

Electrodisintegration of ^{16}O and the Rate Determination of the Radiative Alpha Capture on ^{12}C at Stellar Energies

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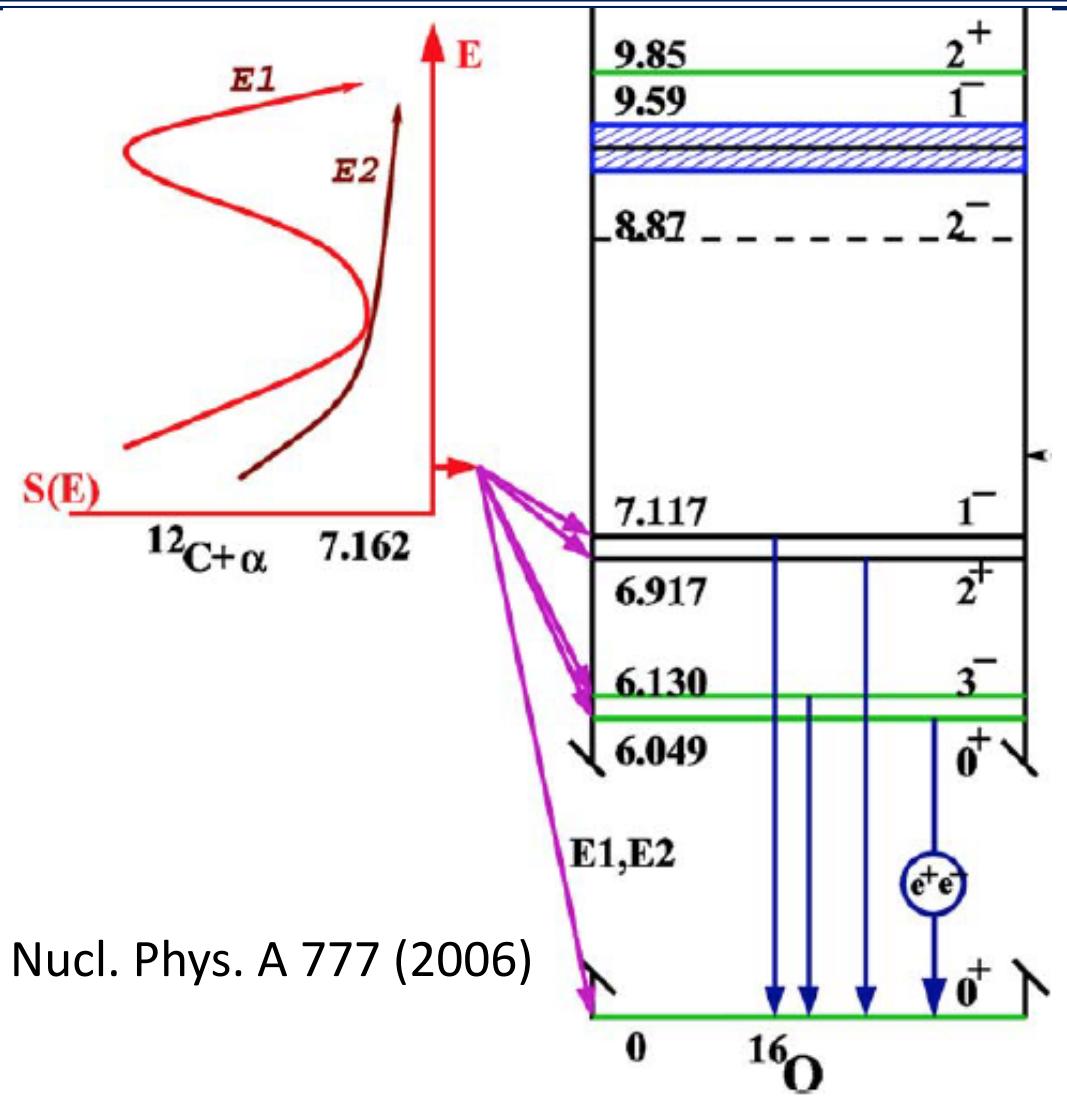
$^{12}\text{C}/^{16}\text{O}$ abundance

- Helium burning stage: $3\alpha \rightarrow ^{12}\text{C} + \gamma$, and $\alpha + ^{12}\text{C} \rightarrow ^{16}\text{O} + \gamma$
- At $T \simeq 2 \cdot 10^8$ K, Gamow window for $\alpha + ^{12}\text{C}$ is around $E_G \simeq 300$ keV and at the moment the rate is known with an uncertainty between 20% and 30%
- Affects evolution of massive stars -> nucleosynthesis of heavier elements
- White dwarfs: ignition of super nova type Ia
- End of stars: ^{16}O rich star - black hole, ^{12}C rich star – neutron star

T. A. Weaver and S. E. Woosley, Phys. Rep. 227 (1993) 335.

$\alpha + {}^{12}\text{C} \rightarrow {}^{16}\text{O} + \gamma$ cross section around E_G

- Large Coulomb barrier $\rightarrow \sigma \simeq 10^{-5}$ pb (direct measurement is not feasible)
- Around E_G the cross section is dominated by two components:
 - E1 component, $J^\pi = 1^-$: subthreshold state at 7.117 MeV and broad resonance at 9.59 MeV
 - E2 component, $J^\pi = 2^+$: subthreshold state at 6.917 MeV and narrow resonance at 9.85 MeV



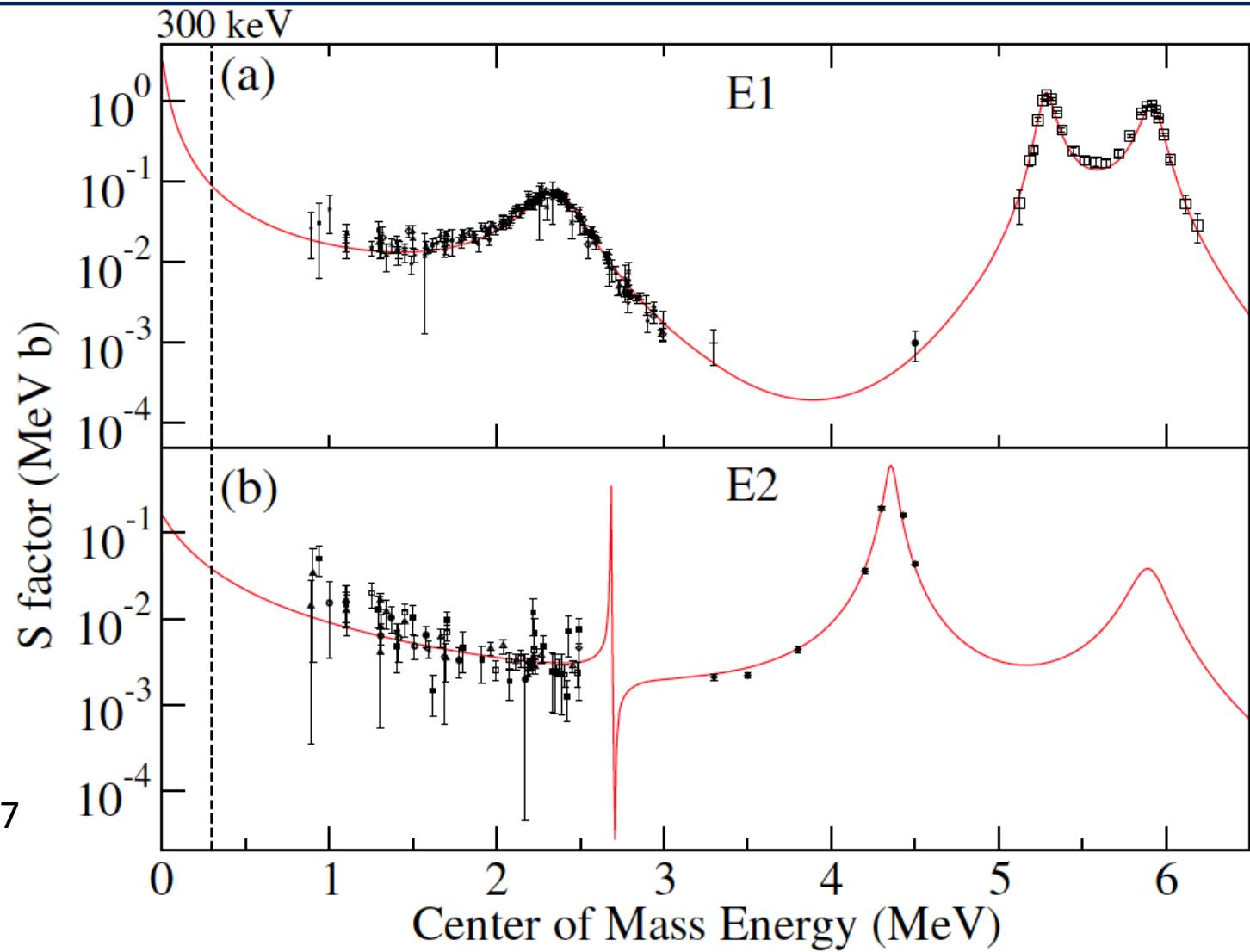
L.R. Buchmann, C.A. Barnes, Nucl. Phys. A 777 (2006)

$\alpha + {}^{12}\text{C} \rightarrow {}^{16}\text{O} + \gamma$ S-factors

$$\sigma(E) = \frac{S(E)}{E} e^{-2\pi\eta}$$

$$\eta = \frac{2\pi Z_1 Z_2 e^2}{h\nu_{rel}}$$

R. J. deBoer et al., Rev. Mod. Phys. 89, 035007
(2017) and references therein



How to extract $\sigma(E_G)$

- Direct measurements

- a) $^{12}\text{C}(\alpha, \gamma)^{16}\text{O}$; (α beam)

- angular distribution of γ is measured

- $\rightarrow S_{E1}$ and S_{E2}

- b) $\alpha(^{12}\text{C}, ^{16}\text{O})\gamma$; (^{12}C beam)

- detection of ^{16}O recoils $\rightarrow S_{\text{tot}}$

- Indirect measurements

- a) β decay of ^{16}N : $^{16}\text{O}^* \rightarrow \alpha + ^{12}\text{C}$; $\rightarrow S_{E1}$

- b) inverse reaction:

- photodisintegration of ^{16}O : $^{16}\text{O}(\gamma, \alpha)^{12}\text{C}$

- \rightarrow Bubble chamber R. J. Holt et al., (2018), arXiv:1809.10176

- \rightarrow Time project. chamber M. Gai et al., JINST 5, P12004 (2010)

- **electrodisintegration of ^{16}O : $^{16}\text{O}(e, e'\alpha)^{12}\text{C}$; THIS TALK**

I. Friščić, W. T. Donnelly and R. G. Milner, Phys. Rev. C 100, (2019) 025804

$$\frac{\sigma(\gamma + ^{16}\text{O})}{\sigma(\alpha + ^{12}\text{C})} = \frac{\mu c^2 E_\alpha^{cm}}{E_\gamma^2} \approx 42 \text{ (for } E_\alpha^{cm} = 1 \text{ MeV)}$$

Advantage of $^{16}\text{O}(\text{e},\text{e}'\alpha)^{12}\text{C}$

- inverse reaction: larger cross section than direct reaction
- new generation of e^- energy recovery linear (ERL) accelerators with $I \geq 10 \text{ mA}$:
MESA, Univ. of Mainz, Germany, F. Hug et al., Proc. of LINAC'16 28, 313 (2017).
CBETA, Cornell Univ., USA, D. Trbojevic et al., Proc. of IPAC'17 8, 1285 (2017).
- oxygen cluster gas-jet target with thickness $> 10^{18} \text{ atoms/cm}^2$
MAGIX, Univ. of Mainz, Germany, S. Grieser et al., Nucl. Inst. Meth. Phys. Res. A 906, 120 (2018).

Luminosity $> 10^{35} \text{ 1/(cm}^2\text{s)}$

Systematics from oxygen isotopes

- Oxygen isotope abundance: ^{16}O 99.757%, ^{17}O 0.038% and ^{18}O 0.205%

K. J. R. Rosman, P. D. P. Taylor, Pure Appl. Chem. 71 (1999) 1593

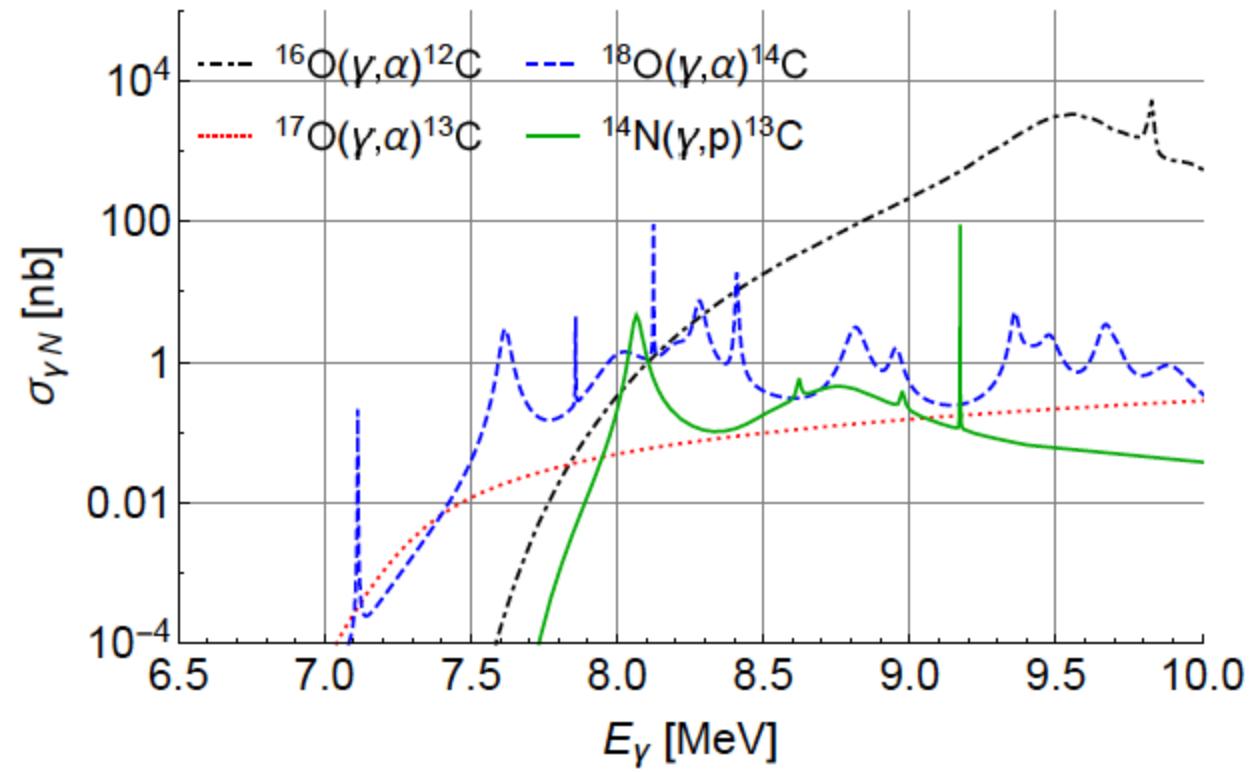
$$Q(^{16}\text{O} \rightarrow \alpha + ^{12}\text{C}) = -7.162 \text{ MeV},$$

$$Q(^{17}\text{O} \rightarrow \alpha + ^{13}\text{C}) = -6.357 \text{ MeV}$$

$$Q(^{18}\text{O} \rightarrow \alpha + ^{14}\text{C}) = -6.228 \text{ MeV}$$

- Photonuclear cross sections: natural abundance of O isotopes + depletion of ^{17}O and ^{18}O by factor 1000, and 5 ppmv for ^{14}N

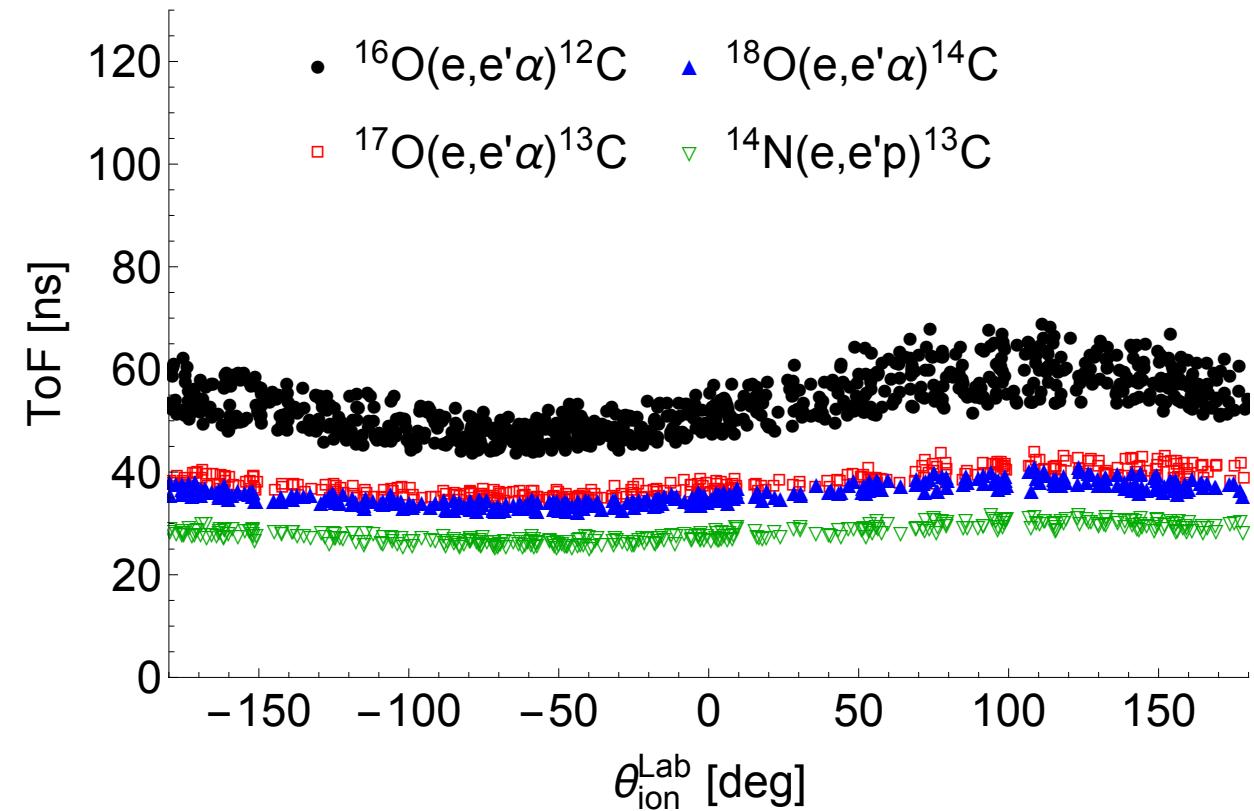
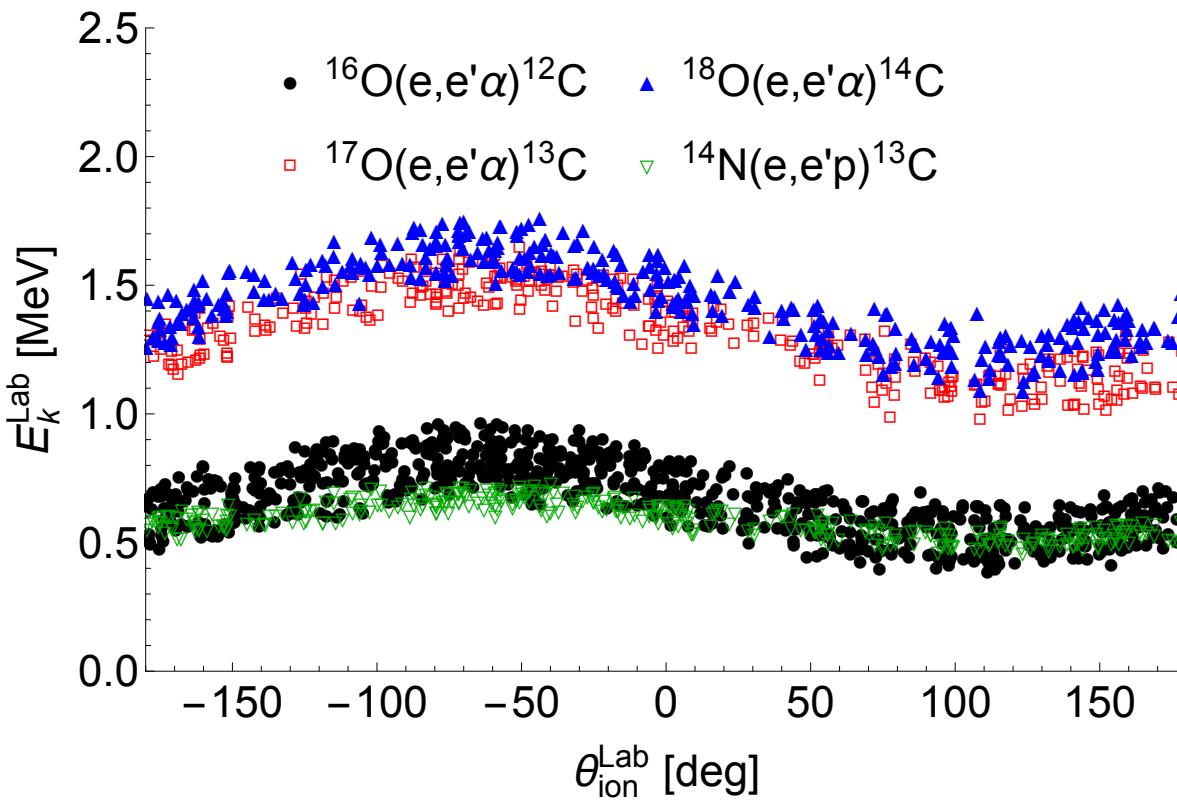
$$E_\gamma = E_\alpha^{cm} + 7.162 \text{ MeV}$$



https://wiki.jlab.org/ciswiki/index.php/Simulations_and_Backgrounds#Relevant_Theoretical_Cross_Sections

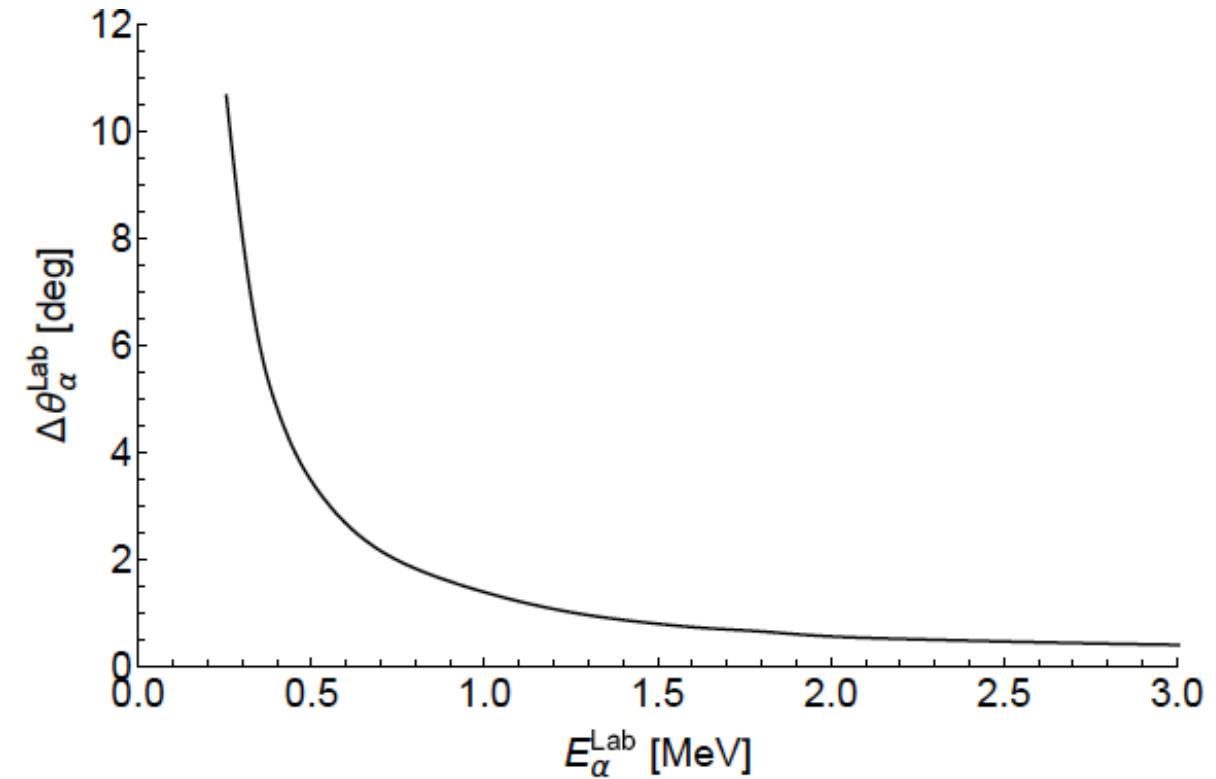
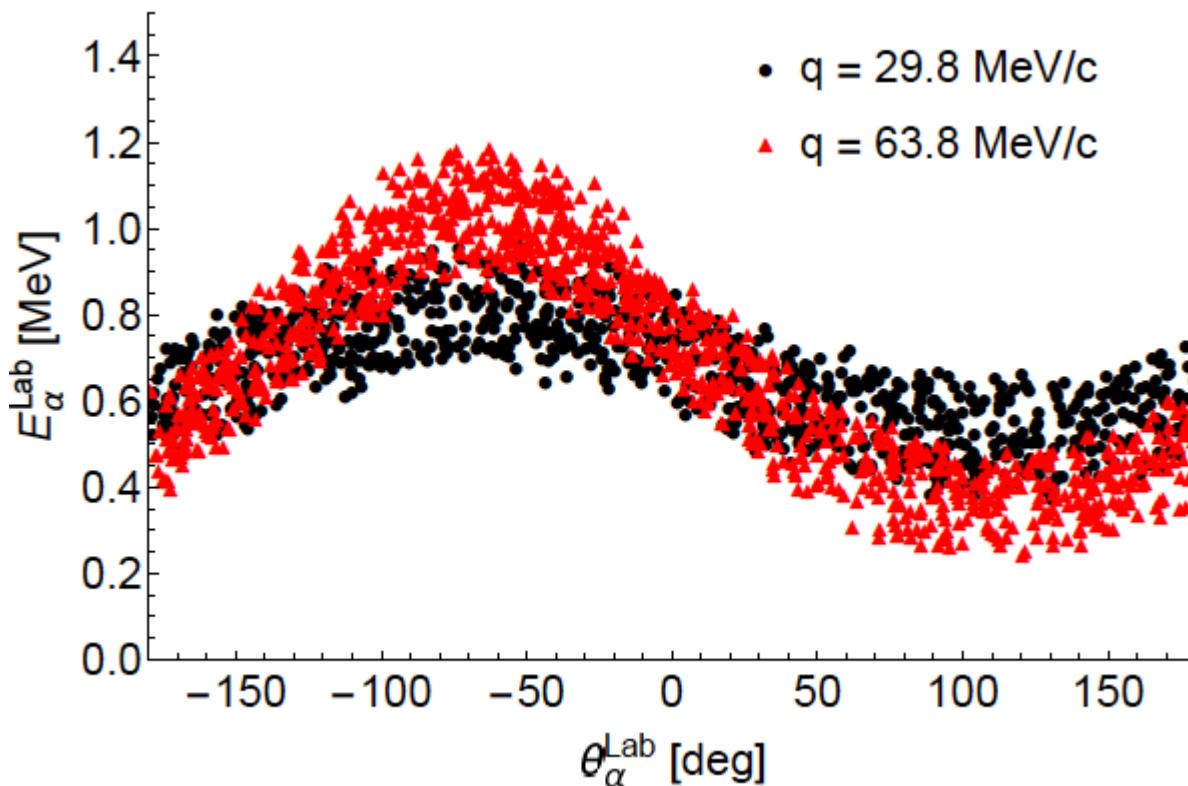
Systematics from oxygen isotopes: Solution

- SRIM simulation: energy loss of α -particles in 2 mm wide oxygen jet, with a density of $6.65 \cdot 10^{-4} \text{ g/cm}^3$, $E_e = 114 \text{ MeV}$, $\theta_e = 15^\circ$, $1.0 \leq E_\alpha^{cm} \leq 1.1 \text{ MeV}$



Virtual photon advantage

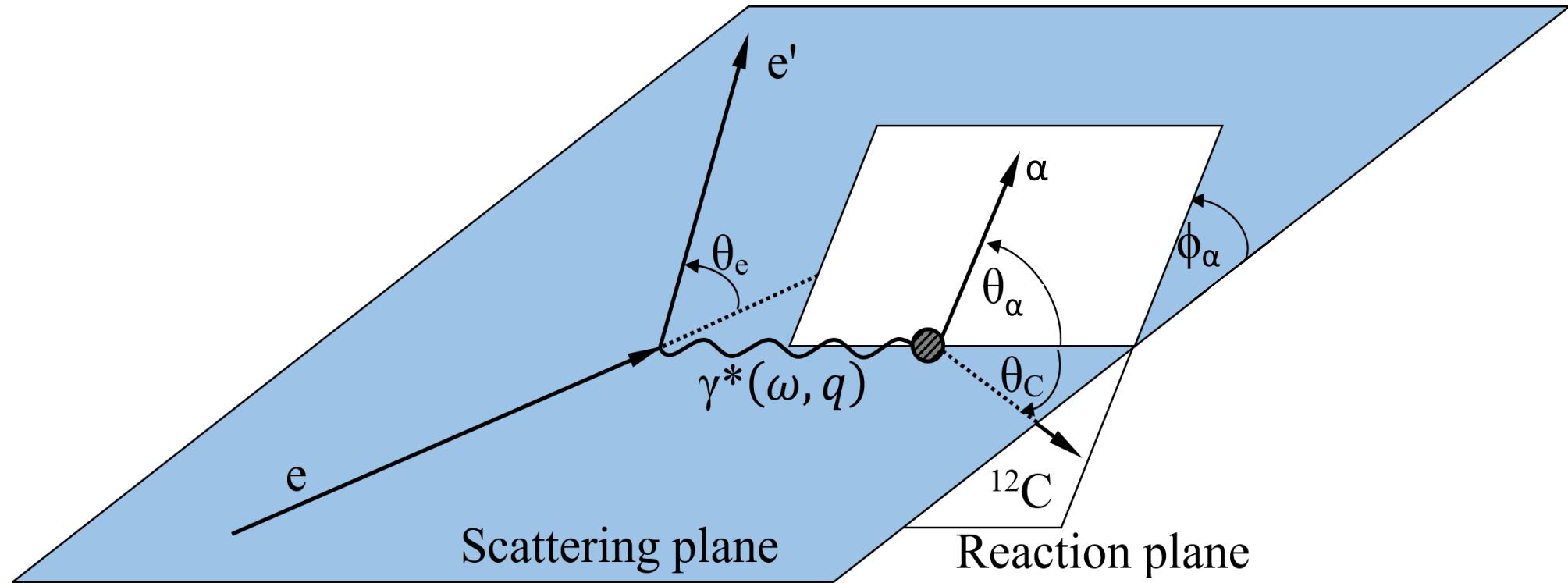
- SRIM simulation: angular spread of α -particles in 2 mm wide oxygen jet, with a density of $6.65 \cdot 10^{-4} \text{ g/cm}^3$, $E_e = 114 \text{ MeV}$, $\theta_e = 15^\circ$ and 35° , $1.0 \leq E_\alpha^{cm} \leq 1.1 \text{ MeV}$



Detection of α -particles and other ions

- Requirements:
 - measure the total energy of the α -particles to about $\sim 10\%$
 - distinguish between protons, α -particles and C-isotopes
 - measure the position to $\sim \text{mm}$ and the timing to a few ns
 - ion detection system have to be blind to scattered e^- and photons
- Options:
 - Silicon detectors → high position resolution, needs to be cooled to min. radiation damage
 - Micro-channel-plate electron (MCP) detector → good timing resolution
 - Parallel-plate avalanche counter (PPAC) → good timing resolution and position resolution
 - Time Projection Chamber → reconstruction of the ion's trajectory

Kinematics: $^{16}\text{O}(\text{e},\text{e}'\alpha)^{12}\text{C}$



The cross section formulas

- Electrodissintegration of ^{16}O :

$$\frac{d\sigma}{dE'_e d\Omega_e d\Omega_\alpha^{cm}} = \frac{M_\alpha M_{12C}}{8\pi^3 W} \frac{p_\alpha^{cm}}{(\hbar c)^3} \sigma_{Mott} (\tilde{\nu}_L R_L + \tilde{\nu}_T R_T + \tilde{\nu}_{LT} R_{LT} + \tilde{\nu}_{TT} R_{TT})$$
$$W = \sqrt{(M_{16O} + \omega)^2 - q^2} \quad E_\alpha^{cm} = W - W_{th}$$

A. S. Raskin and T. W. Donnelly, Ann. of Phys. 191 (1989)

- Direct reaction $^{12}\text{C}(\alpha, \gamma)^{16}\text{O}$:

$$\left. \frac{d\sigma}{d\Omega_\gamma^{cm}} \right|_{(\alpha,\gamma)} = \frac{M_\alpha M_{12C}}{2\pi W} \frac{p_\alpha^{cm}}{\hbar c} \frac{\alpha}{E_\gamma} R_T$$
$$W = \sqrt{M_\alpha^2 + M_{12C}^2 + 2M_{12C}E_\alpha^{lab}} \quad E_\alpha^{cm} = \frac{M_{12C}}{M_{12C} + M_\alpha} E_\alpha^{lab}$$

Response functions for $J^\pi = 0^+$ nuclei

$$R_L = P_0(\cos \theta_\alpha) \left(|t_{C0}|^2 + |t_{C1}|^2 + |t_{C2}|^2 \right)$$

$$+ P_1(\cos \theta_\alpha) \left(2\sqrt{3}|t_{C0}||t_{C1}| \cos(\delta_{C1} - \delta_{C0}) + 4\sqrt{\frac{3}{5}}|t_{C1}||t_{C2}| \cos(\delta_{C2} - \delta_{C1}) \right)$$

$$+ P_2(\cos \theta_\alpha) \left(2|t_{C1}|^2 + \frac{10}{7}|t_{C2}|^2 + 2\sqrt{5}|t_{C0}||t_{C2}| \cos(\delta_{C2} - \delta_{C0}) \right)$$

$$+ P_3(\cos \theta_\alpha) \left(6\sqrt{\frac{3}{5}}|t_{C1}||t_{C2}| \cos(\delta_{C2} - \delta_{C1}) \right)$$

$$+ P_4(\cos \theta_\alpha) \left(\frac{18}{7}|t_{C2}|^2 \right)$$

$$R_T = P_0(\cos \theta_\alpha) \left(|t_{E1}|^2 + |t_{E2}|^2 \right)$$

$$+ P_1(\cos \theta_\alpha) \left(\frac{6}{\sqrt{5}}|t_{E1}||t_{E2}| \cos(\delta_{E2} - \delta_{E1}) \right)$$

$$+ P_2(\cos \theta_\alpha) \left(-|t_{E1}|^2 + \frac{5}{7}|t_{E2}|^2 \right)$$

$$+ P_3(\cos \theta_\alpha) \left(-\frac{6}{\sqrt{5}}|t_{E1}||t_{E2}| \cos(\delta_{E2} - \delta_{E1}) \right)$$

$$+ P_4(\cos \theta_\alpha) \left(-\frac{12}{7}|t_{E2}|^2 \right)$$

$$R_{TT} = -R_T \cos(2\phi_\alpha)$$

Matrix elements and coefficients

- Multipole matrix elements ($q_0 = 1.2 \text{ fm}^{-1}$):

$$t_{EJ} = \frac{\omega}{q} \left(\frac{q}{q_0} \right)^J a'_{EJ} \left[1 + \left(\frac{q}{q_0} \right)^2 b'_{EJ}(q) \right] e^{-\left(\frac{q}{q_0} \right)^2} \quad t_{CJ} = \left(\frac{q}{q_0} \right)^J a'_{CJ} \left[1 + \left(\frac{q}{q_0} \right)^2 b'_{CJ}(q) \right] e^{-\left(\frac{q}{q_0} \right)^2}$$

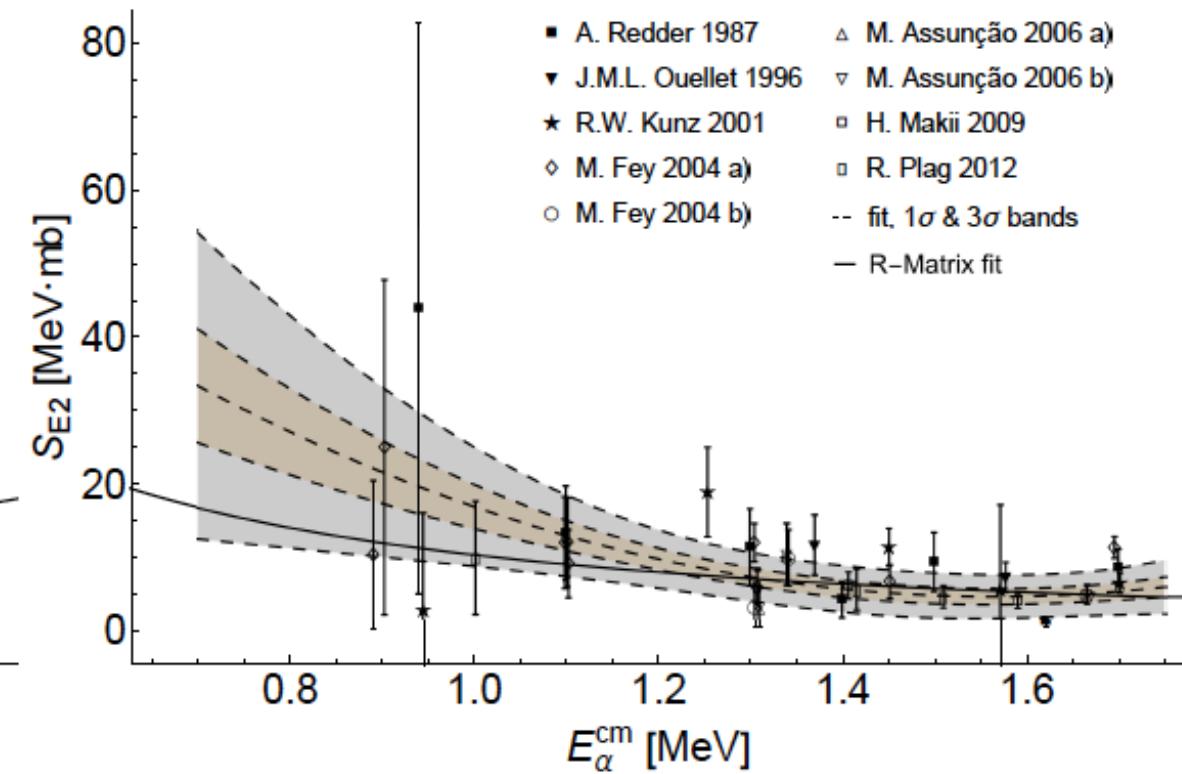
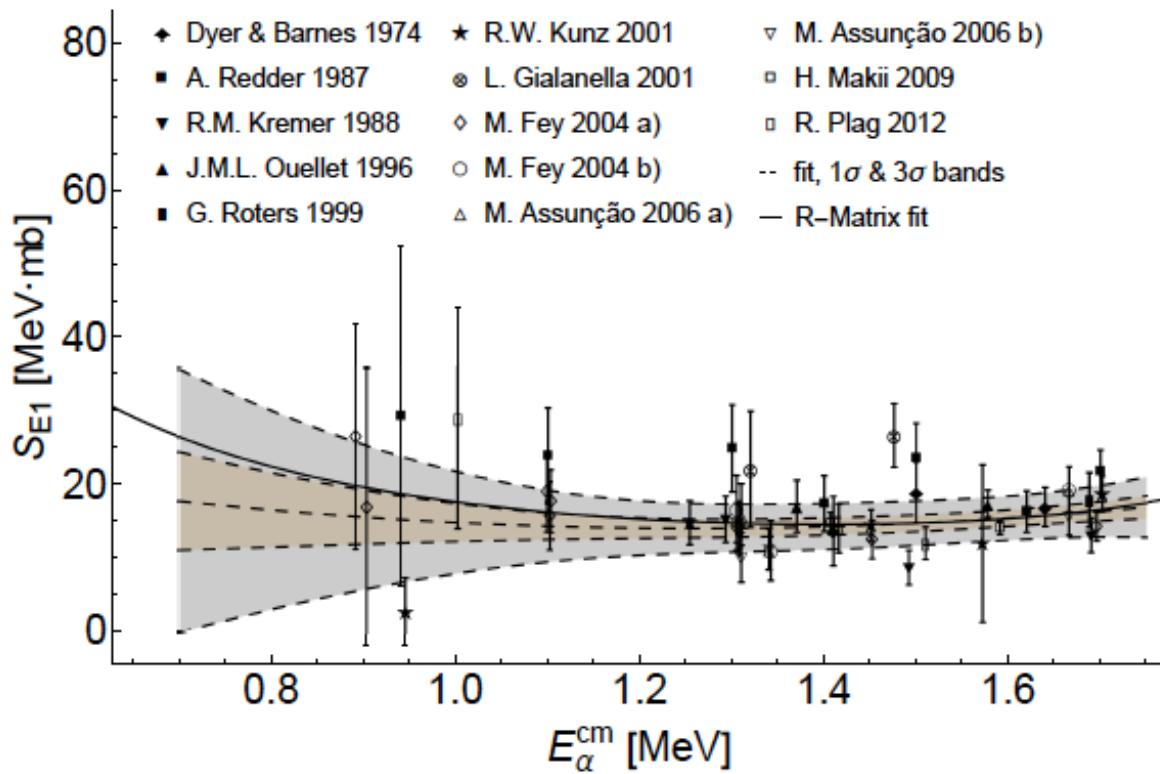
(t_{C0} leading dependence cannot occur due to orthogonality of initial and final state)

- Long wavelength limit ($q \rightarrow 0$) and continuity equation:

$$t_{EJ} \rightarrow -\sqrt{\frac{J+1}{J}} \left(\frac{\omega}{q} \right) t_{CJ} \quad a'_{EJ} = -\sqrt{\frac{J+1}{J}} a'_{CJ}$$

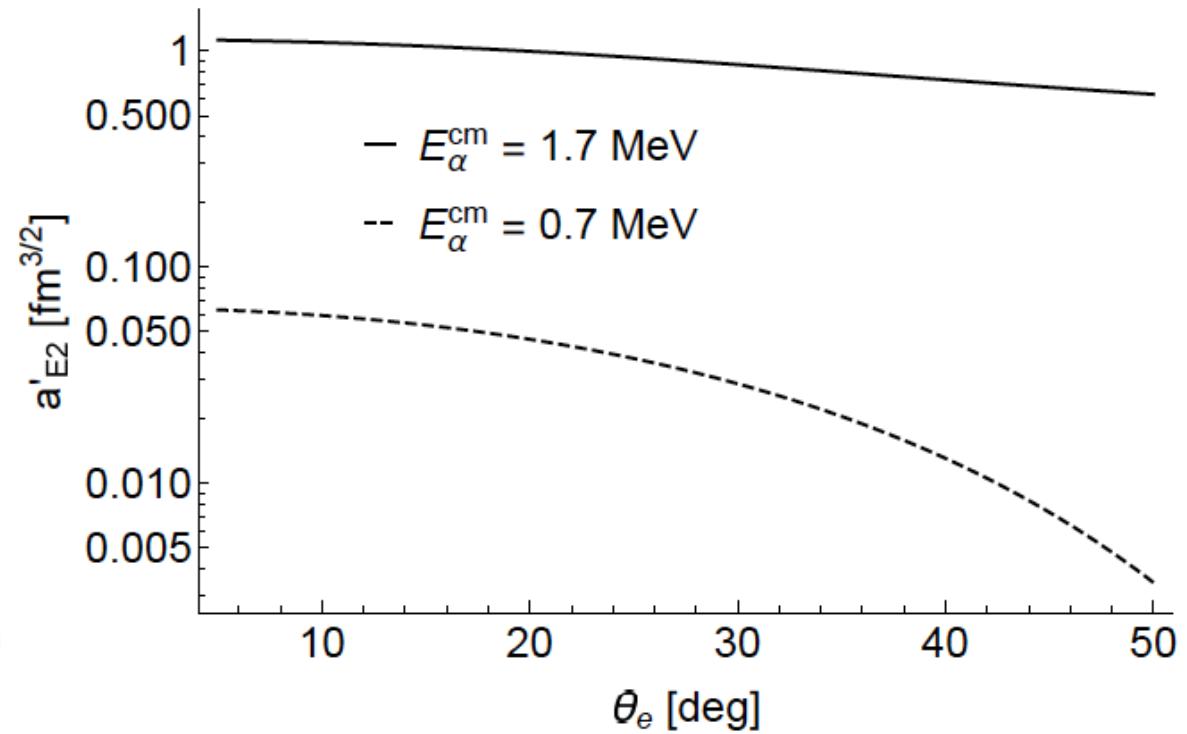
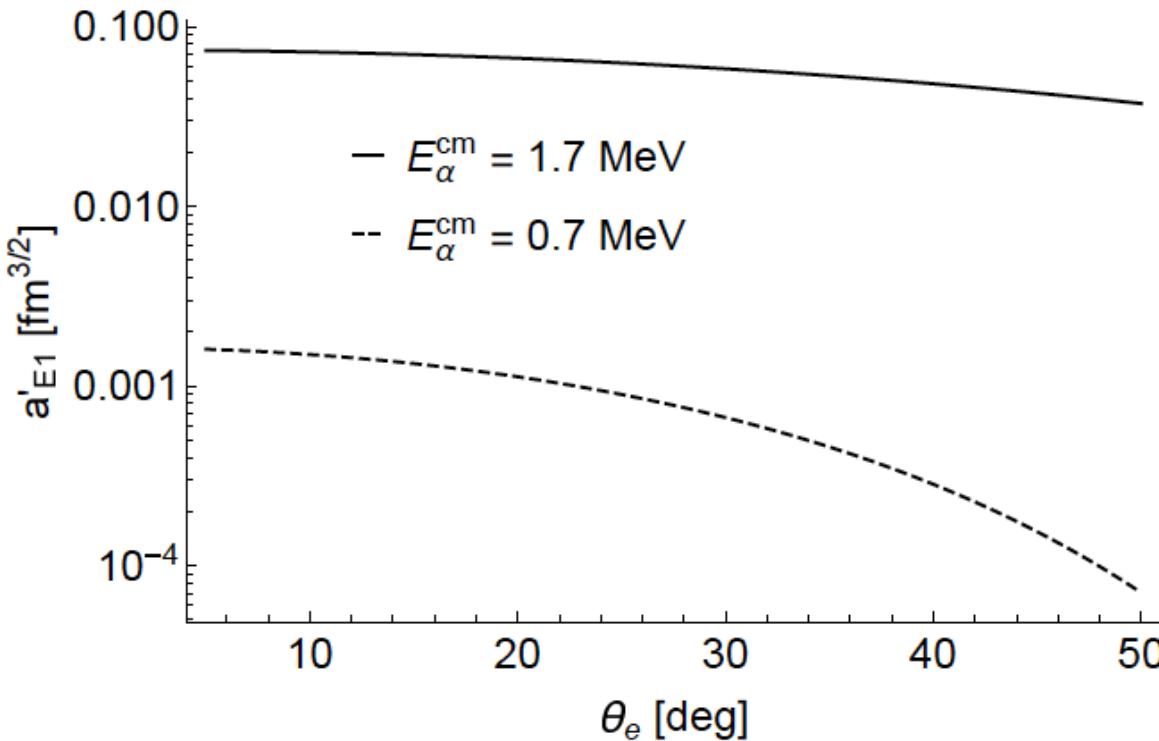
S-factor modeling

- Second order polynomial fit to data $E_{\alpha}^{cm} < 1.7$ MeV



Leading order coefficients

$$a'_{EJ} = \left(\frac{q_0}{\omega} \right)^J \sqrt{\frac{\hbar c \cdot p_\alpha^{cm} \cdot W}{2\alpha \cdot \omega \cdot M_\alpha M_{12} C}} \frac{S_{EJ}(E_\alpha^{cm}) \cdot e^{-2\pi\eta(E_\alpha^{cm})}}{E_\alpha^{cm}} ; \quad J = 1, 2.$$



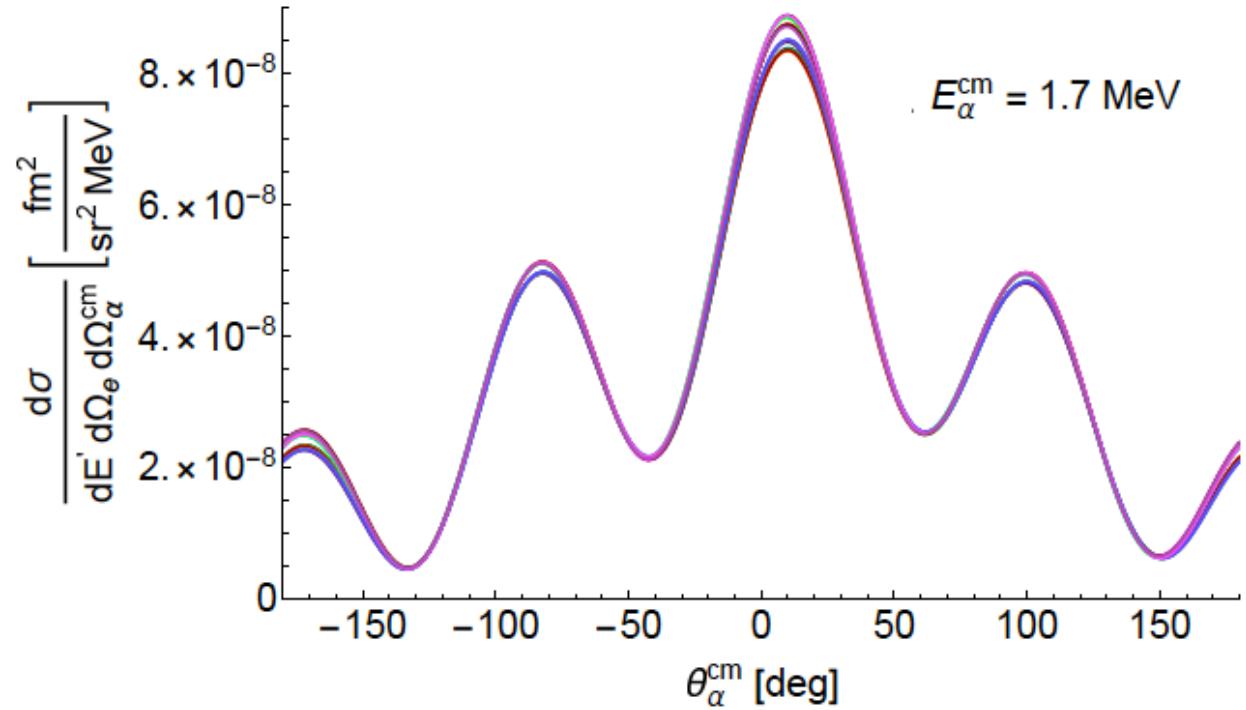
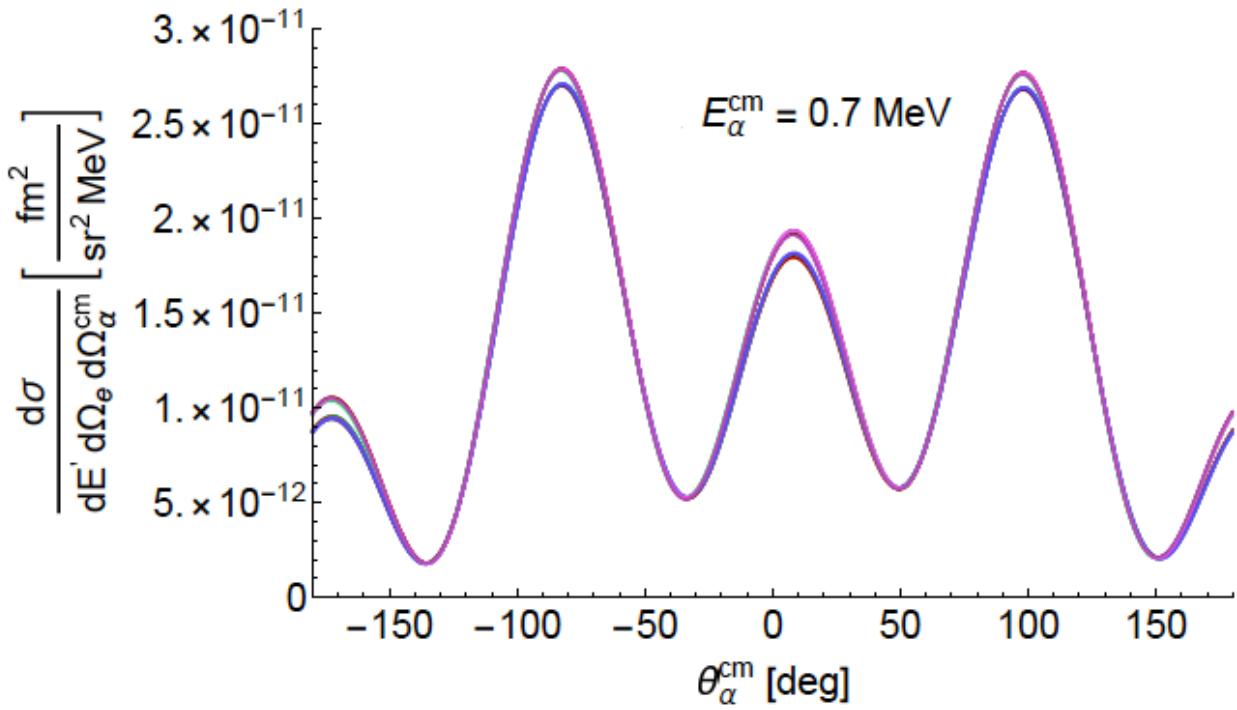
Next-to-leading order coefficients

- No knowledge about next to leading order coefficients $b'_{EJ,CJ}$ with $J = 1, 2$
→ Assuming $b'_{EJ,CJ} \approx 1$ and “+” sign
- No knowledge about C0 multipole and $b'_{C0} \cdot a'_{C0}$
→ Assuming $b'_{C0} \approx 1$ and “+” sign, **Case A** $a'_{C0} = a'_{E2}$ and **Case B** $a'_{C0} = 0.5a'_{E2}$
- For $E_\alpha^{cm} < 1.7$ MeV only Coulomb phase contributes:

$$\delta_{Cl} - \delta_{C0} = \delta_{El} - \delta_{E0} = \sum_{n=1}^l \arctan \frac{\eta}{l}$$

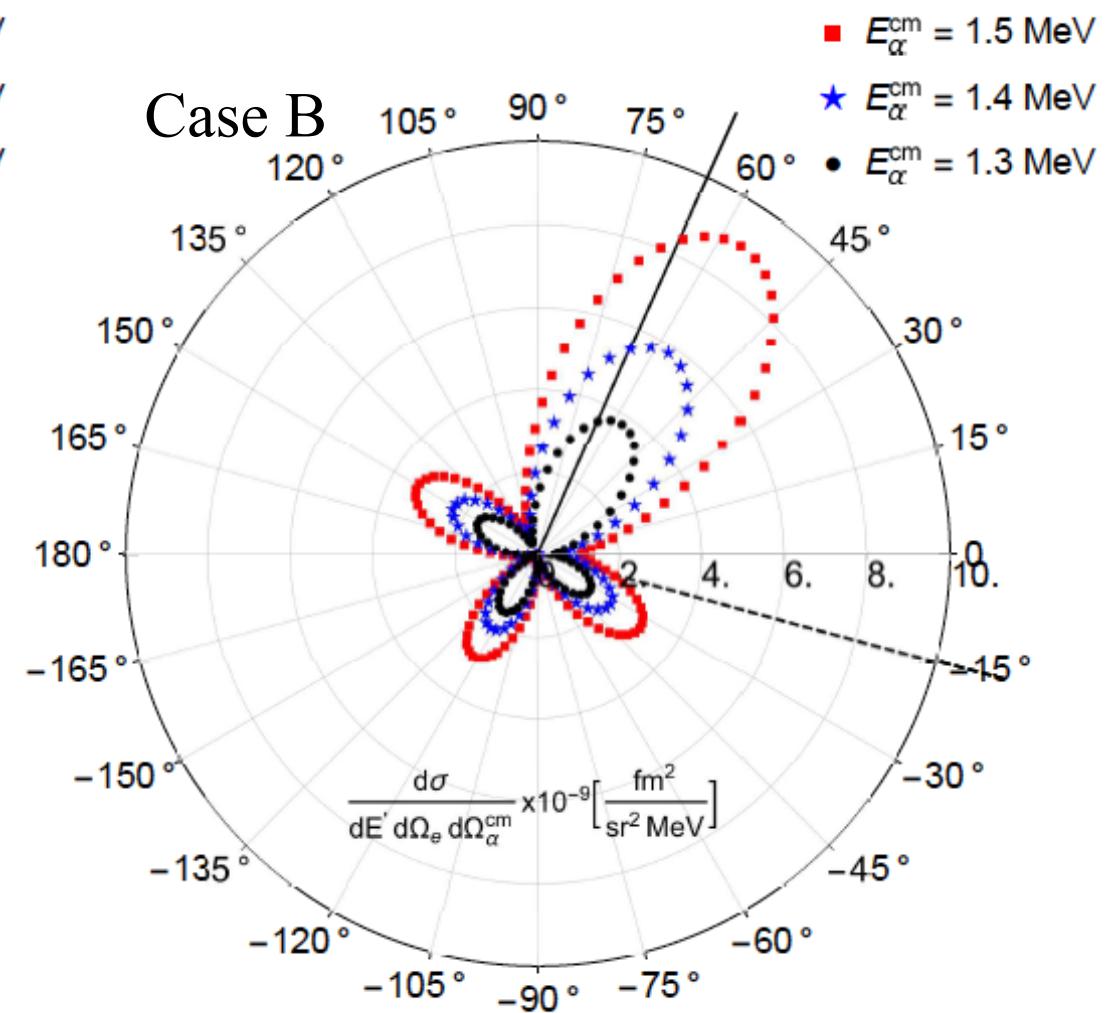
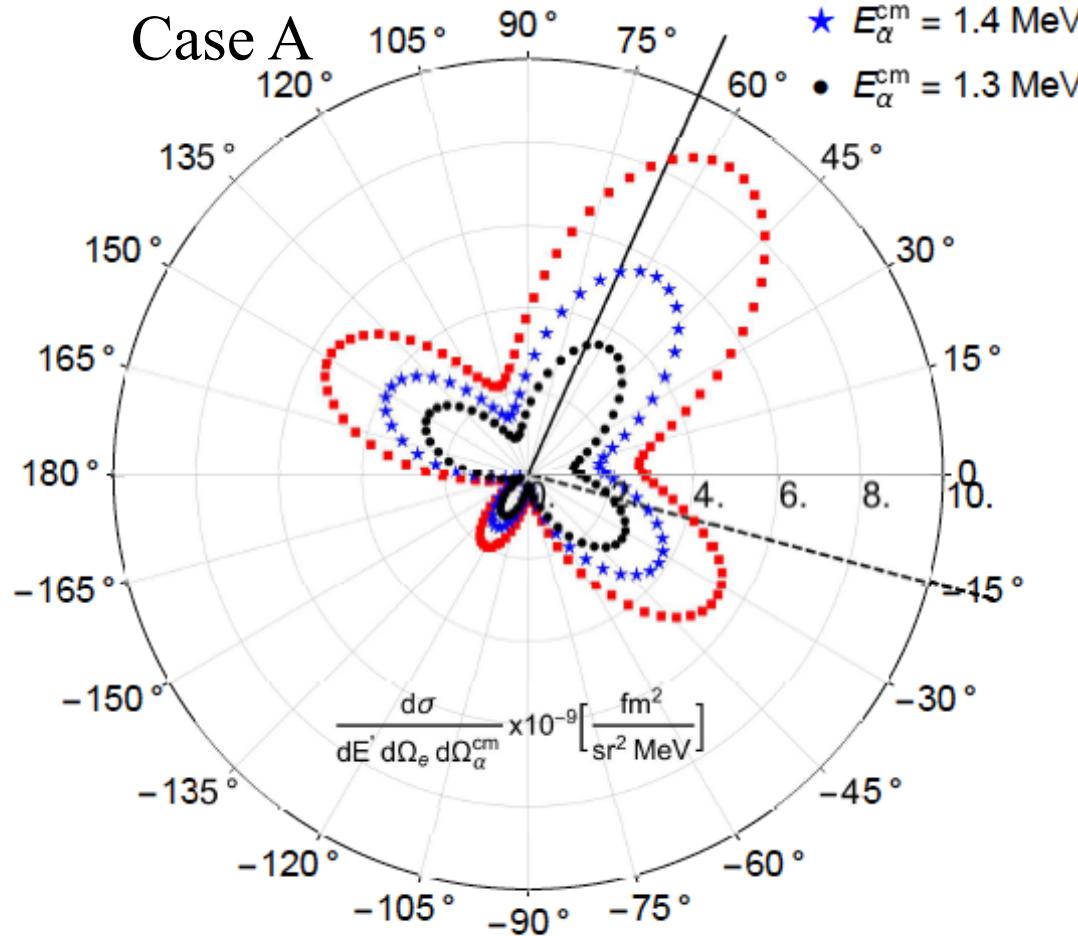
Differential cross section: $^{16}\text{O}(\text{e},\text{e}'\alpha)^{12}\text{C}$

- $E_e = 114 \text{ MeV}$, $\theta_e = 15^\circ$
- Minor differences corresponding to the local maxima from $+b'_{C2}$ and $-b'_{C2}$



Differential cross section: $^{16}\text{O}(\text{e},\text{e}'\alpha)^{12}\text{C}$

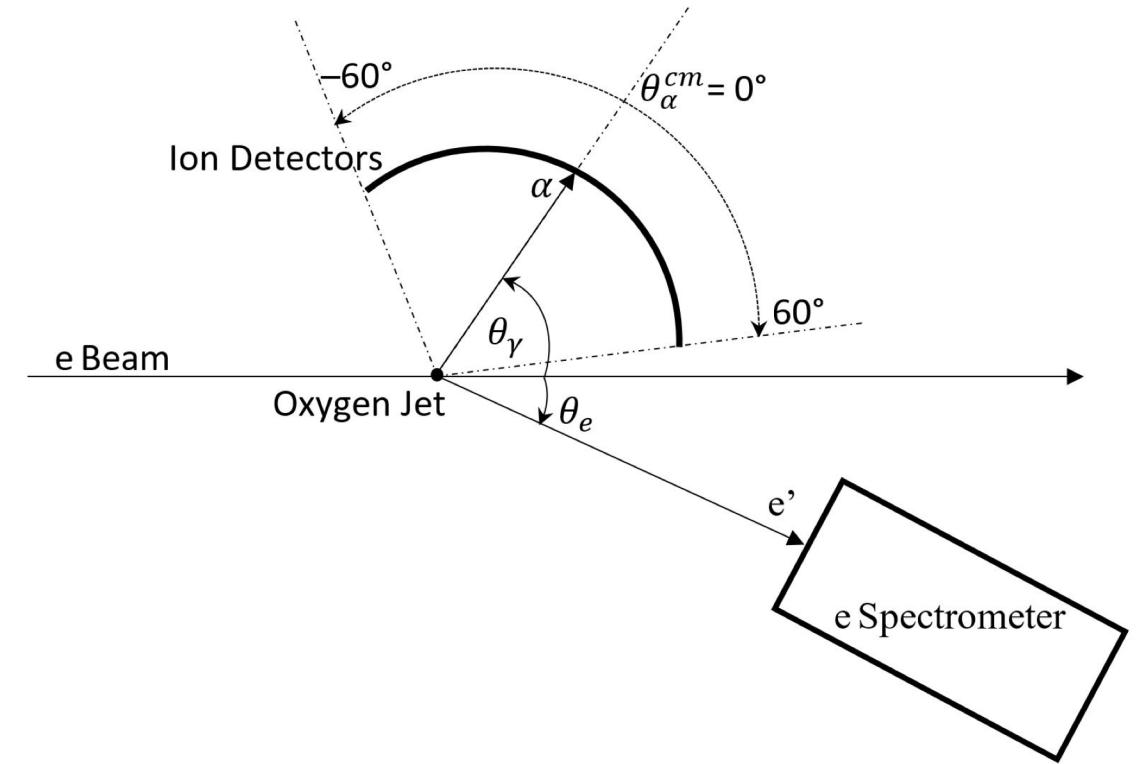
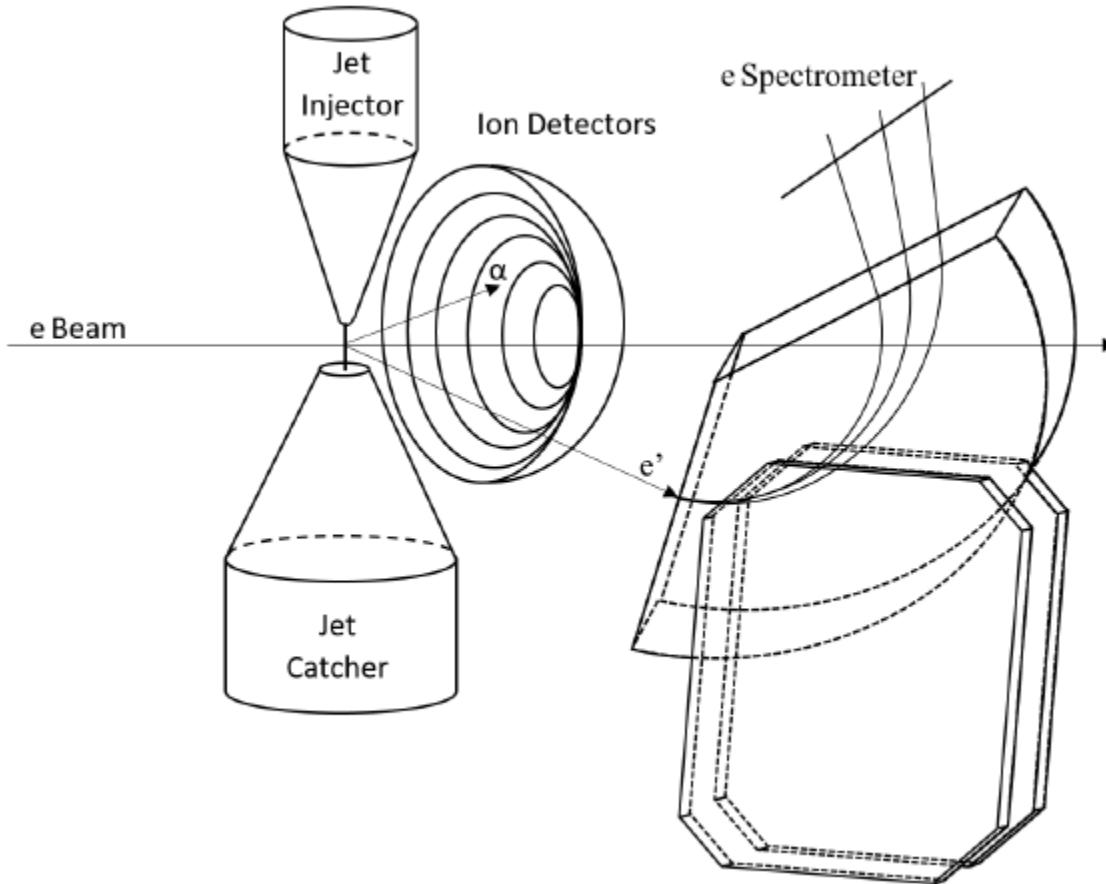
- $E_e = 114 \text{ MeV}, \theta_e = 15^\circ$



Previous experiments and proposals

- G. De Meyer et al., Phys. Lett. B 513 (2001): α -knockout experiment at NIKHEF; $3\mu\text{A}$ e^- beam at 639 MeV and 615 MeV; E_α : from 20 to 35 MeV; target: CO_2 at 1.6 bar and 300 K $\rightarrow 1.46 \cdot 10^{34} (\text{cm}^{-2}\text{s}^{-1})$
- E. Tsentalovich et al., (2000), MIT-Bates PAC proposal 00-01: 100 mA e^- beam at 400 MeV in a storage ring, cluster jet target $2 \cdot 10^{16} \text{ at/cm}^2 \rightarrow 10^{34} (\text{cm}^{-2}\text{s}^{-1})$; Blast detector for the scattered e^- and silicon detectors for α
- T. W. Donnelly, "Electron scattering and the nuclear many-body problem", in The Nuclear Many-Body Problem 2001, edited by W. Nazarewicz and D. Vretenar (Springer Netherlands, Dordrecht, 2002) pp. 19
- S. Lunkenheimer, "Studies of the nucleosynthesis $^{12}\text{C}(\alpha,\gamma)^{16}\text{O}$ in inverse kinematics for the MAGIX experiment at MESA", (2017), 650. WE-Heraeus-Seminar.

Schematic layout of the proposed experiment



Parameters for the rate calculation

Oxygen Target	Thickness Density	5×10^{18} atoms/cm ² 6.65×10^{-4} g/cm ³
Electron Beam	Current Energies	40 mA 78, 114, 150 MeV
Electron arm	In-plane acceptance	$\pm 2.08^\circ$
	Out-of-plane acceptance	$\pm 4.16^\circ$
	Solid angle acceptance	10.5 msr
α -particle arm	In-plane acceptance	60°
	Out-of-plane acceptance	360°
	Solid angle acceptance	3.14 sr
Luminosity		1.25×10^{36} cm ⁻² s ⁻¹
Integrated Luminosity (100 days)		1.08×10^7 pb ⁻¹
Central electron scattering angles		15°, 25°, 35°
E_α^{cm} -range of interest		$0.7 \leq E_\alpha^{cm} \leq 1.7$ MeV

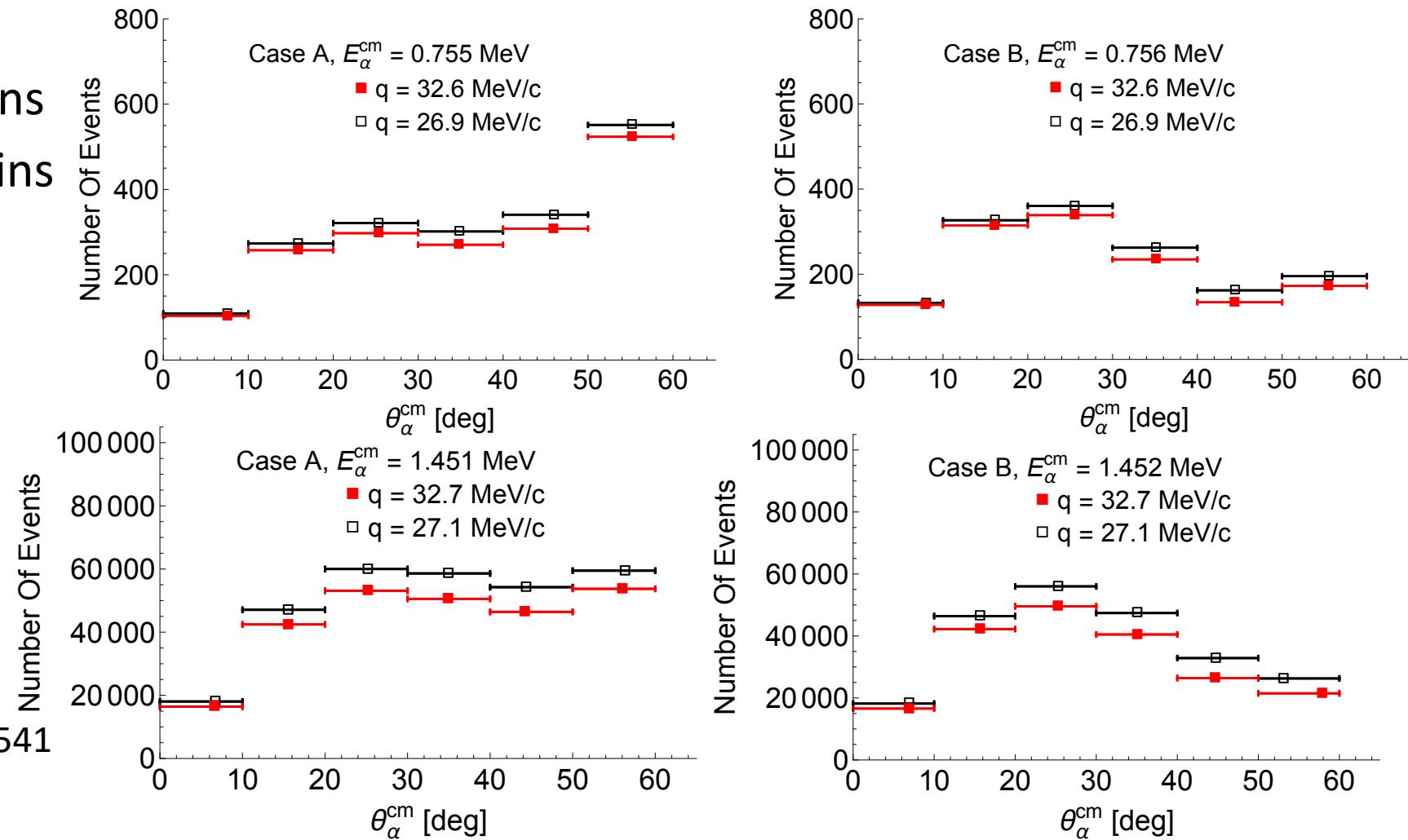
Number of events after 100 days

- Events were sorted in:
 - four 1.91 MeV wide q -bins
 - ten 100 keV wide E_α^{cm} -bins
 - six 10° wide $\theta_\alpha^{\text{cm}}$ -bins
- $E_e = 114 \text{ MeV}$, $\theta_e = 15^\circ$,
Case A and Case B

- Now we can compute statistical uncertainties**

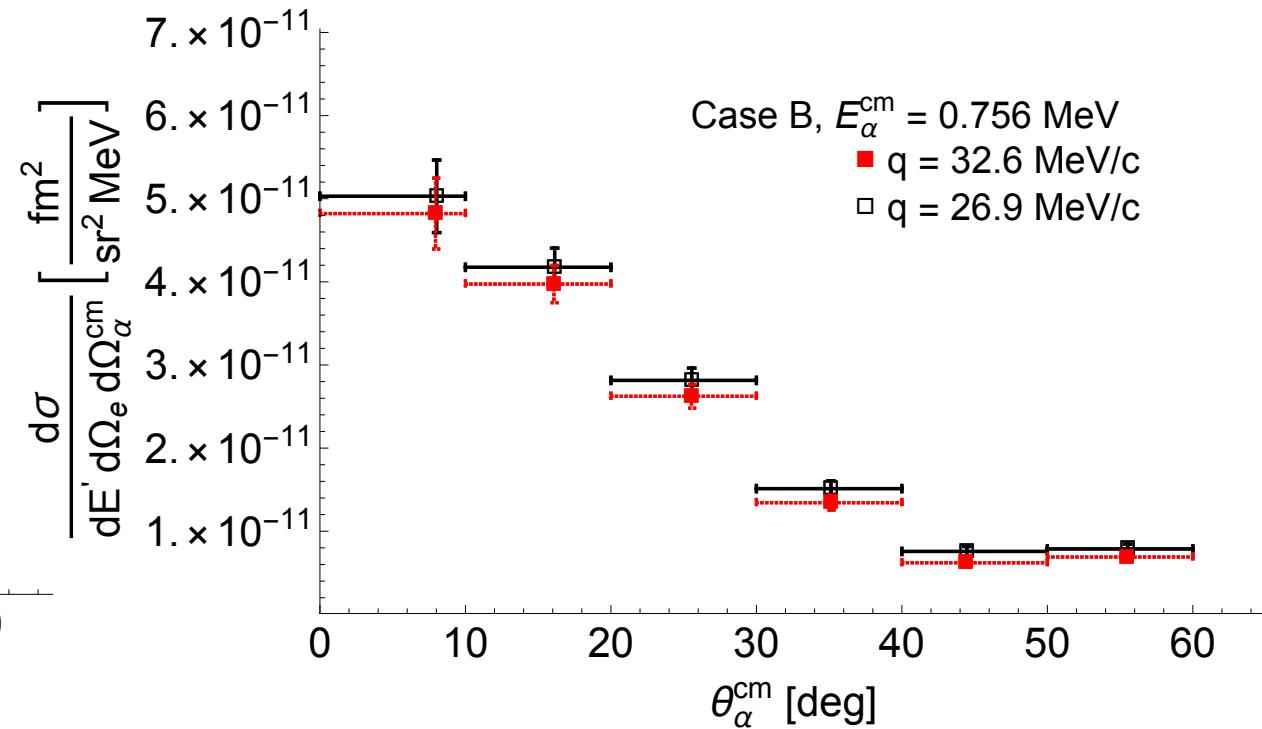
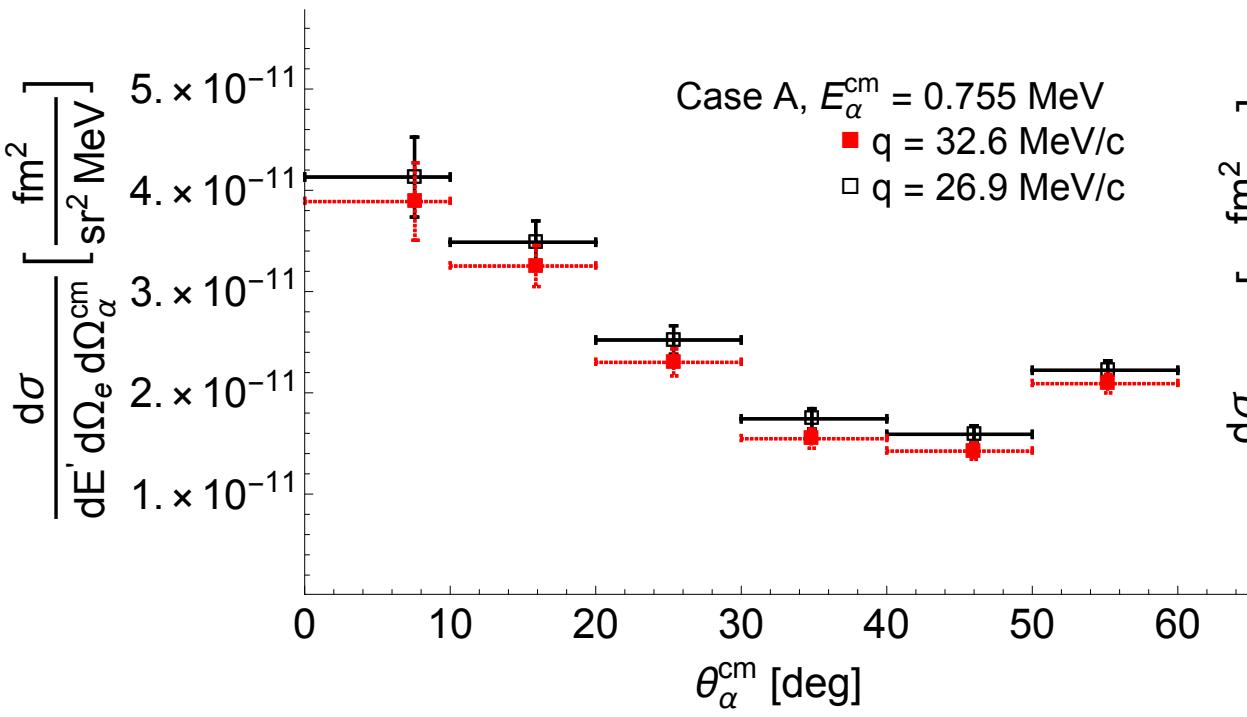
- Horizontal placement of data points according to:

G. D. Lafferty and T. R. Wyatt, Nucl.
Instrum. Methods Phys. Res. A 355, 541
(1995).



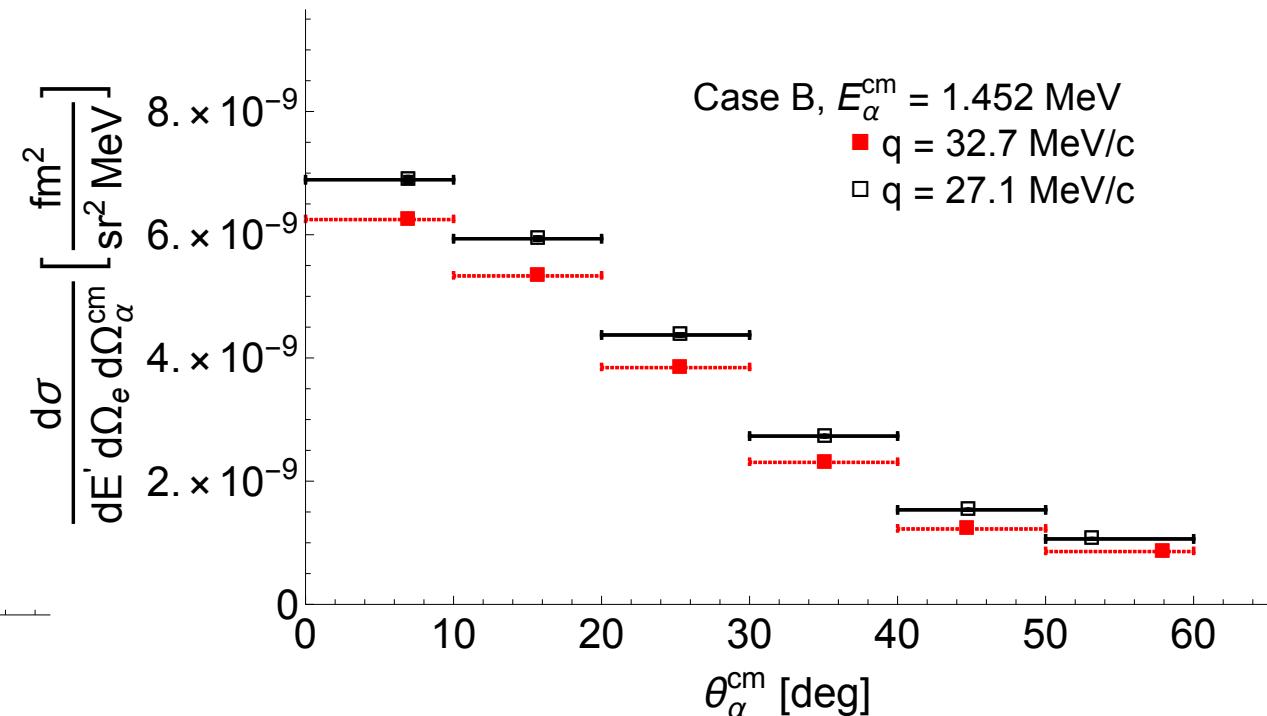
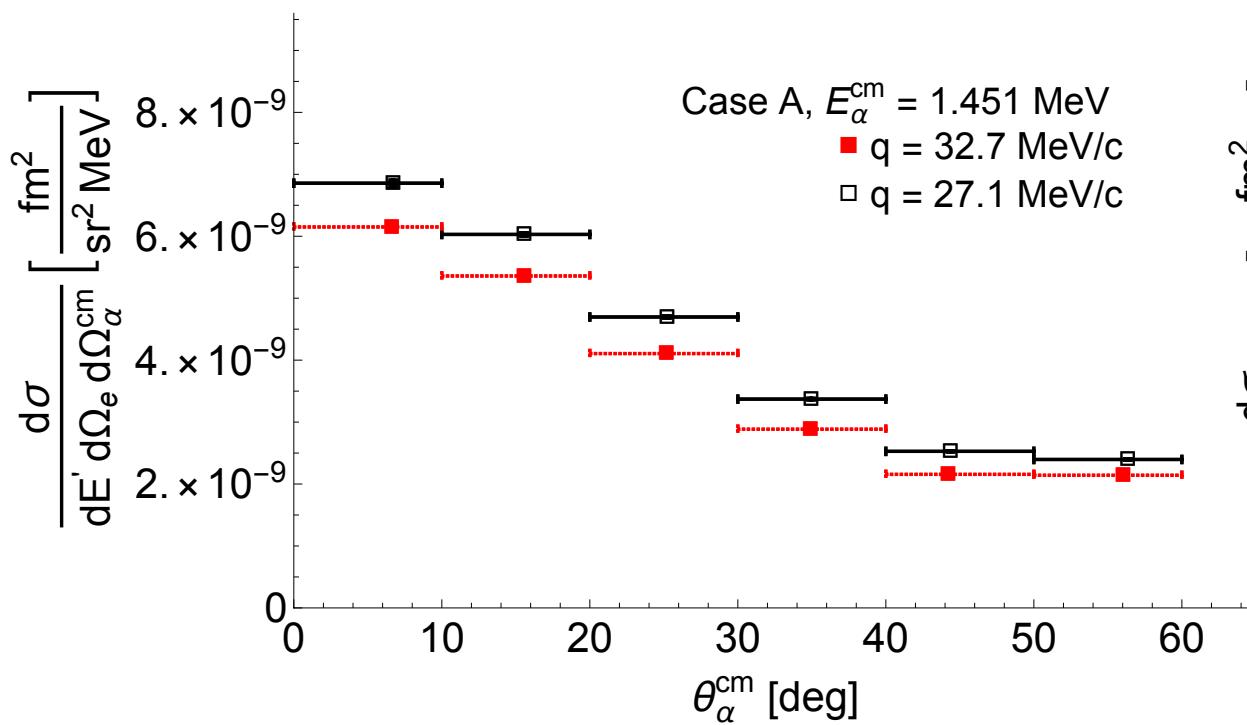
Differential cross section after 100 days

- $E_e = 114 \text{ MeV}$, $\theta_e = 15^\circ$, Case A and Case B



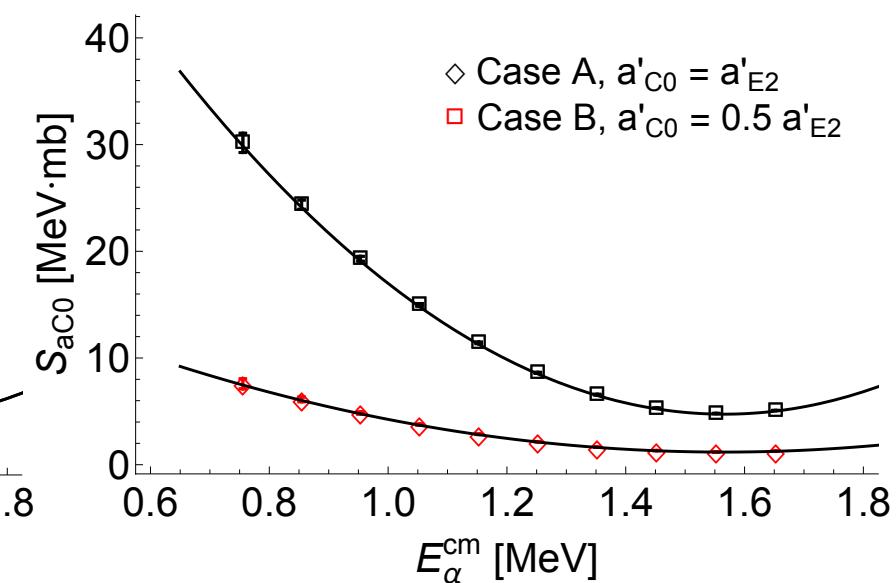
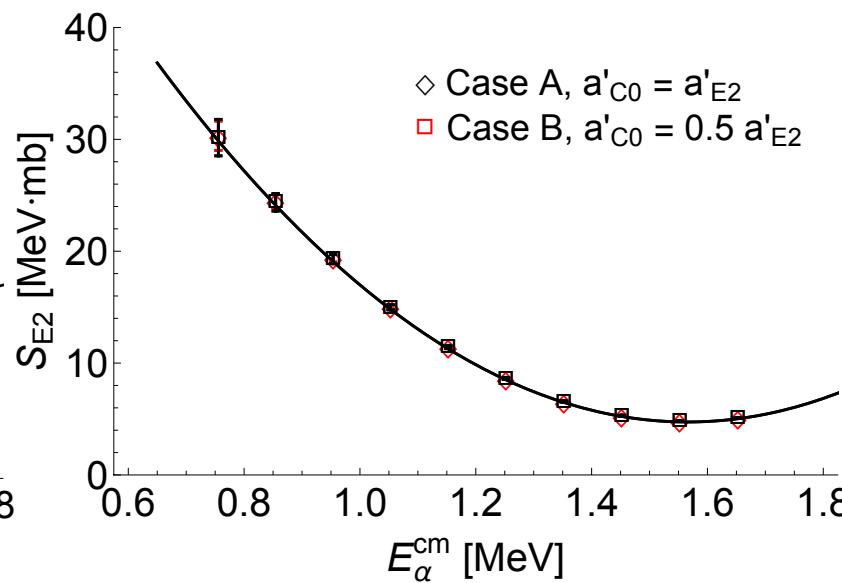
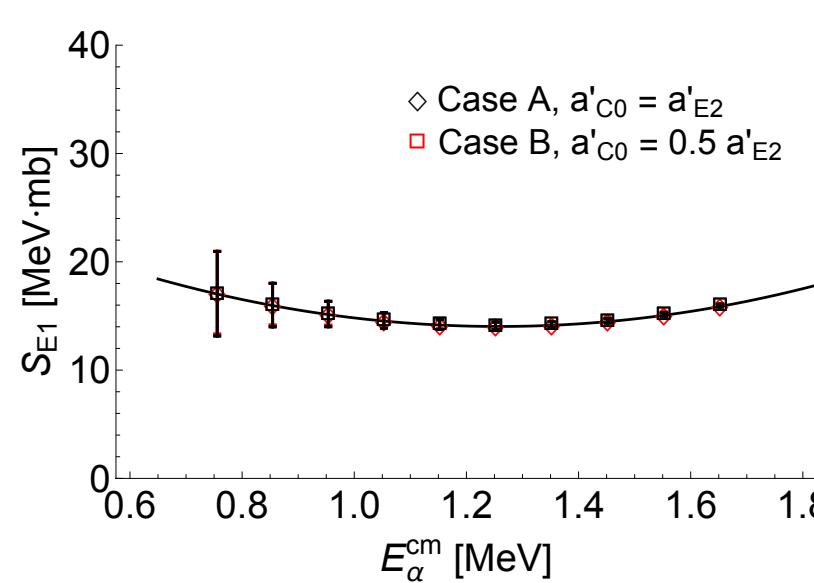
Differential cross section after 100 days

- $E_e = 114 \text{ MeV}$, $\theta_e = 15^\circ$, Case A and Case B



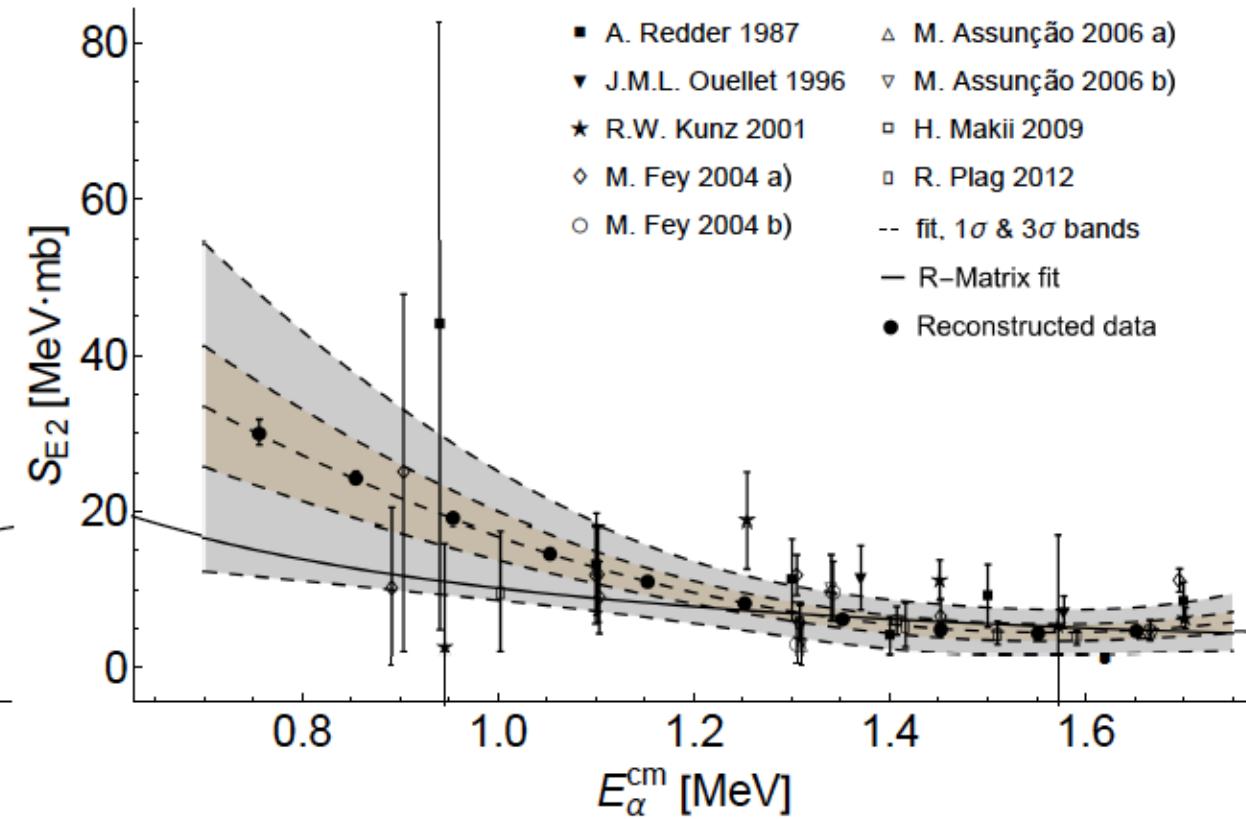
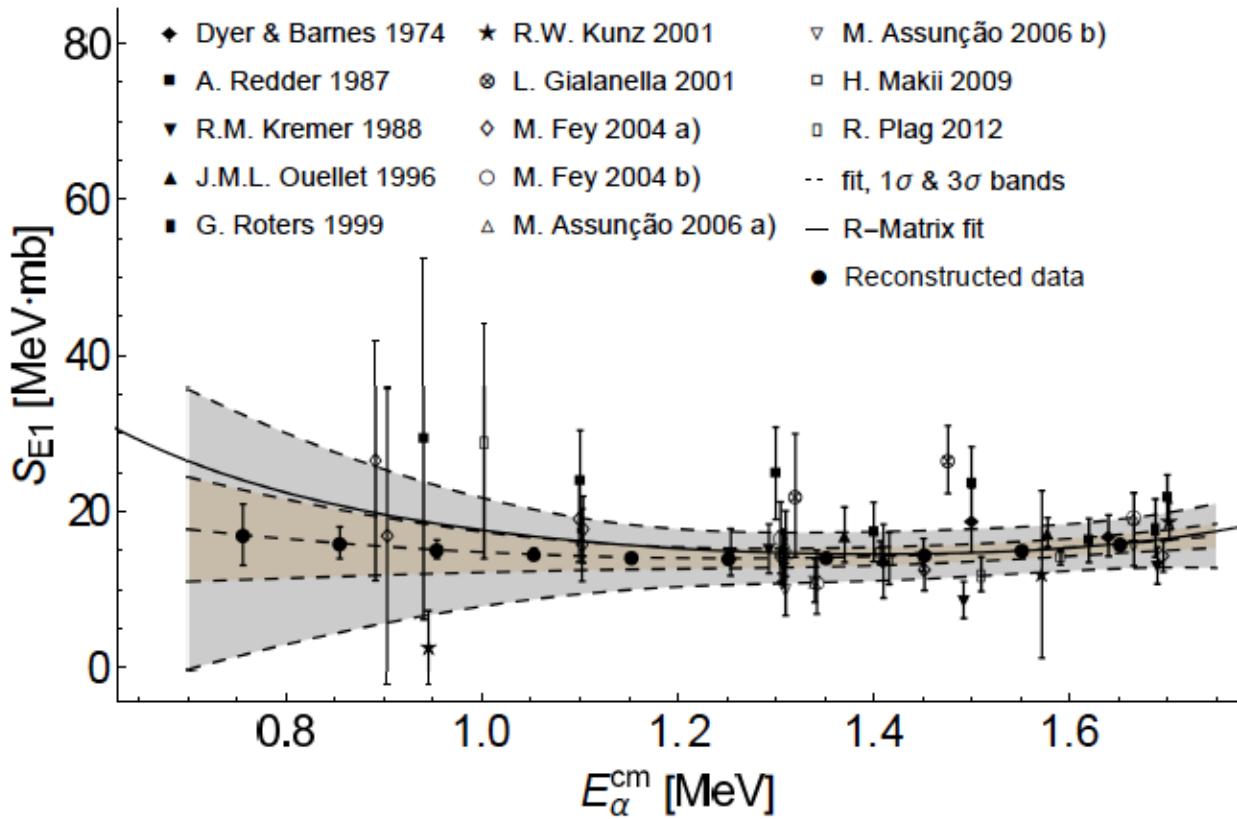
S-factors with projected statistical uncertainties

- $E_e = 114 \text{ MeV}$, $\theta_e = 15^\circ$, Case A and Case B
- Three fitting parameters a'_{E1} , a'_{E2} and $a'_{C0} \rightarrow S_{E1}$, S_{E2} and S_{aC0} non-astrophysical factor



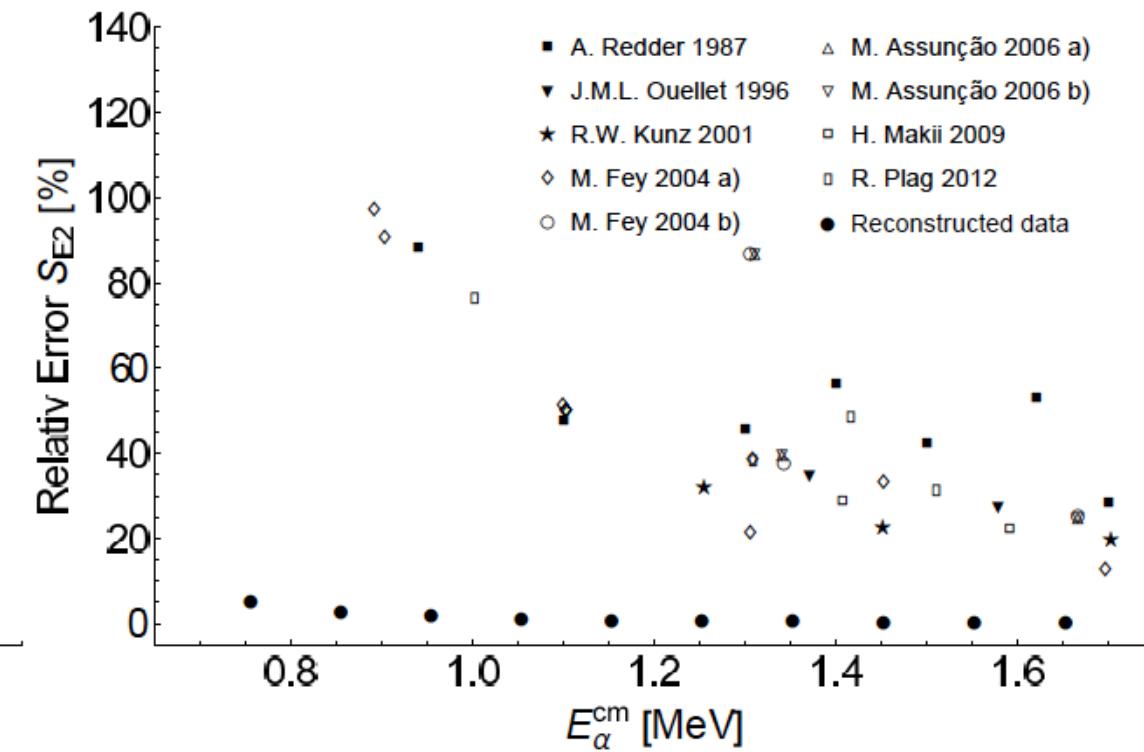
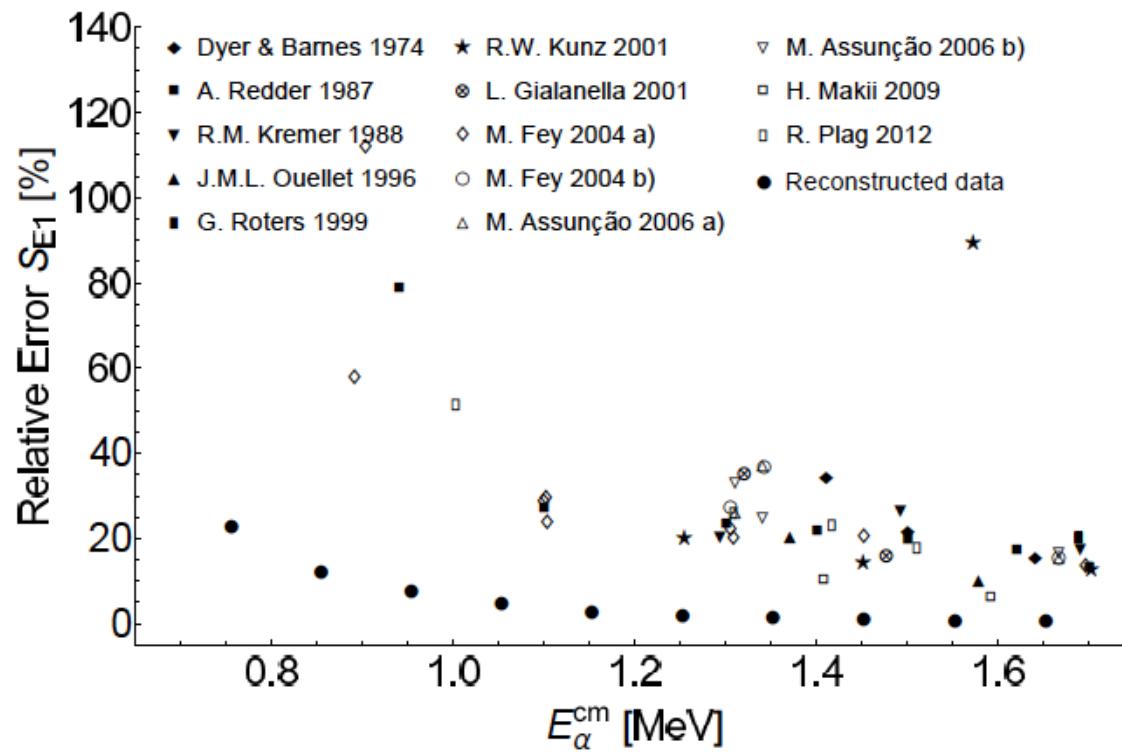
S-factors with projected statistical uncertainties

- $E_e = 114 \text{ MeV}$, $\theta_e = 15^\circ$, Case A



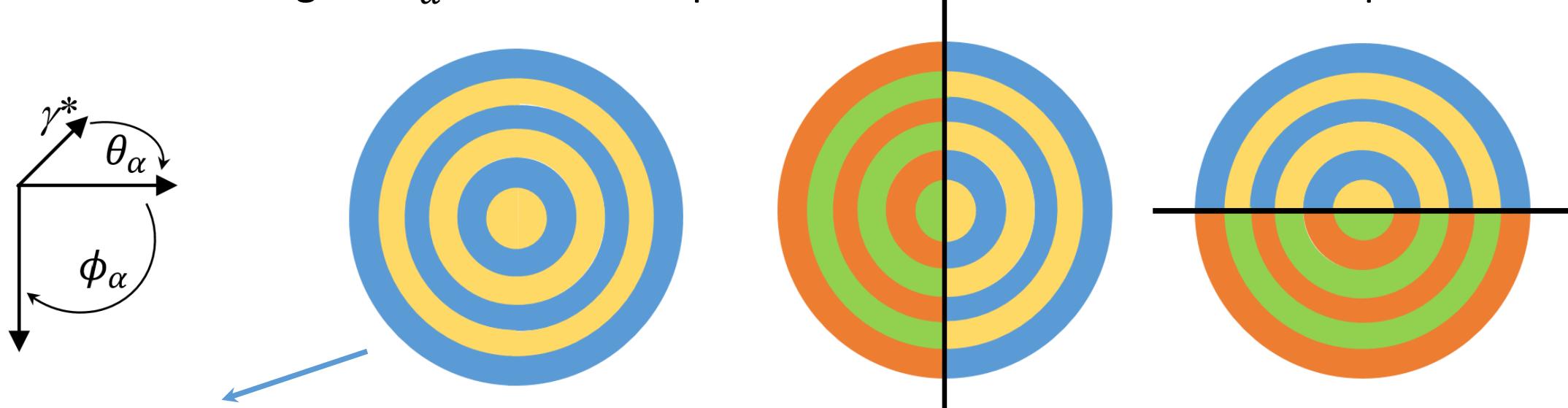
S-factors with projected statistical uncertainties

- $E_e = 114 \text{ MeV}$, $\theta_e = 15^\circ$, Case A
- Compared to most accurate measurements, statistical uncertainties of S_{E1} and S_{E2} are improved at least by factors 5.6 and 23.9, respectively



Outlook

- Shorter run at higher E_α^{cm} to test the particle identification and all assumptions



- Only R_T and R_L contribute to the rate
- Rosenbluth separation of R_T and R_L to extract the S-factors and the phases separately

- $\sigma_{\text{Right}} - \sigma_{\text{Left}} \rightarrow 2R_{TL}$
 - $\sigma_{\text{Right}} + \sigma_{\text{Left}} \rightarrow 2(R_T + R_L)$
 - Form an asymmetry
- $\sigma_{\text{Up}} = \sigma_{\text{Down}}$ \rightarrow test of efficiency and systematics

Conclusion

- Using a simple model, possibilities of new ERL accelerators and the gas-jet target, we studied the rate of $^{16}\text{O}(\text{e},\text{e}'\alpha)^{12}\text{C}$ reaction in range $0.7 < E_\alpha^{cm} < 1.7$ MeV and showed how to determine $^{12}\text{C}(\alpha, \gamma)^{16}\text{O}$ reaction rate with unprecedented statistical precision
- At $E_e = 114$ MeV and electron spectrometer with 10% E'_e acceptance the full range $0. < E_\alpha^{cm} < 10.2$ MeV is accessible
- For more details: **I. Friščić, W. T. Donnelly and R. G. Milner, Phys. Rev. C 100, (2019) 025804; or arXiv:1904.05819**

Work supported by the DOE Office of Nuclear Physics under grant No. DE-FG02-94ER40818.

Backup

The cross section formulas

A. S. Raskin and T. W. Donnelly, Ann. of Phys. 191 (1989)

$$\frac{d\sigma}{dE'_e d\Omega_e d\Omega_\alpha^{cm}} = \frac{M_\alpha M_{12C}}{8\pi^3 W} \frac{p_\alpha^{cm}}{(\hbar c)^3} \sigma_{Mott} (\tilde{\nu}_L R_L + \tilde{\nu}_T R_T + \tilde{\nu}_{LT} R_{LT} + \tilde{\nu}_{TT} R_{TT})$$

$$W = \sqrt{(M_{16O} + \omega)^2 - q^2} \quad E_\alpha^{cm} = W - W_{th}$$

$$\frac{d\sigma}{d\Omega_\alpha^{cm}} \Big|_{(\gamma, \alpha)} = \frac{M_\alpha M_{12C}}{4\pi W} \frac{p_\alpha^{cm}}{\hbar c} \frac{\alpha}{E_\gamma} R_T$$

$$W = \sqrt{M_{16O}(M_{16O} + 2E_\gamma)}$$

$$E_\alpha^{cm} = E_\gamma - 7.162 \text{ MeV}$$

$$\frac{d\sigma}{d\Omega_\gamma^{cm}} \Big|_{(\alpha, \gamma)} = \frac{M_\alpha M_{12C}}{2\pi W} \frac{p_\alpha^{cm}}{\hbar c} \frac{\alpha}{E_\gamma} R_T$$

$$W = \sqrt{M_\alpha^2 + M_{12C}^2 + 2M_{12C}E_\alpha^{lab}}$$

$$E_\alpha^{cm} = \frac{M_{12C}}{M_{12C} + M_\alpha} E_\alpha^{lab}$$

Differential Cross Section: $^{16}\text{O}(\text{e},\text{e}'\alpha)^{12}\text{C}$

A. S. Raskin and T. W. Donnelly, Ann. of Phys. 191 (1989)

$$\frac{d\sigma}{dE'_e d\Omega_e d\Omega_\alpha^{cm}} = \frac{M_\alpha M_{12C}}{8\pi^3 W} \frac{p_\alpha^{cm}}{(\hbar c)^3} \sigma_{Mott} (\tilde{v}_L R_L + \tilde{v}_T R_T + \tilde{v}_{LT} R_{LT} + \tilde{v}_{TT} R_{TT})$$

$$\rho \equiv |Q^2/q^2| = 1 - (\omega/q)^2$$

$$v_L = \rho^2$$

$$\tilde{v}_L = (W/M_{^{16}\text{O}})^2 v_L$$

Total invariant mass:

$$W = \sqrt{(M_{^{16}\text{O}} + \omega)^2 - q^2}$$

$$v_T = \frac{1}{2}\rho + \tan^2 \theta_e / 2$$

$$\tilde{v}_T = v_T$$

$$E_\alpha^{cm} = W - W_{th}$$

$$v_{TL} = -\frac{1}{\sqrt{2}}\rho \sqrt{\rho + \tan^2 \theta_e / 2}$$

$$\tilde{v}_{TL} = (W/M_{^{16}\text{O}}) v_T$$

$$v_{TT} = -\frac{1}{2}\rho$$

$$\tilde{v}_{TT} = v_{TT}$$