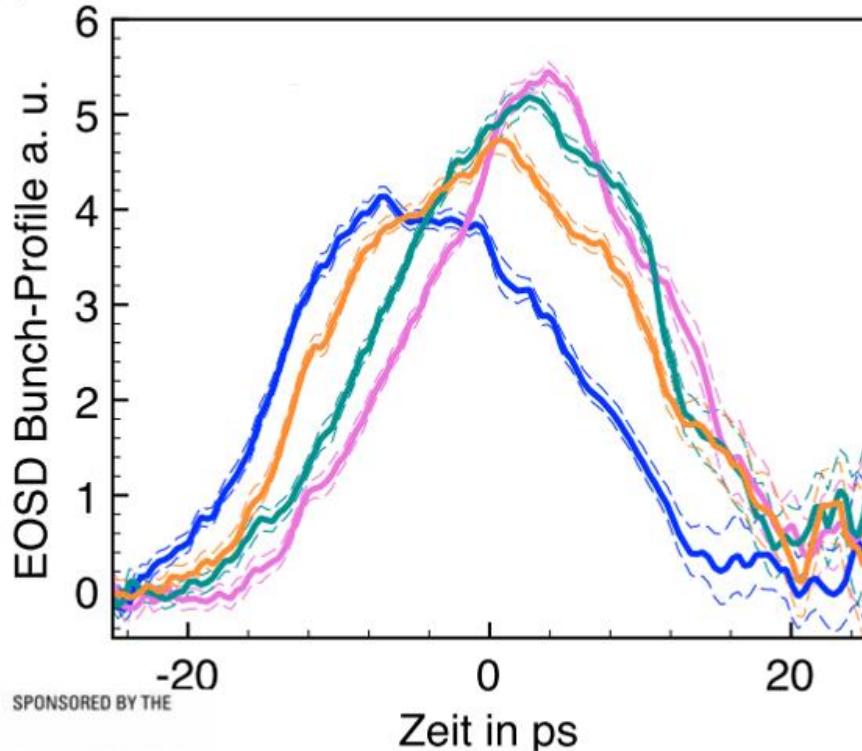


- Pockels effect: EO crystal (for example, ZnTe) birefringency  $\mathbf{Q}$  bias voltage
- The THz signal works as modulation signal for the crystal birefringency
- Polarization status of an ultrafast (fs) probe beam is modulated by the THz radiation
- WP separates P and S components of the encoded bunch profile
- Balanced detector P-S Q THz ampl.
- Huge bandwidth: 100 GHz – 37 THz

S. Funkner, G. Niehues, M. J. Nasse, E. Bründermann, M. Caselle, B. Kehrer, L. Rota, P. Schönenfeldt, M. Schuh, B. Steffen, J. L. Steinmann, M. Weber, A.-S. Müller arXiv preprint, arXiv:1912.01323

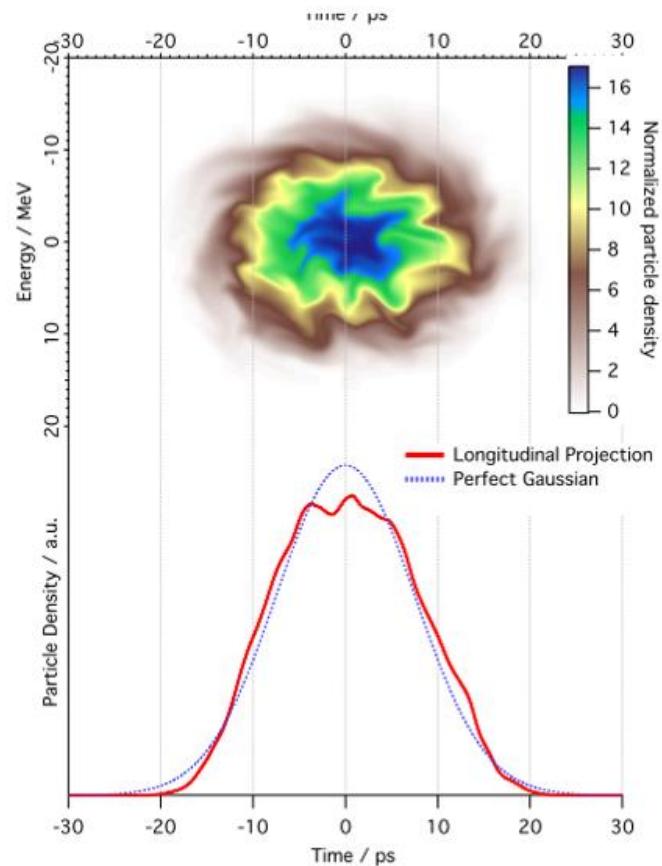
# Sampling Detection – EO @ KARA

- Single-shot EOSD measurements show dynamic sub-structures



N. Hiller, et al., IBIC 2014, MOPD17

simulated phase space



courtesy J. Steinmann, P. Schönfeldt

# Summary

- Large variety of detectors and techniques exist or are under development in THz detection:
  - Thermal
  - Direct
  - Heterodyne
  - Sampling
- Direct detection techniques generally offer very good performances (sensitivity, speed, etc.) with the advantage of being relatively simple
- Innovative devices like YBaCuO and KID are being developed as low-cost simple solutions with very good perspective



**THANK YOU!**  
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