



Deceleration of Highly Charged Heavy Ion Beams

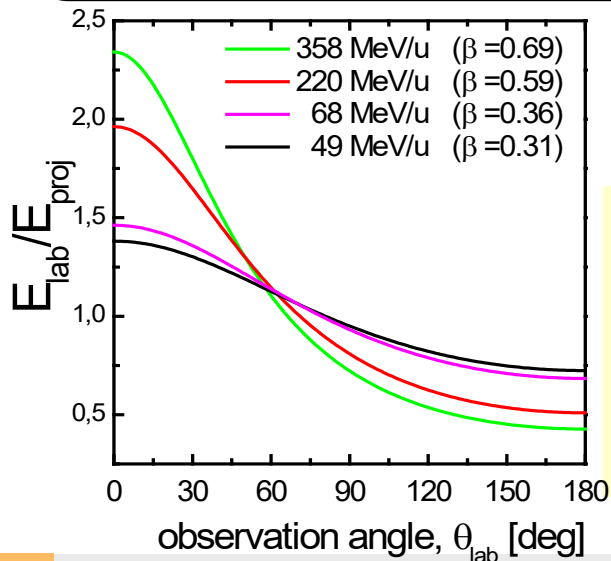
M. Steck, Accelerator Operations, Storage Rings,
GSI Helmholtzzentrum, Darmstadt

relativistic Doppler transformation

$$E_{\text{lab}} = \frac{E_{\text{proj}}}{\gamma \cdot (1 - \beta \cdot \cos \theta_{\text{lab}})}$$

relativistic transformation of solid angle:

$$\Delta\Omega_{\text{lab}} = \frac{\Delta\Omega_{\text{proj}}}{\gamma^2 \cdot (1 - \beta \cdot \cos \theta_{\text{lab}})^2}$$

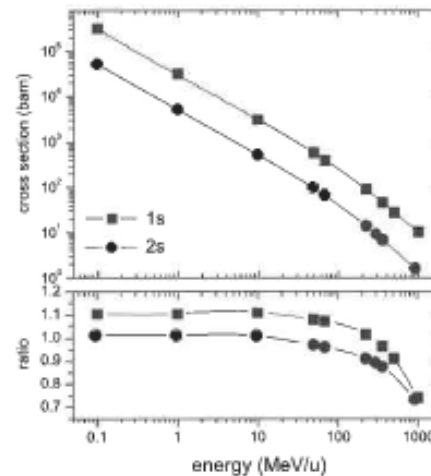


E_{lab} : photon energy in lab system

E_{proj} : photon energy in emitter system

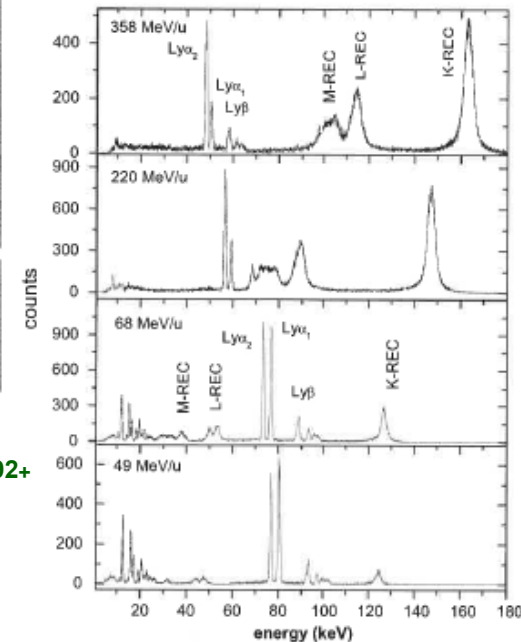
deceleration reduces:

- dependence on velocity and observation angle θ_{lab}
- influence of energy and intensity



radiative recombination U^{92+}

J. Eichler, T. Stöhler / Physics Reports 439 (2007) 1–99



X-rays of REC $U^{92+} \rightarrow N_2$

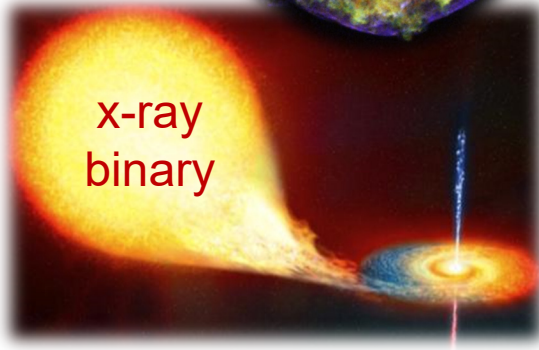
deceleration:

- provides increased reaction cross sections
- opens new reaction channels (transitions, polarization)

Decelerated Beams for Nuclear Astrophysics



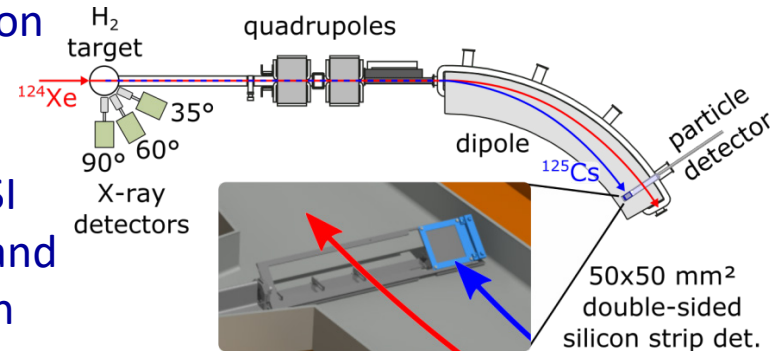
supernova
remnant



x-ray
binary

To investigate the origin of elements in stars, nuclear astrophysics aims for challenging reaction studies on rare ion beams.

Heavy ion storage rings at GSI provide unique possibilities and unrivalled conditions for such experiments.



Experimental setup of the reaction study $^{124}\text{Xe}(p,\gamma)$ in the ESR at beam energies as low as 5.5 MeV/u.

- high energy production & separation of radioactive beams in FRS
 - most efficient and versatile technique available
- deceleration to energies below 10 MeV/u
 - access to the famous Gamow window relevant for nuclear physics of stars
- cooled beam in combination with a thin gas jet target
 - inverse kinematics studies at unmatched energy resolution
- storage and recycling of the rare ion beam
 - extremely efficient technique for studies on beams of limited intensity

Jan Glorius

The GSI Deceleration Facility

11.4 MeV/u → 400 MeV/u → 400–X MeV/u

UNILAC



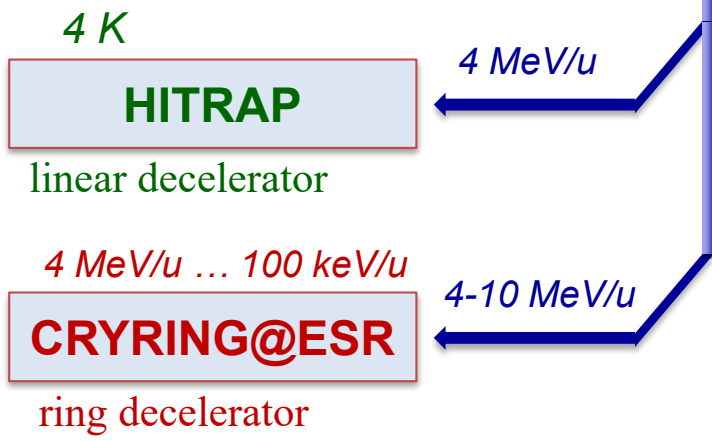
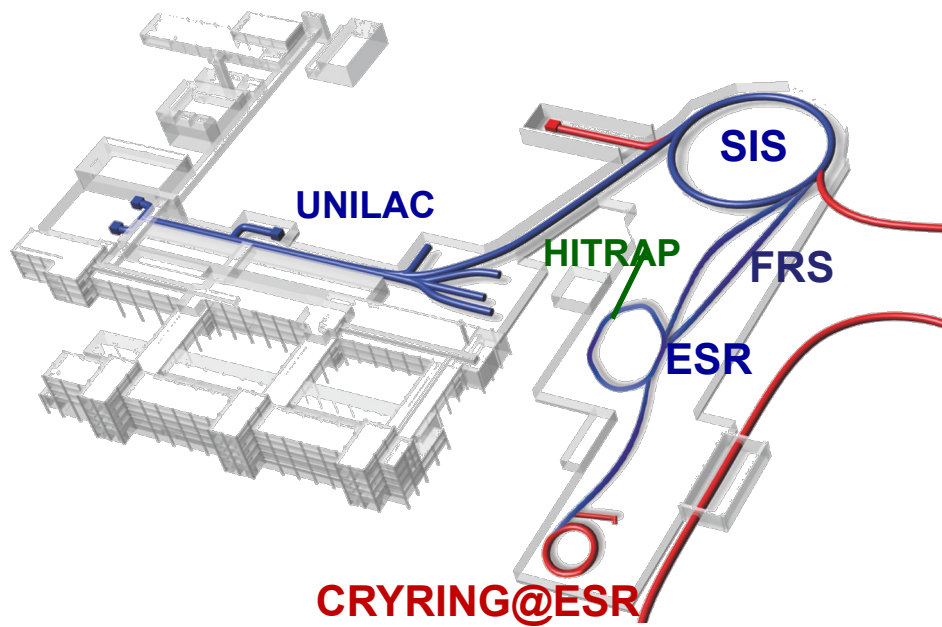
stripper/
production target

FRS



ESR

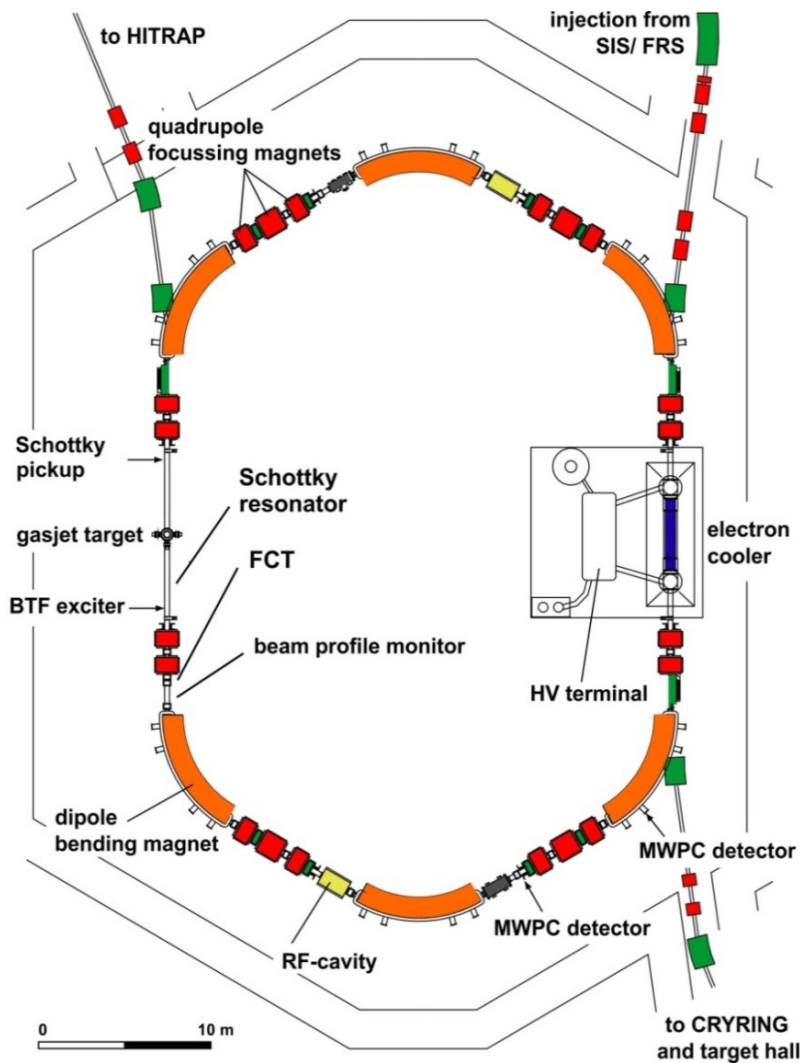
Slow, heavy, highly charged ions stored and well controlled
 Energy range 10 MeV/u to sub eV/ion
 10⁵ to 10⁷ highly charged ions. e.g. U⁹¹⁺



Requests for the Deceleration of HCI and RI Beams in the ESR

- good beam quality of stored beam \Rightarrow high precision experiments**
- high luminosity with internal target for internal experiment**
- long lifetime of stored beam \Rightarrow vacuum pressure 10^{-11} mbar or below**
- fast extraction at 4 MeV/u for HITRAP, defined by the linac injection energy**
- fast extraction below ~ 10 MeV/u for CRYRING, defined by bending power**
- slow extraction employing charge changing processes (cooler, target)**
- maximum intensity (particles per time interval) to HITRAP/CRYRING**
- i.e. large particle number, short cycle time**

The Heavy Ion Storage Ring ESR



Fast injection (stable HCIs / RIBs)

Stochastic cooling (≥ 400 MeV/u)

Electron cooling (3 - 430 MeV/u)

Laser cooling (C^{3+} 120 MeV/u)

Deceleration (down to 3 MeV/u)

Fast extraction (HITRAP/CRYRING@ESR)

Slow (resonant) extraction

Ultraslow extraction (charge change)

Beam accumulation

Internal gas jet target

Multi charge state operation

Schottky mass spectrometry of RIBs

Isochronous mode (TOF detector)

Beam Cooling at the ESR

stochastic pre-cooling on the injection orbit

energy 400 (-550) MeV/u
 bandwidth 0.8 GHz (range 0.9-1.7 GHz)
 $\delta p/p = \pm 0.35\%$ \rightarrow $\delta p/p = \pm 0.01\%$
 $\epsilon = 10\ \mu\text{m}$ \rightarrow $\epsilon = 2\ \mu\text{m}$



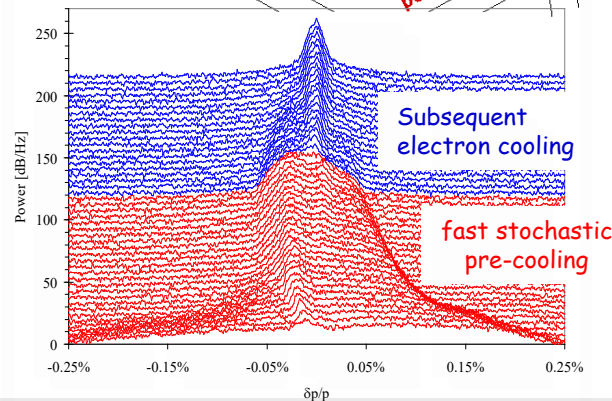
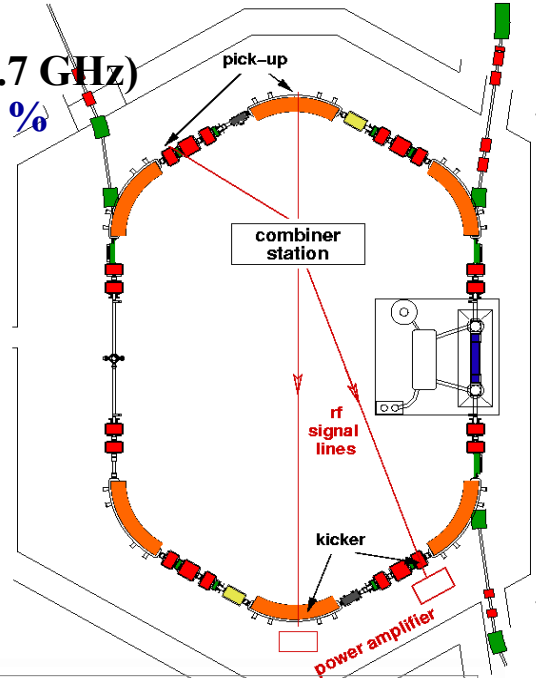
electrodes



combiner station



power amplifiers



electron cooling

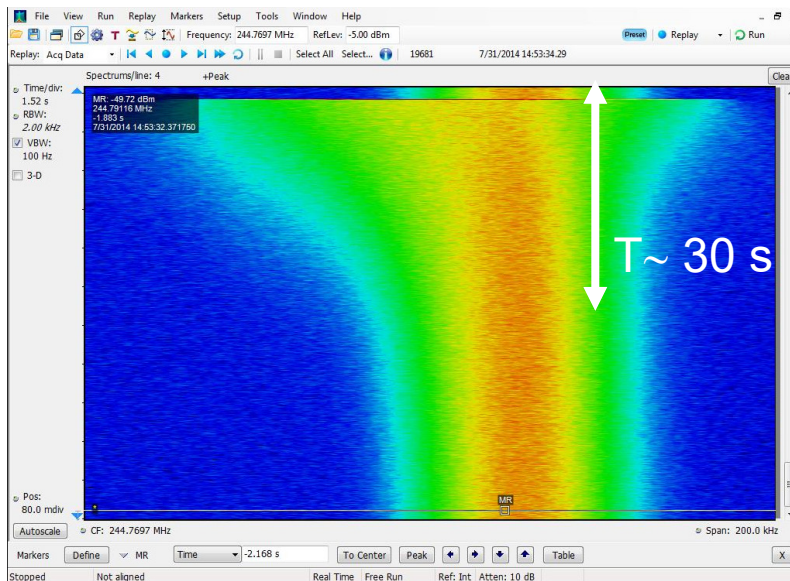


electron energy 1.6 – 250 keV
 electron current 0.001 – 1 A
 beam diameter 50.8 mm
 magnetic field 0.01 - 0.2 T
 collection efficiency 0.9998
 transverse temperature 0.1 eV
 longitudinal temperature $\sim 0.1\ \text{meV}$
 vacuum $2 \times 10^{-11}\ \text{mbar}$

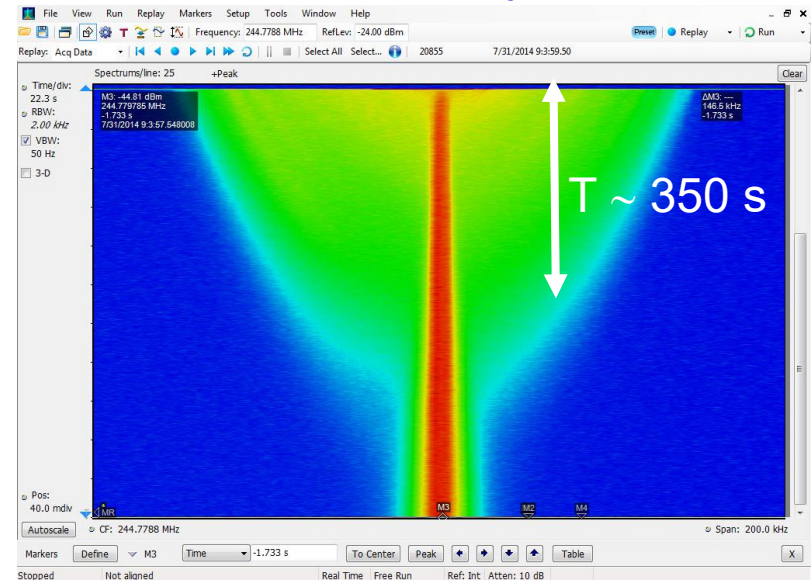
Cooling after Injection

- The injection energy of the ESR is flexible (30 - 400 MeV/u)
- the production rate of highly charged heavy ions increases with energy, but at higher energy cooling time increases \Rightarrow trade-off for maximum luminosity
- electron cooling can be applied for any energy ≤ 430 MeV/u
- stochastic cooling is presently available at 400 MeV/u but could be extended to higher energies (up to 550 MeV/u)
- at 400 MeV/u stochastic cooling is faster than electron cooling

protons 400 MeV stochastic cooling



electron cooling ($I_{el}=0.25$ A)

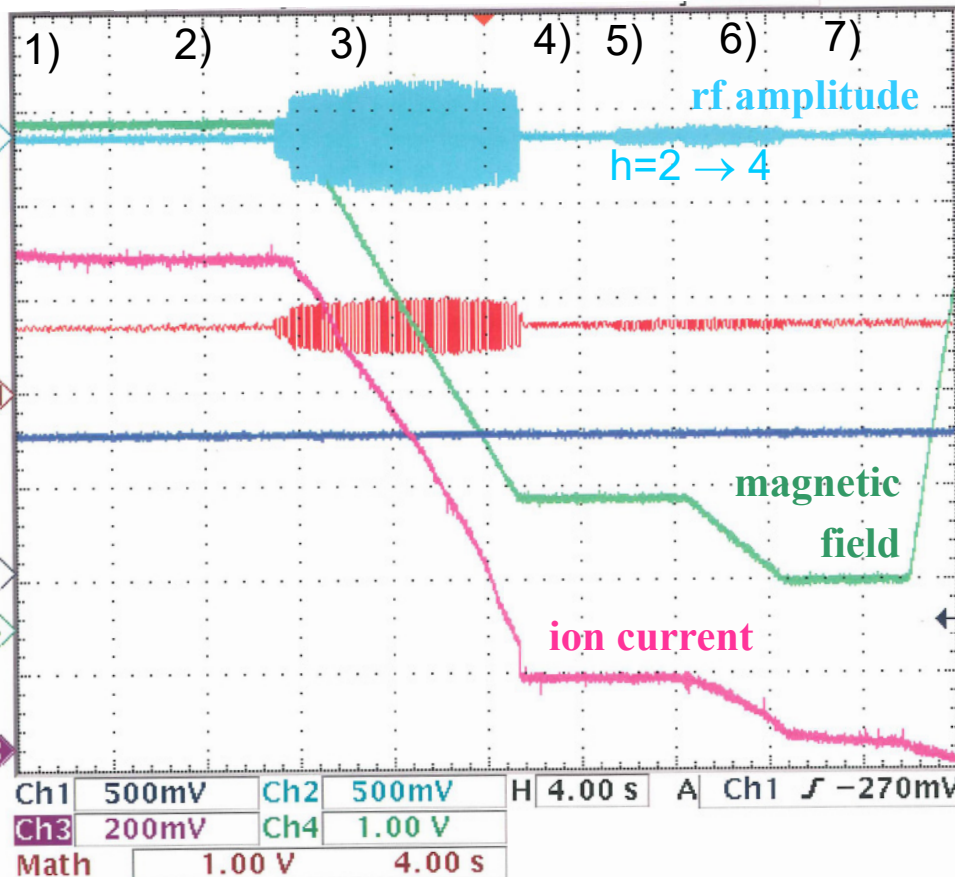


Typical ESR Deceleration Cycle

Ni^{28+} 400 \rightarrow 30 \rightarrow 4 MeV/u

1100 μA \rightarrow 180 μA \rightarrow 25 μA
 45% 37 %

cycle time 45 s

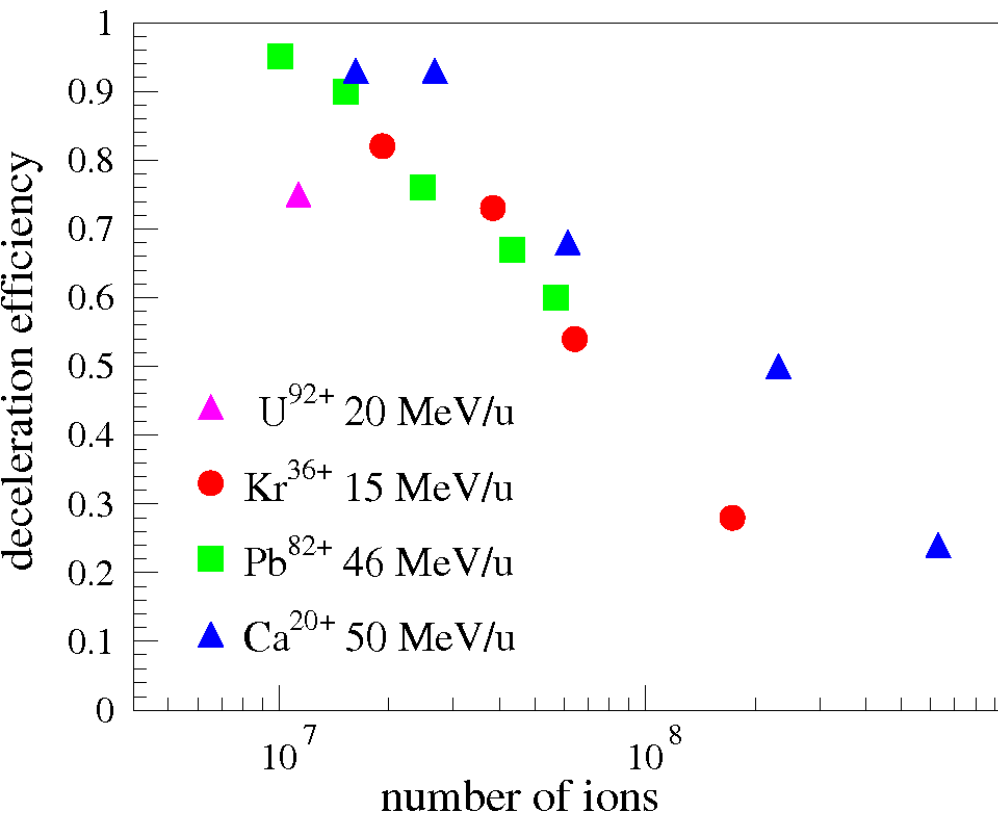


- 1) injection (at 400 MeV/u)
- 2) stochastic/electron cooling
(controlled ramp of cooler HV is difficult)
- 3) deceleration to 30 MeV/u
- 4) electron cooling
- 5) changing rf harmonic 2 \rightarrow 4
- 6) deceleration to final energy
- 7) electron cooling, bunching, extraction

maximum change of p ($B\rho$): 13

Efficiency of Deceleration (2000-2015)

Losses increase with increase of initial stored particle number most likely due to IBS



in a bunched beam during deceleration transverse emittance is growing due to

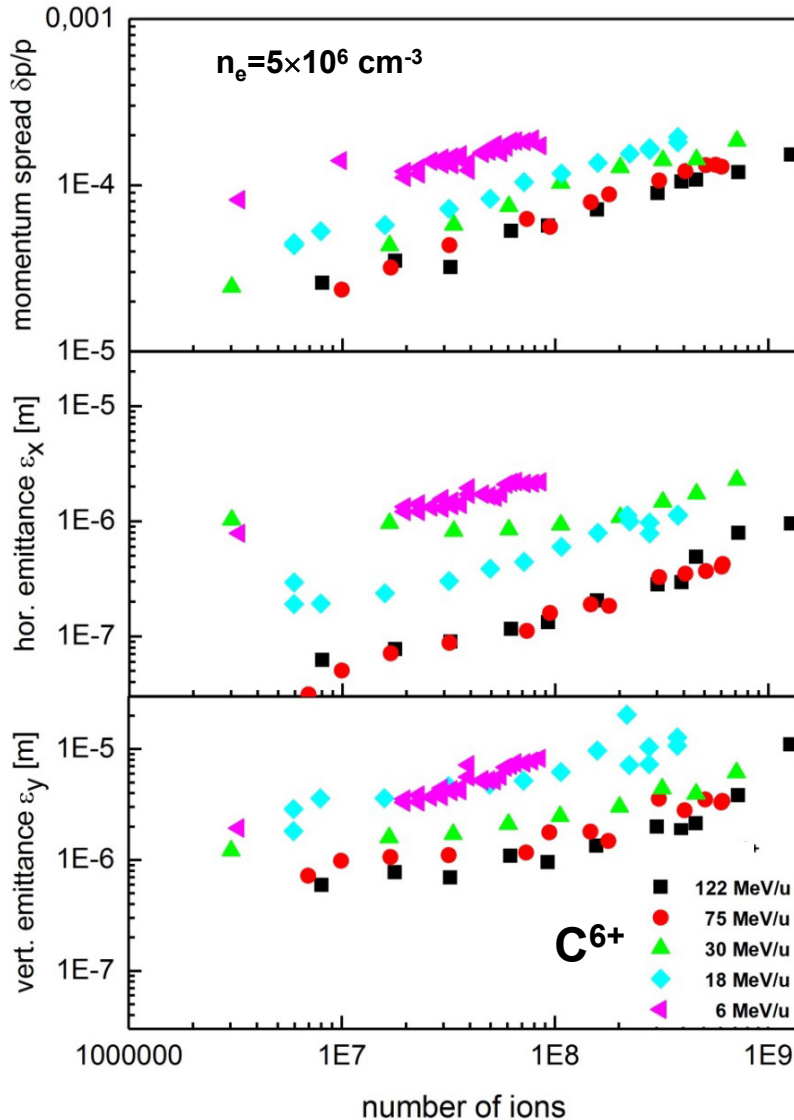
intrabeam scattering (no cooling during deceleration),
larger emittance of bunched beam
adiabatic emittance growth ($\propto 1/\beta\gamma$)

during deceleration the beam is also approaching the space charge dominated regime

$$\Delta Q_x = \frac{r_p Z^2 N g}{\pi A \beta^2 \gamma^3 B (\epsilon_x + \sqrt{\epsilon_x \epsilon_y} Q_x / Q_y)}$$

$$\Delta Q_x \simeq \frac{r_p Z^2 N}{2\pi A \beta^2 \gamma^3 B \epsilon_x},$$

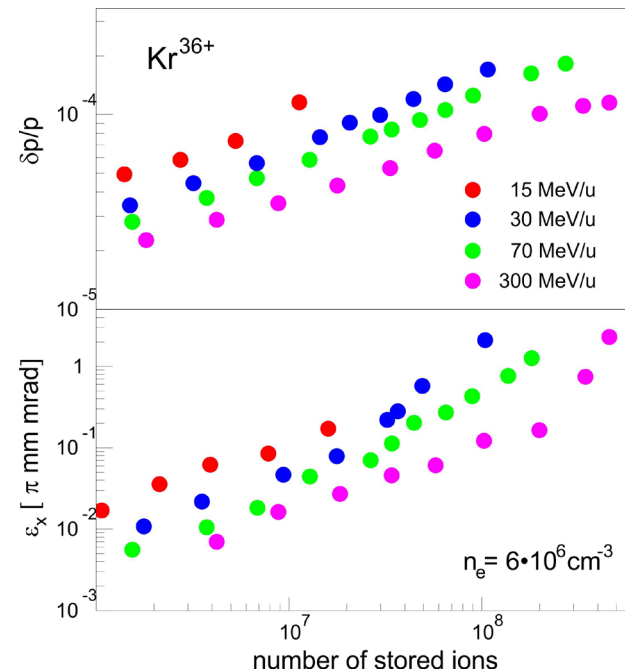
$$g = 1, \epsilon_x = \epsilon_y, Q_x = Q_y$$



Intrabeam Scattering rate

$$\tau_{IBS}^{-1} = \frac{Q^4 e^4}{(Am_i)^2} \cdot \frac{N}{C \varepsilon_h \varepsilon_v \delta p / p} \cdot \frac{1}{(\gamma^4 \beta^3 c^3)} \cdot 4\pi L_C^{IBS}$$

in case of equipartitioning:
emittance increases \approx proportional $\beta\gamma$

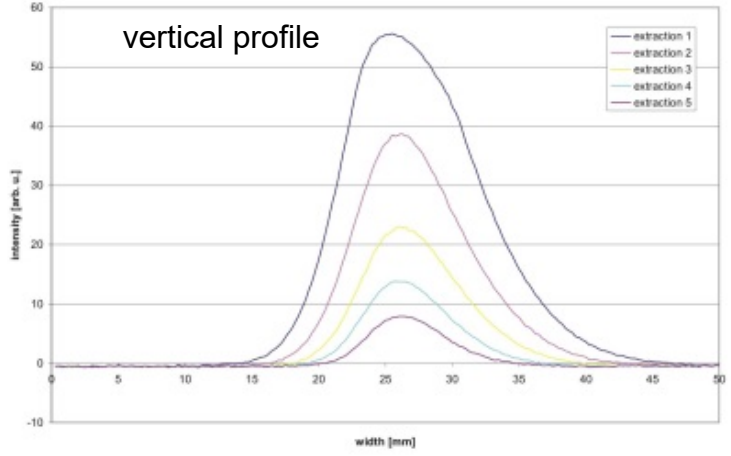
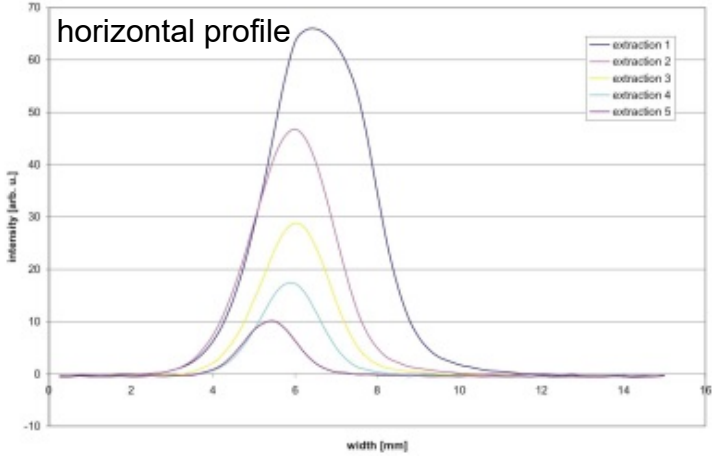


Emittance of Extracted Decelerated Beam

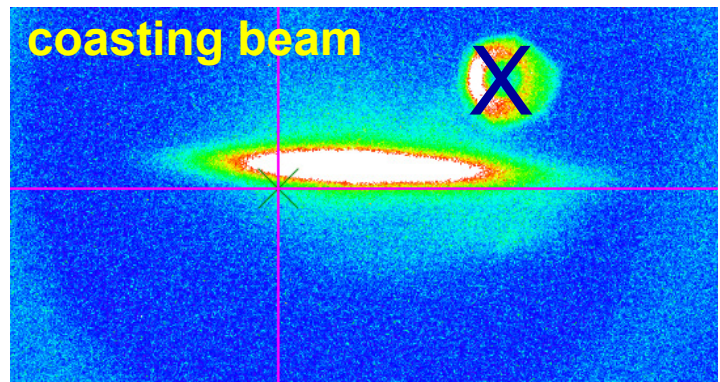
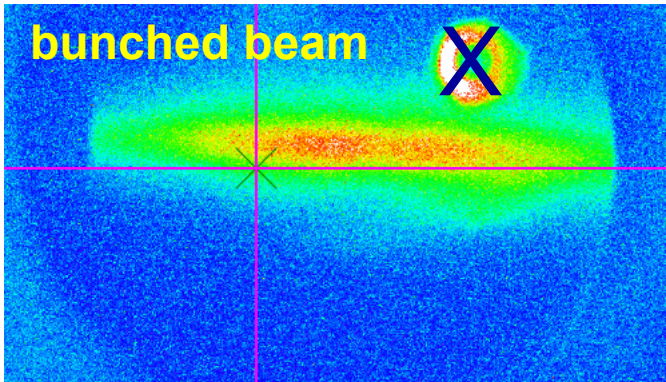


fast extracted beam
 Ca^{20+} 4 MeV/u

successive extractions with decreasing intensity



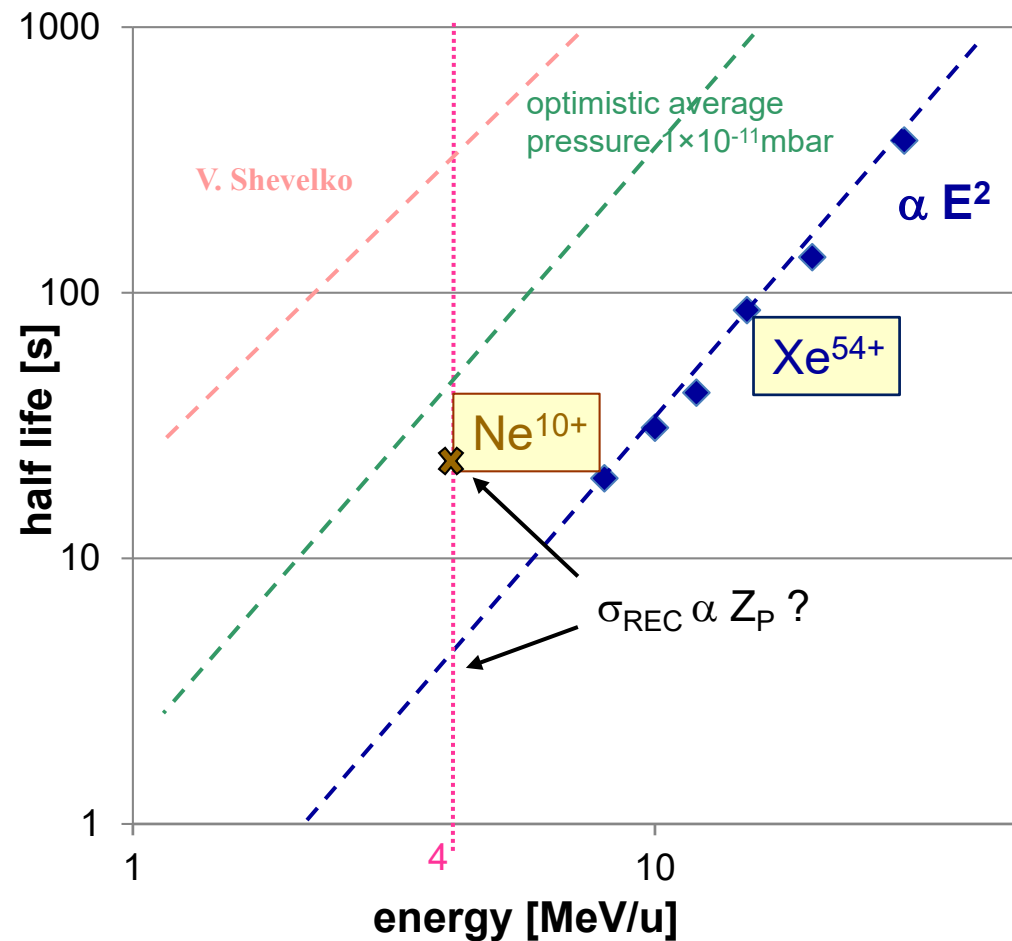
fast extracted beam
 Ni^{28+} 4 MeV/u
about 1×10^7 ions



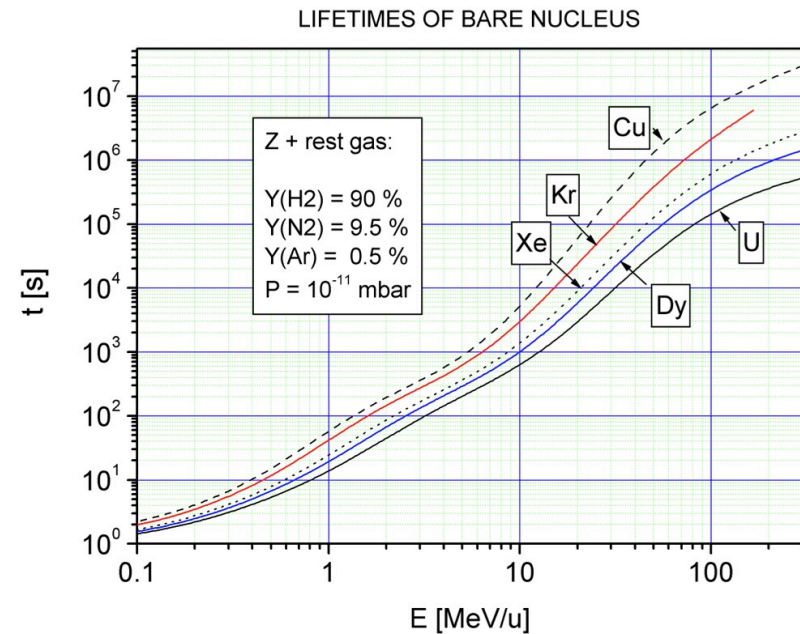
the increase of the beam emittance with intensity due to IBS reflects in the beam size of the extracted cooled beam

Low Energy Beams – Vacuum

measurement of beam half life with an average pressure of $p \approx 1 \times 10^{-10}$ mbar



simulation for CRYRING V. Shevelko



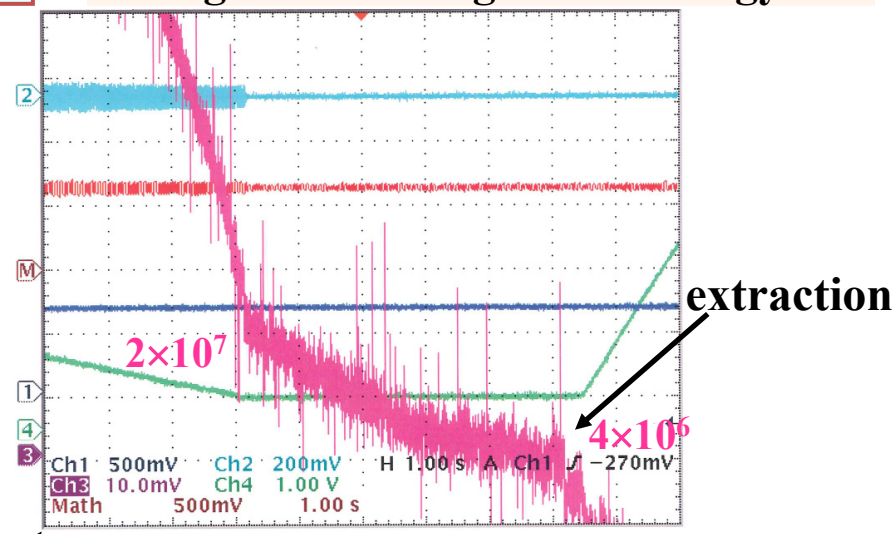
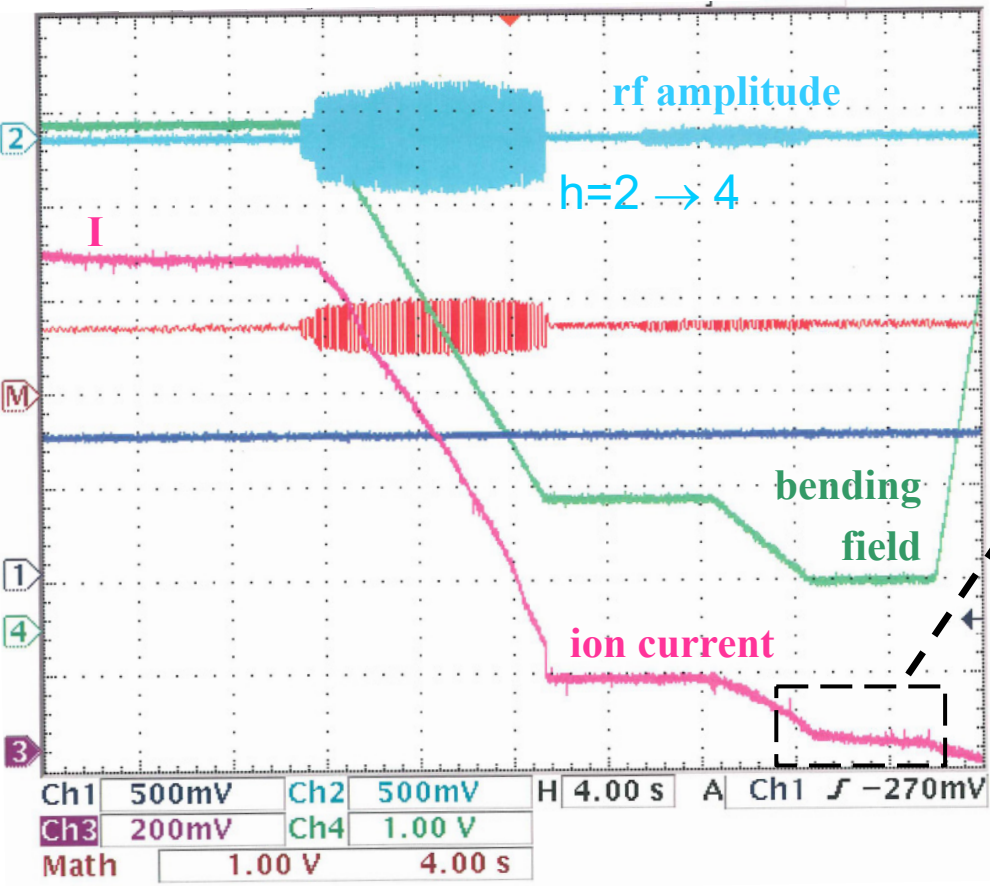
lifetime in the residual gas and cooling time can be of similar value
 \Rightarrow excellent vacuum and fast cooling are crucial for low energy highly charged ions

Deceleration to 4 MeV/u

Ni²⁸⁺ 400 → 30 → 4 MeV/u
1100 μA → 180 μA → 25 μA
45% 37 %

cycle time
45 s

Main losses:
End of ramp
Storage and cooling at low energy

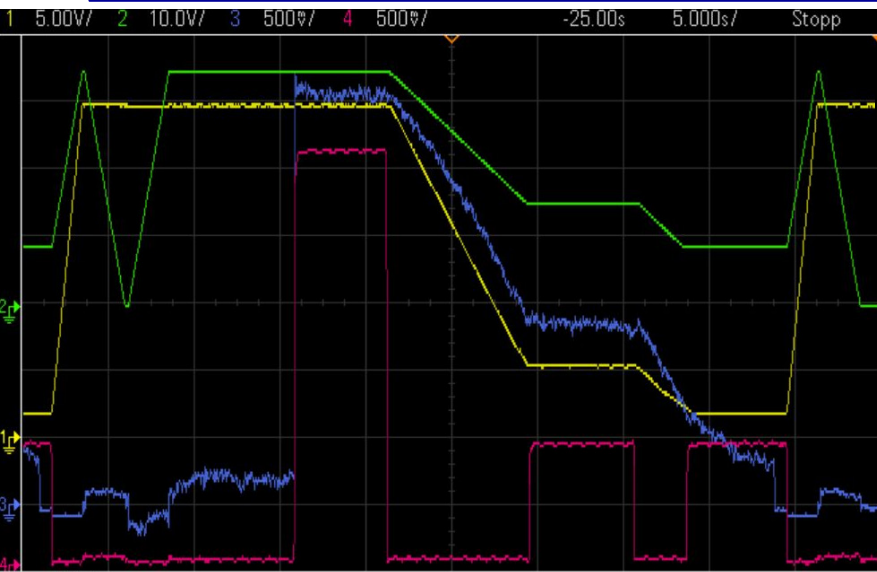


beam half life (vacuum dominated):
30 MeV/u: $T_{1/2} \approx 480$ s
4 MeV/u: $T_{1/2} \approx 2$ s

ESR Deceleration Cycle 2022

Au⁷⁸⁺: 145 → 30 → 10 MeV/u

with fast extraction for CRYRING@ESR



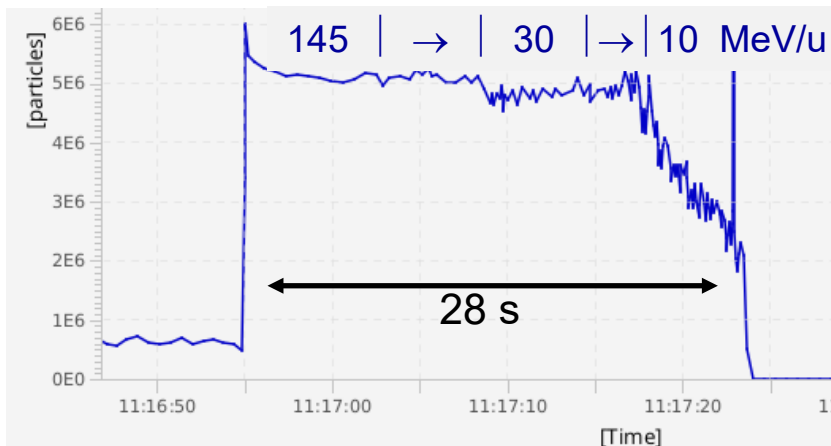
high voltage of electron cooler (electron energy)

electron cooling is applied at all three plateaus

dipole field strength

ion current

electron current (300/100 mA)

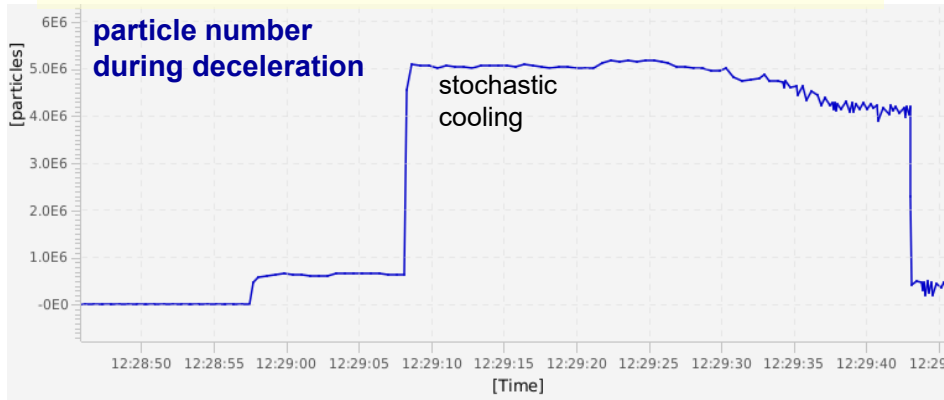


**For intensity below 10⁷
particle loss less than 20 % from 145 to 10 MeV/u**

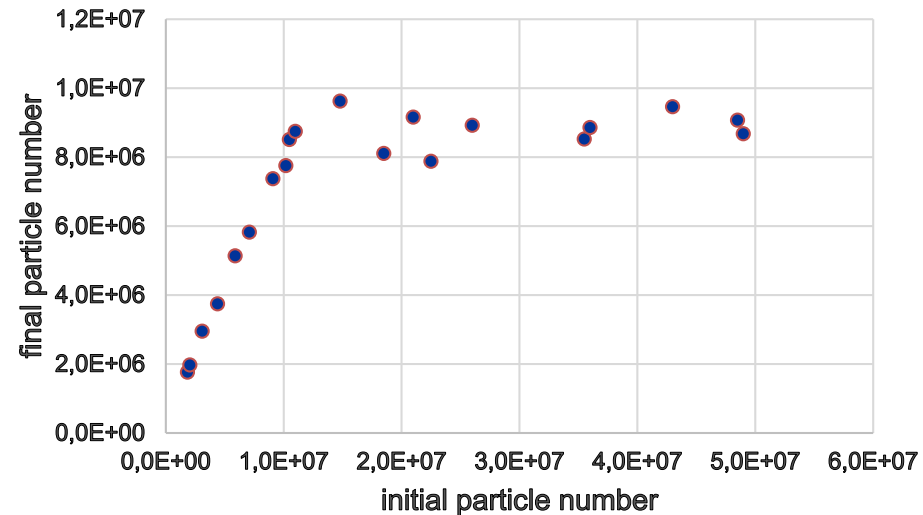
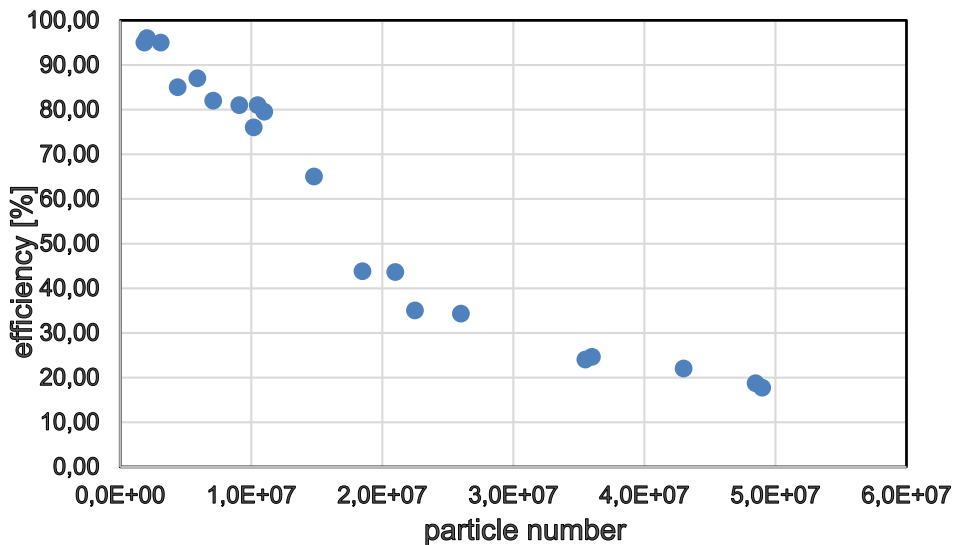
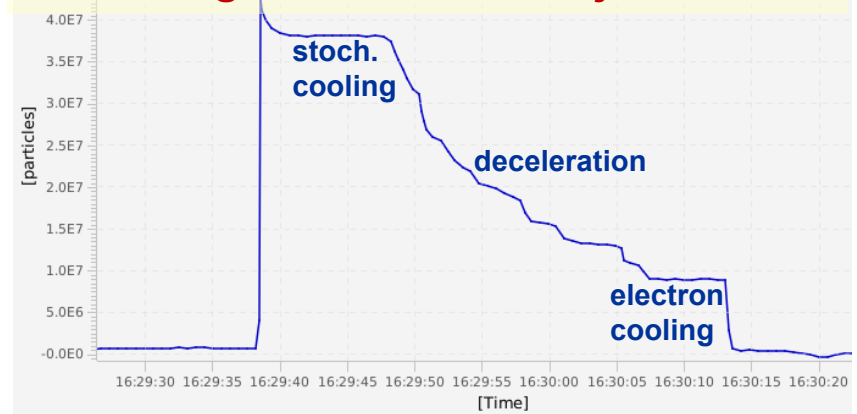
**fast losses at 10 MeV/u
(beam lifetime due to vacuum)**

deceleration cycle Bi^{83+} from 400 to 30 MeV/u

loss-free deceleration of some 10^6 ions



increasing loss above 10^7 injected ions



Space Charge Limit

Space charge limit due to incoherent tune shift

$$\Delta Q_x = \frac{r_p Z^2 N g}{\pi A \beta^2 \gamma^3 B (\epsilon_x + \sqrt{\epsilon_x \epsilon_y} Q_x / Q_y)}$$
$$\Delta Q_x \simeq \frac{r_p Z^2 N}{2\pi A \beta^2 \gamma^3 B \epsilon_x}, \quad g = 1, \epsilon_x = \epsilon_y, Q_x = Q_y$$

at 4 MeV/u, for $\Delta Q_x = 0.1$ and $\epsilon_x = 1$ mm mrad:

$$N \leq 8.7 \times 10^8 \frac{A}{Z^2} \cdot B$$

Coasting (B=1):

$$\text{Ne}^{10+}: 1.8 \times 10^8$$

$$\text{Kr}^{36+}: 5.8 \times 10^7$$

$$\text{U}^{92+}: 2.5 \times 10^7$$

Bunched ($\leq 1 \mu\text{s}$, B=1/5):

$$\text{Ne}^{10+}: 4 \times 10^7$$

$$\text{Kr}^{36+}: 1 \times 10^7$$

$$\text{U}^{92+}: 5 \times 10^6$$

For bunched beam the intensity limit is decreased by the bunching factor.

Experimental observation: with rf amplitude 100 V (h=1) the bunching factor is $B \approx 1/20$.

require precise and simultaneous measurement of:

intensity

bunching factor

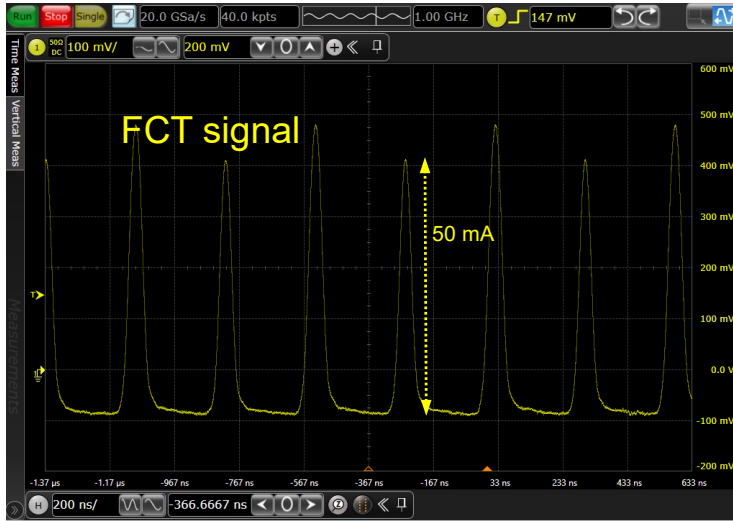
orbit

tune

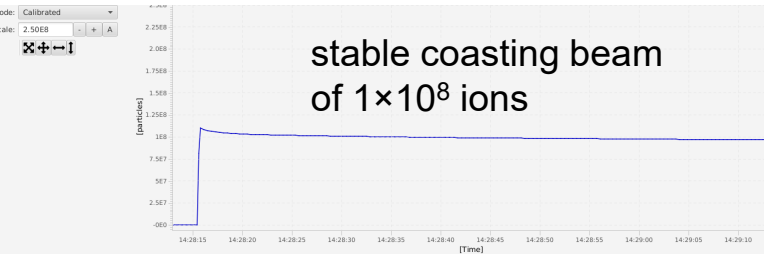
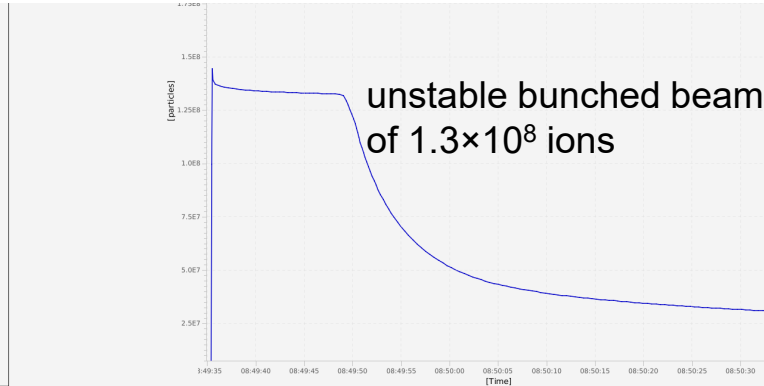
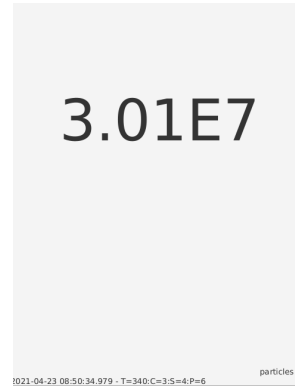
transverse emittances

....and a lot of machine development time

Intensity Issues during Deceleration

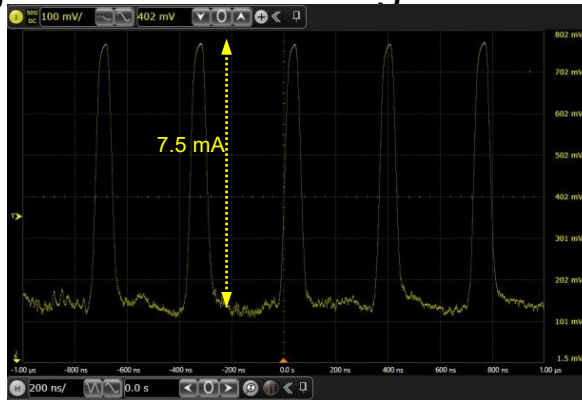


two cooled bunches U^{89+} at 300 MeV/u
with peak current 50 to 60 mA

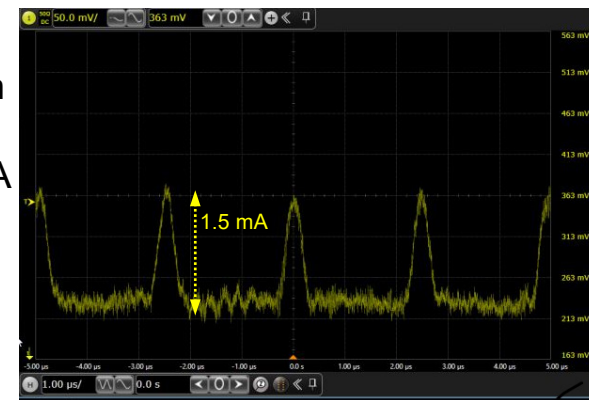


bunches during deceleration of Aq^{47+}

4 cooled bunches
at 30 MeV/u
dc current 0.55 mA

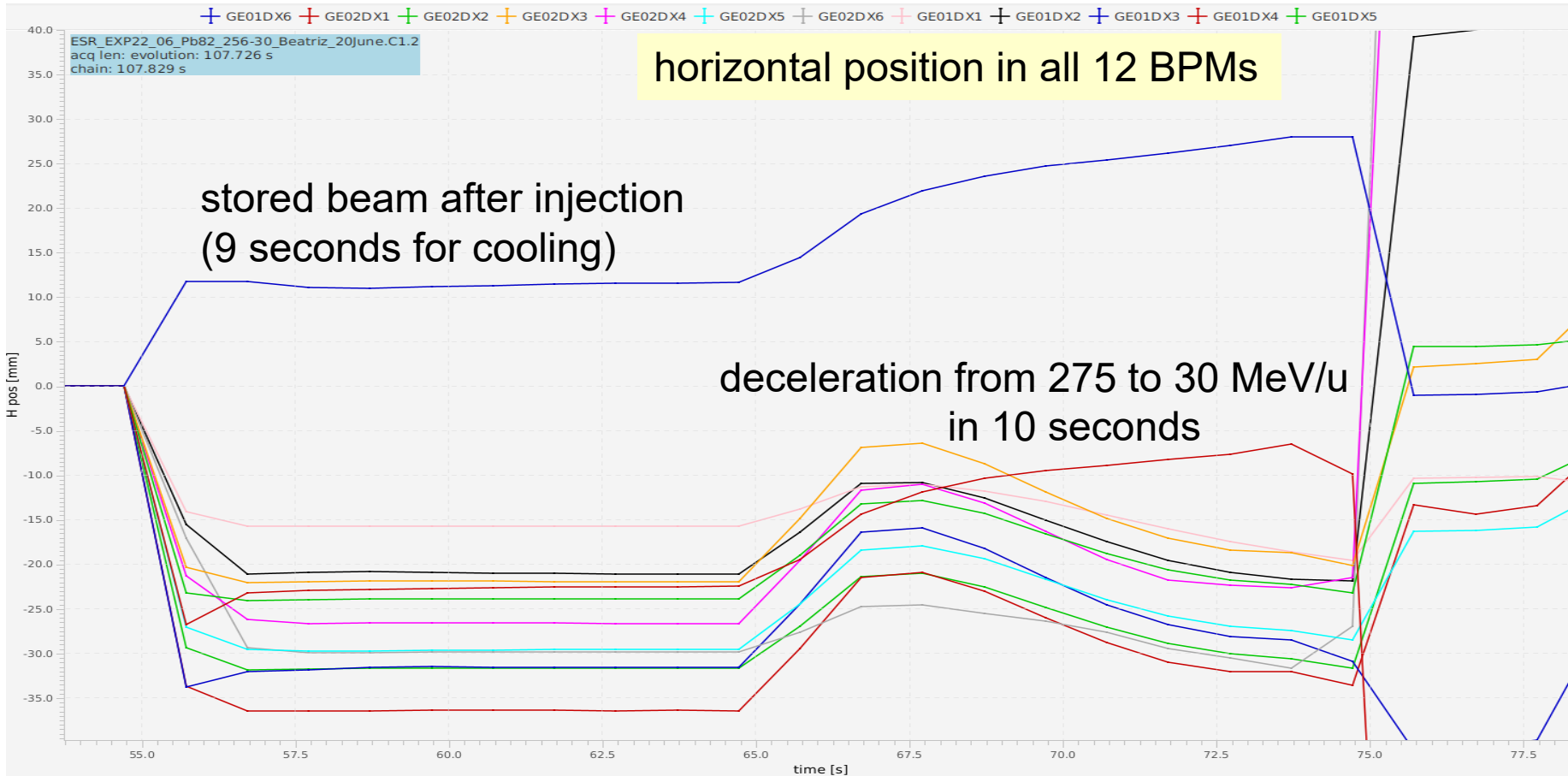


single cooled bunch
at 10 MeV/u
dc current ~ 0.05 mA



Orbit during Deceleration

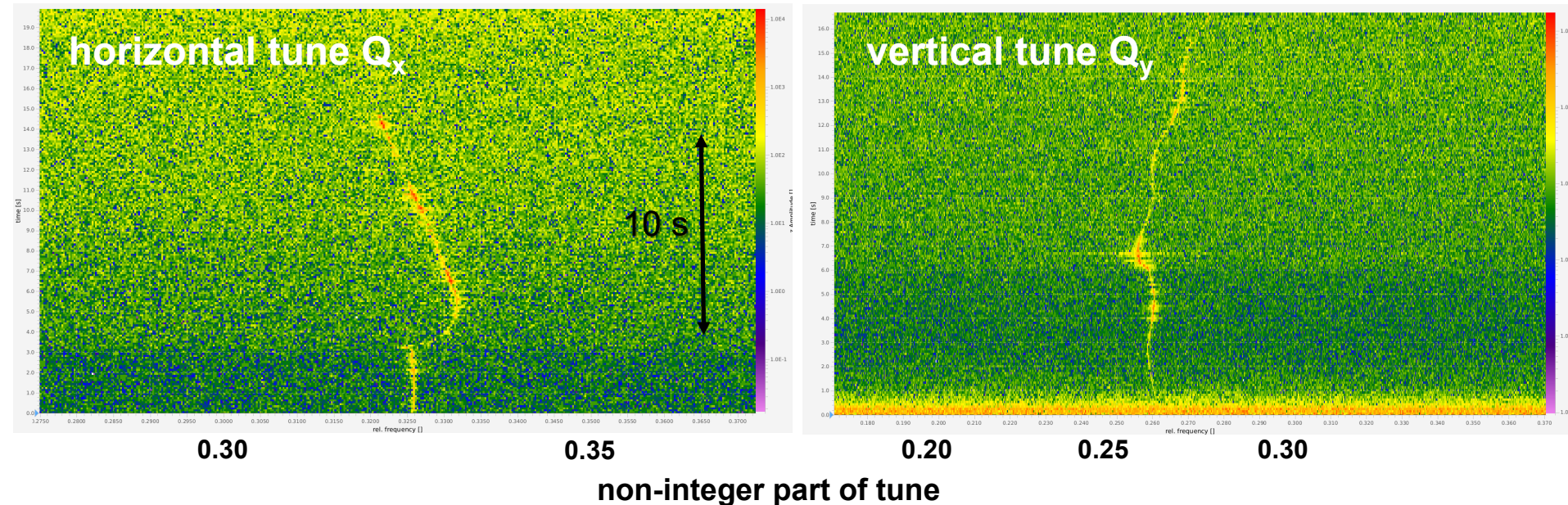
change of position during deceleration cycle Pb⁸²⁺ from 275 to 30 MeV/u



Tune during Deceleration

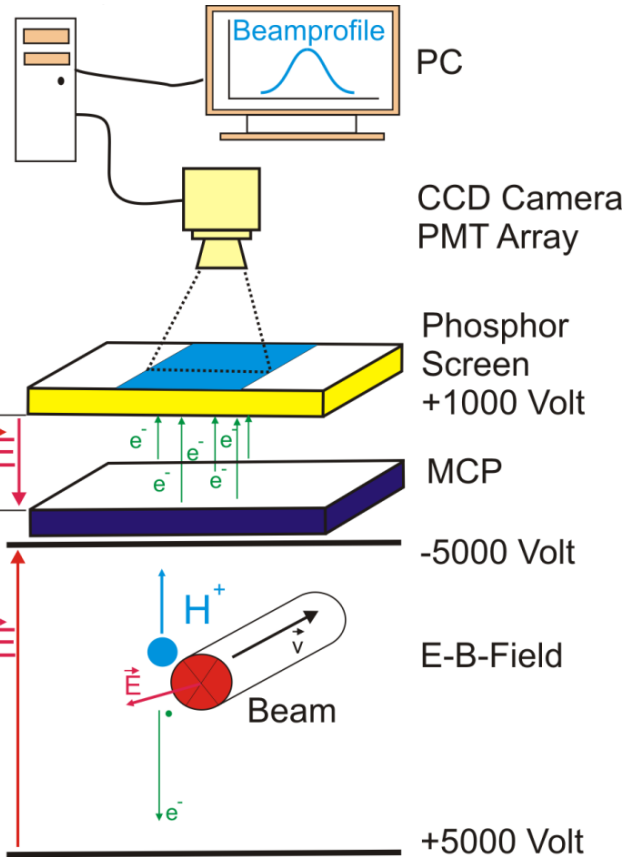
deceleration of Ni^{28+} from 45 to 4 MeV/u (for HITRAP)

excitation of bunched beam with transverse exciter (synchronized with rf frequency)
intensity: a few 10^7 particles

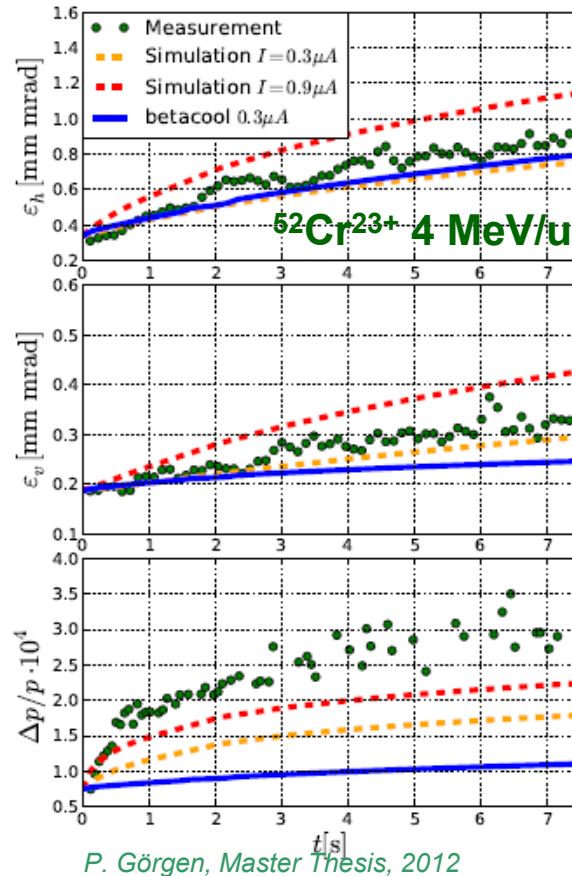


Transverse Beam Emittance

Residual Gas Ionization Beam Profile Monitor



Growth of Beam Emittance due to Intrabeam Scattering

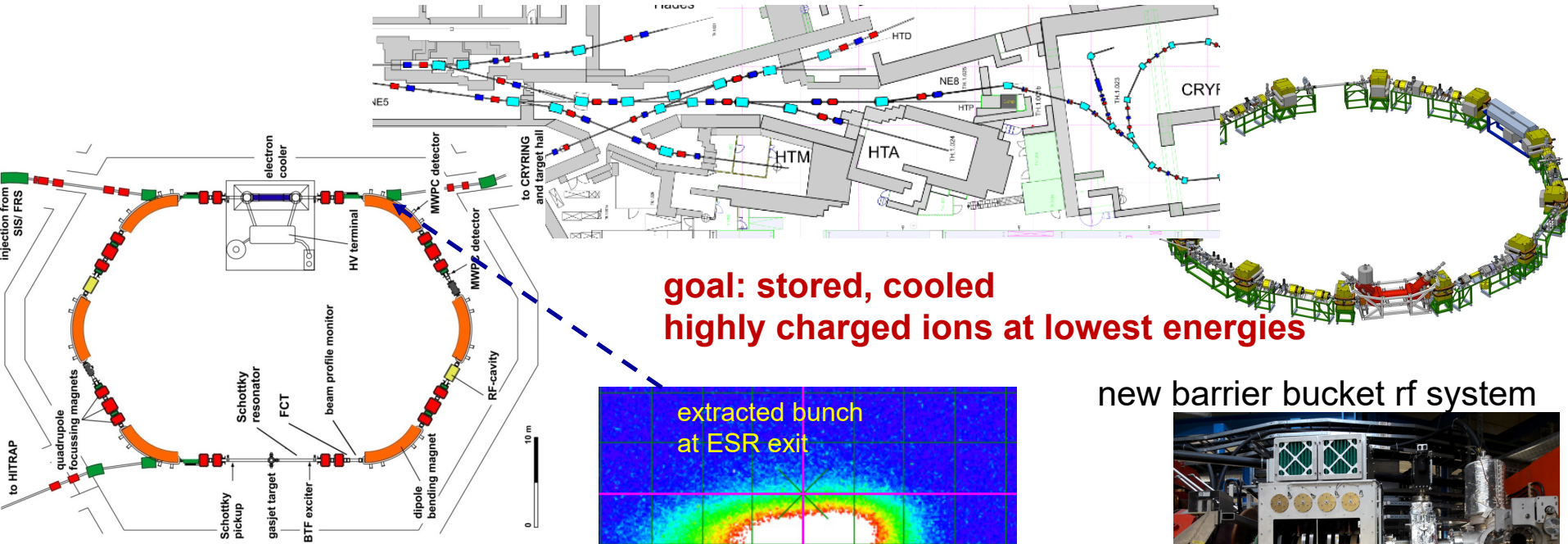


stronger (faster) cooling results in faster IBS heating

for highly charged ions IBS is expected to scale with Q^4/A^2
 \Rightarrow highly charged uranium heats up 10 times faster than Cr^{23+}

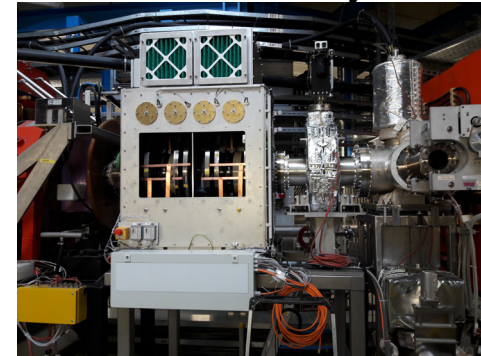
Similar observations are reported from LEIR at CERN with Pb^{53+} at 4.2 MeV/u: heating time of order of 10 ms

Fast Extraction from ESR to CRYRING@ESR

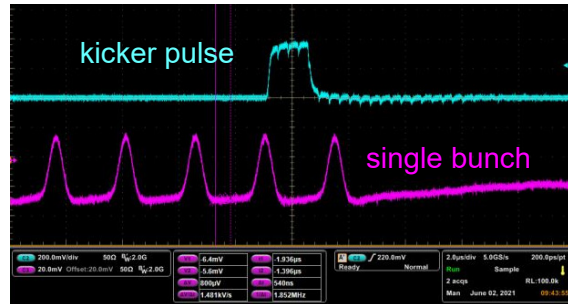
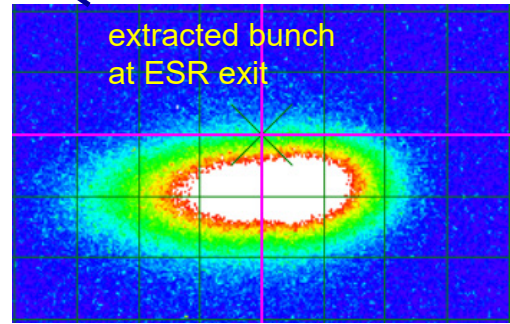


**goal: stored, cooled
highly charged ions at lowest energies**

new barrier bucket rf system



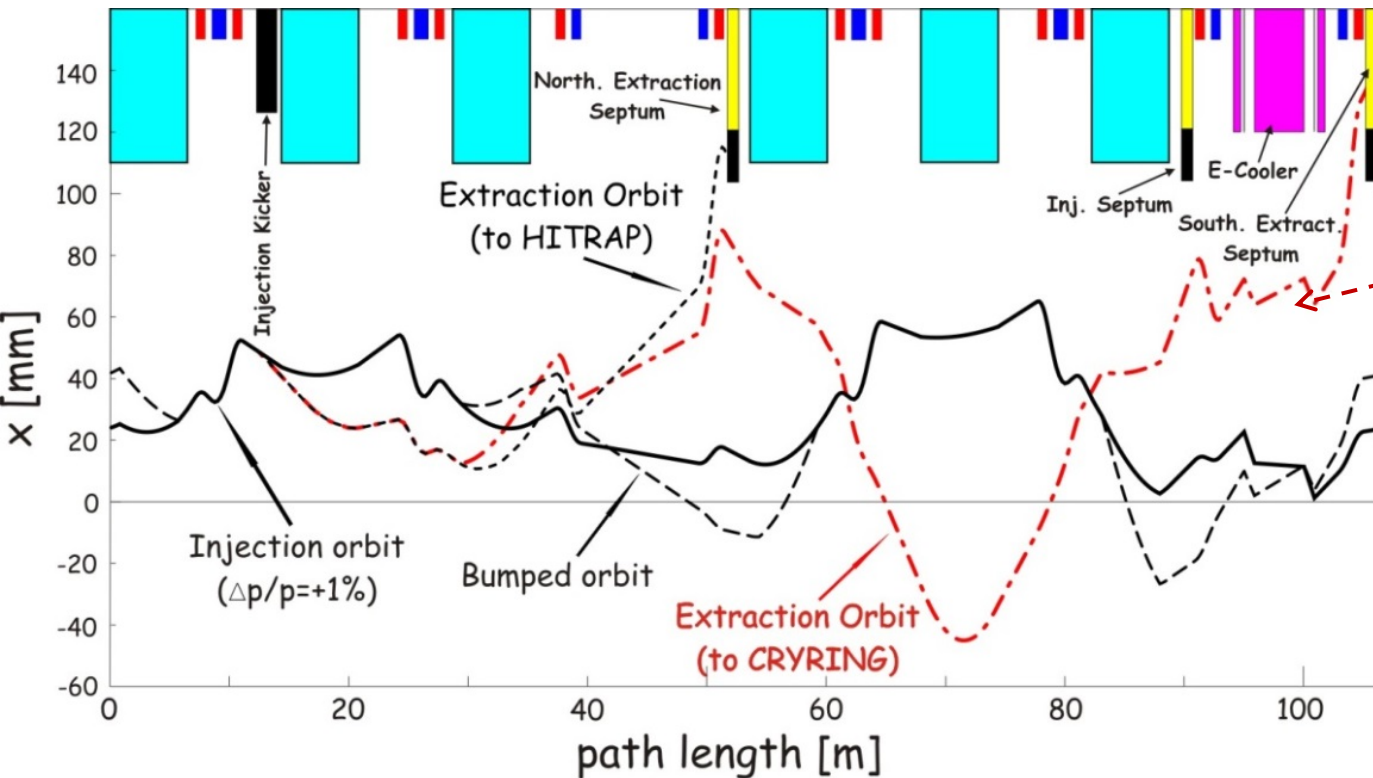
ESR decelerates cooled highly charged ions or rare isotope beams from 400 MeV/u to 4-14 MeV/u



single cooled bunch
before extraction from ESR
compressed by
new barrier bucket rf system

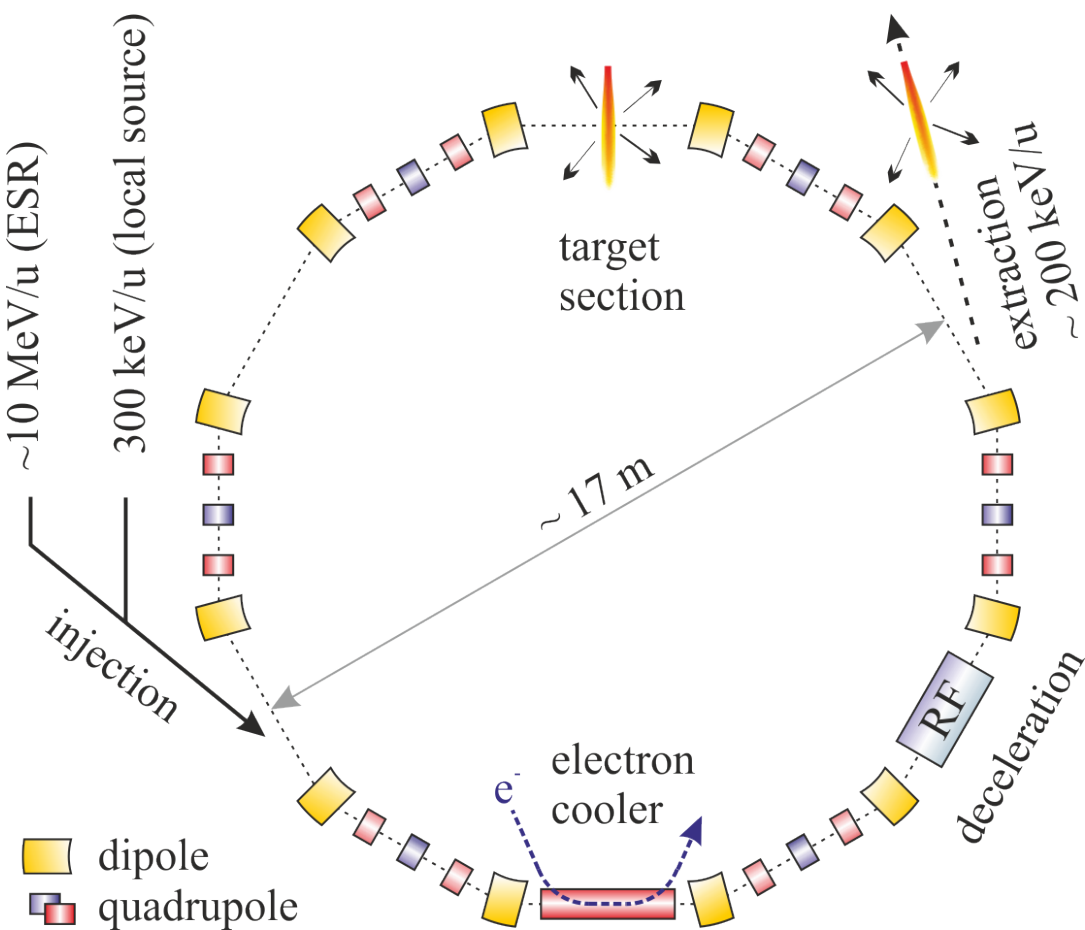
Fast Beam Extraction from ESR to CRYRING

fast beam extraction from ESR to CRYRING is employing a sophisticated scheme with a distorted closed orbit for the circulating beam before extraction to CRYRING
consequence: cooling is difficult during extraction, risk of unacceptable heating and emittance growth, particularly for highly charged ions
(expensive remedy: additional kicker in the northern arc)

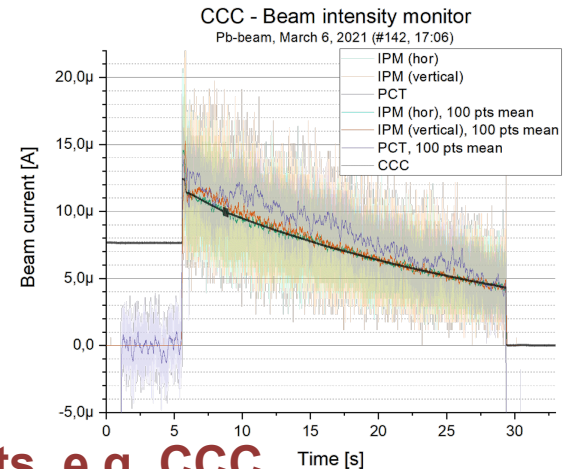


angle in the electron cooling section hampers good cooling

CRYRING@ESR Parameters



Circumference	54.17 m (ESR/2)
Vacuum pressure	10 ⁻¹¹ -10 ⁻¹² mbar
Ion energy	< 300 keV/u - 14 MeV/u
Rigidity for ions	0.054 - 1.44 Tm
Magnet ramping	1 T/s (4 T/s, 7 T/s)
Stand-alone operation	local ion beam (300 keV/u, q/A > 0.25)
Beam injection	multiturn and fast
Beam extraction	slow and fast



test bed for new controls and advanced components, e.g. CCC

Deceleration is a Complex Interplay



initial or intermediate cooling (stochastic \leftrightarrow electron cooling)

determines cooling time and beam emittances

ramp rate and consequently cycle time

is linked to orbit and tune variation which determine losses

is linked to rf amplitude which determines bunching factor (local beam current)

determines losses due to vacuum

can be limited by power supplies (in particular, if electron cooling is employed)

intermediate and final cooling

damps unwanted growth of emittance as a result of deceleration

can result in space charge effects (at low energy)

cooling time should be shorter than beam lifetime

The optimization of a deceleration cycle is aiming at balancing the beneficial and adverse effects of the various deceleration parameters

Thanks to

R. Heß, R. Joseph, S. Litvinov, B. Lorentz, U. Popp, J. Rossbach

F. Herfurt, Z. Andelkovic and the Decelerators team

and all technical departments supporting storage ring operation