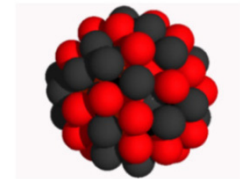


Research opportunities for Nuclear Photonics with an ERL-based hard X-ray source



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Norbert Pietralla, TU Darmstadt

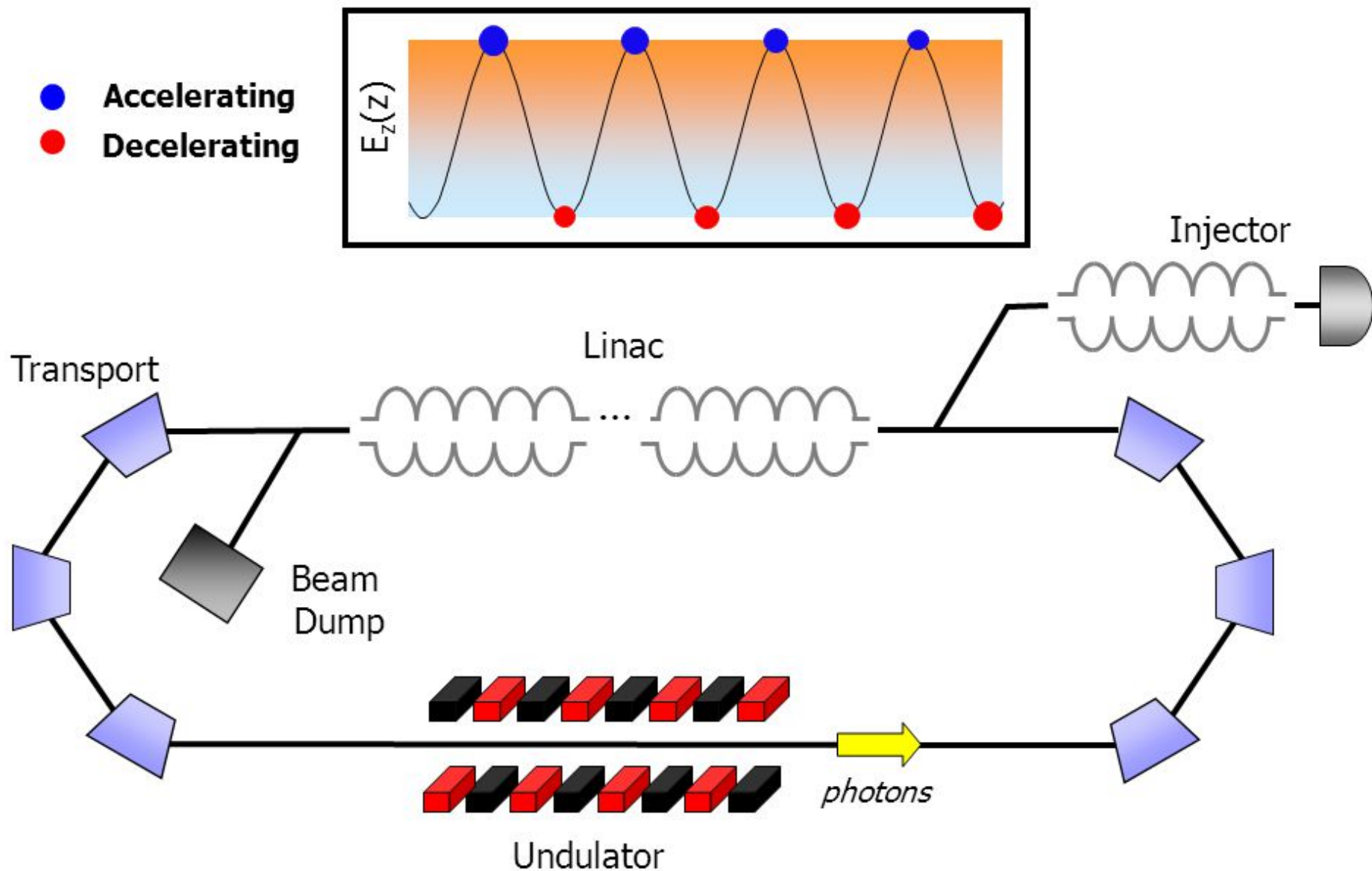


Photonuclear reactions

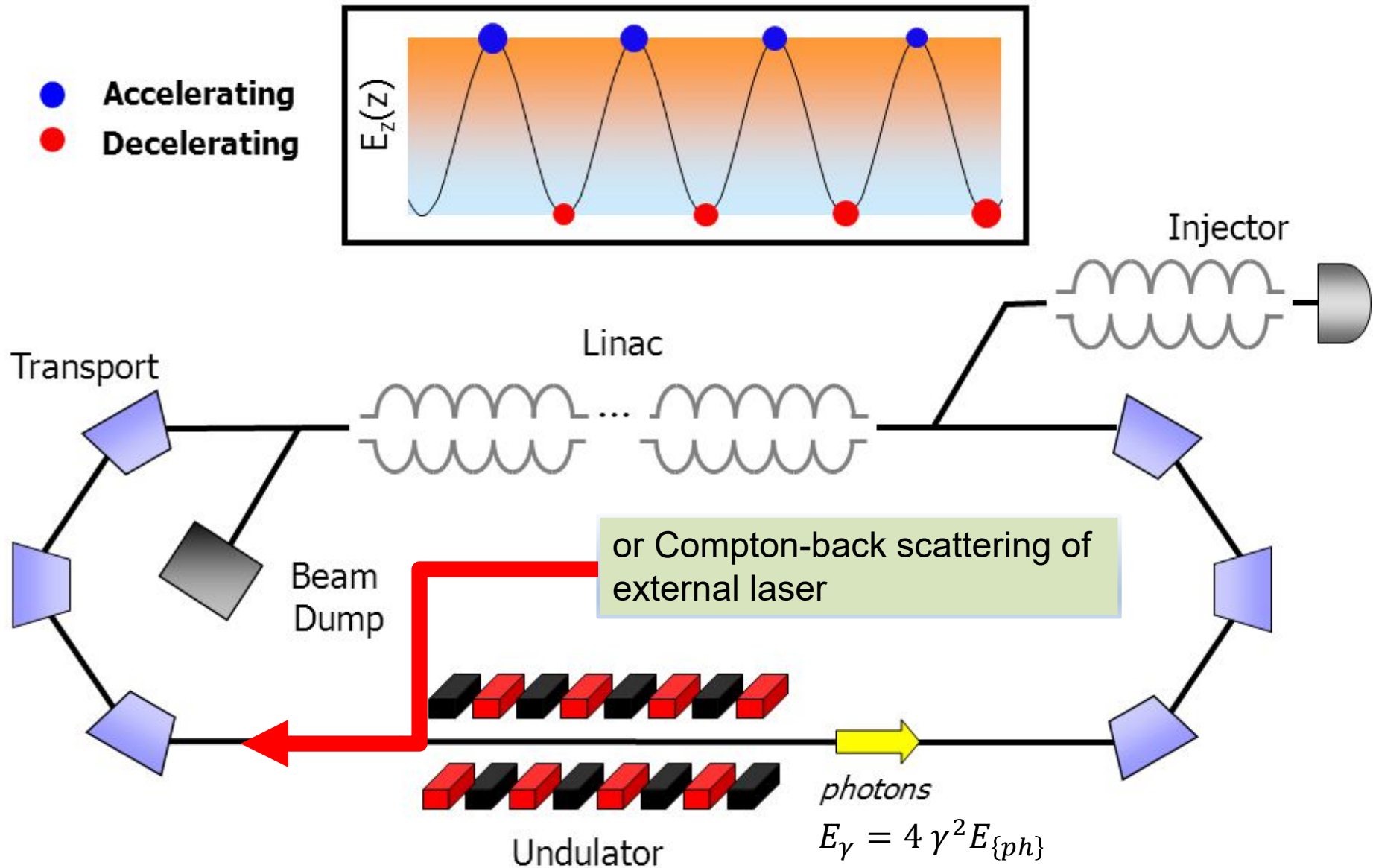


Conference photo
Nuclear Photonics 2018
Brasov, Romania

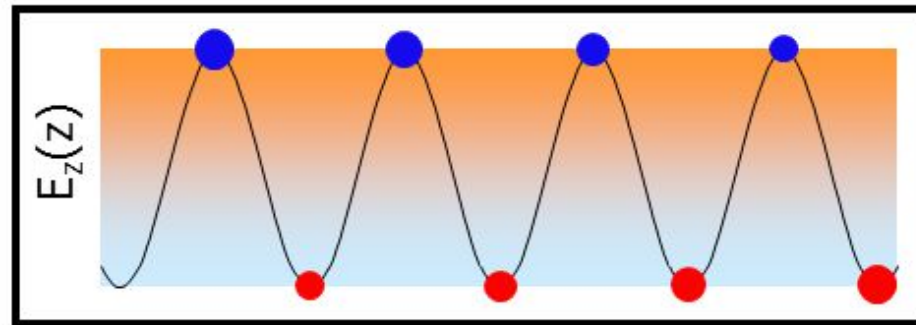
Generic ERL-based Light Source



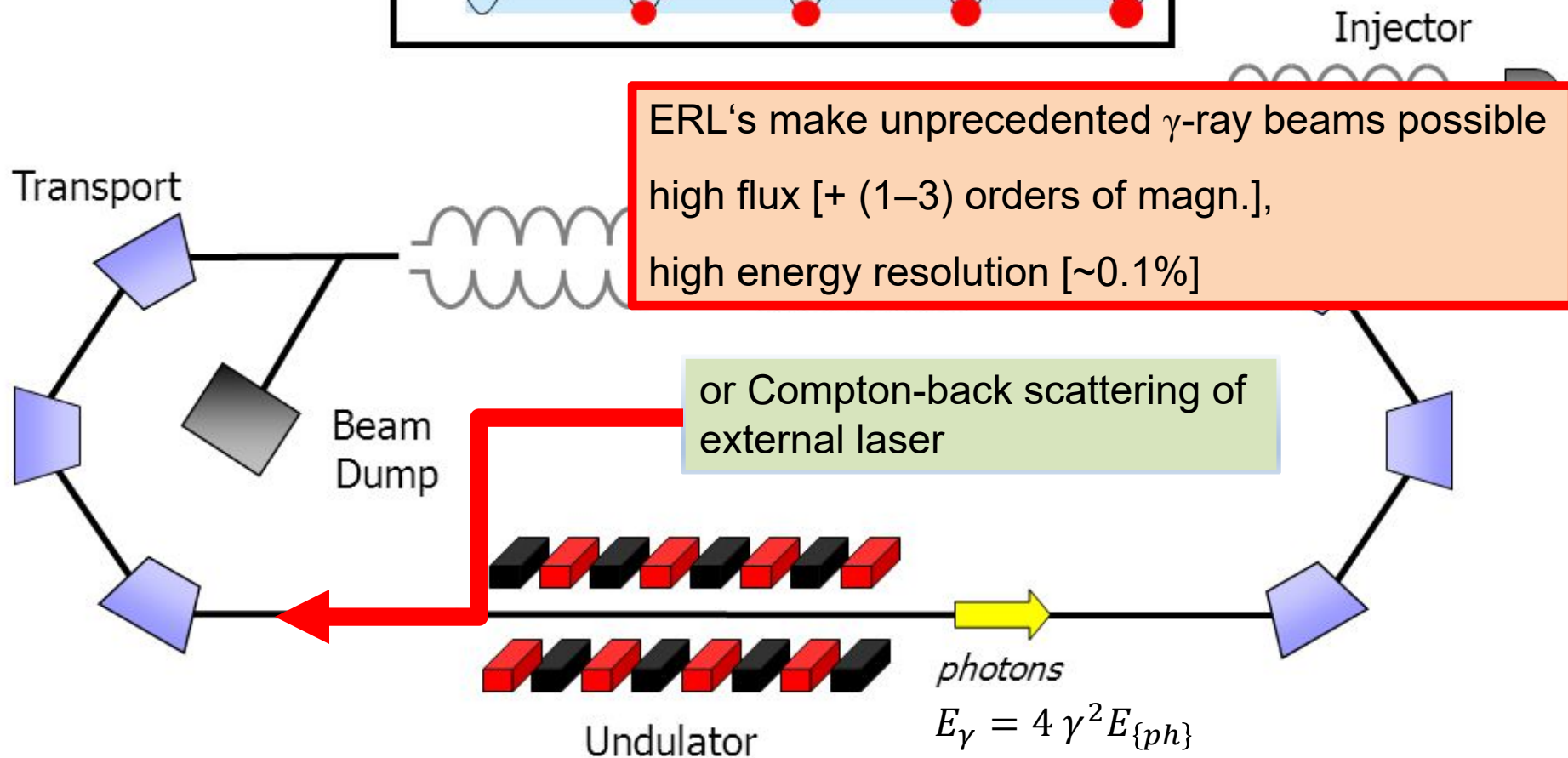
Generic ERL-based Light Source



Generic ERL-based Light Source

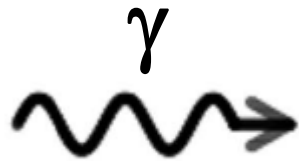


- Accelerating
- Decelerating

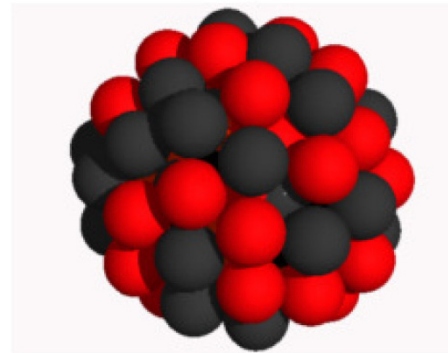


- Photonuclear reactions with MeV-range γ rays
- Methodological developments for nuclear photonics
- Motivation from astrophysics and particle physics
- Examples for recent achievements
- Concluding remarks

Photonuclear Reactions



0.1 – 100 MeV

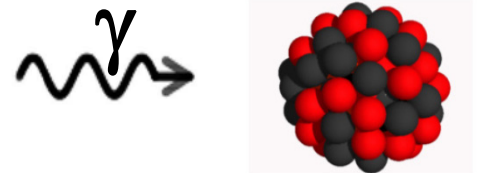


What happens?

Nuclear Physics with photon beams

Pure EM-interaction

- (nuclear-)model independent
- “small” cross sections, intense beams



Minimum projectile mass

- min. angular momentum transfer,
spin-selective: low-spin modes [E1,M1,E2,(E3?)]

Polarisation

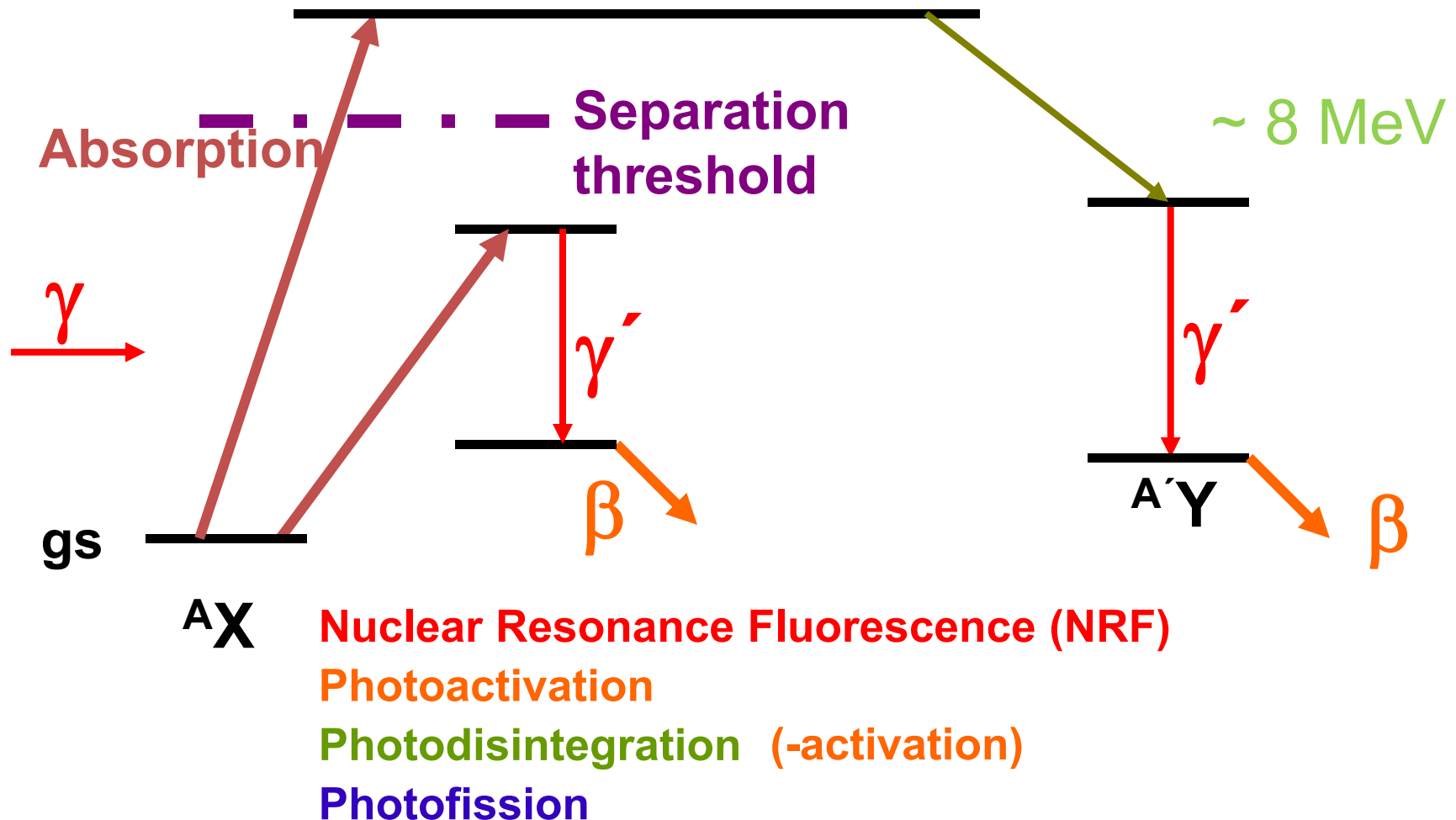
- “Parity Physics”, channel selectivity

Narrow Bandwidth (at 4th generation LCB source)

- Explore specific excitation energy
“Selective Manipulation of Nuclear States”:



Photonuclear Reactions

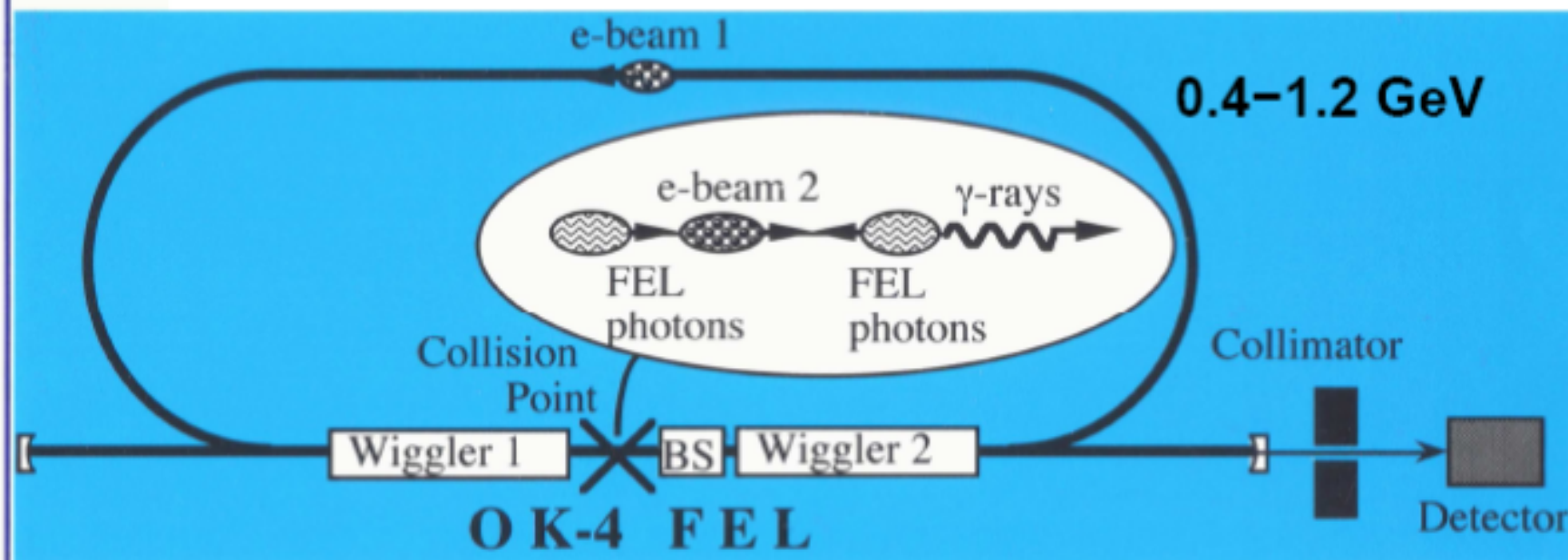


High Intensity γ -Ray Source (HlgS)



H.R.Weller, V.N.Litvinenko
Duke University, Durham, NC, U.S.A.

Compton Backscattering of Intra-cavity Laser Light



2 – 60 MeV

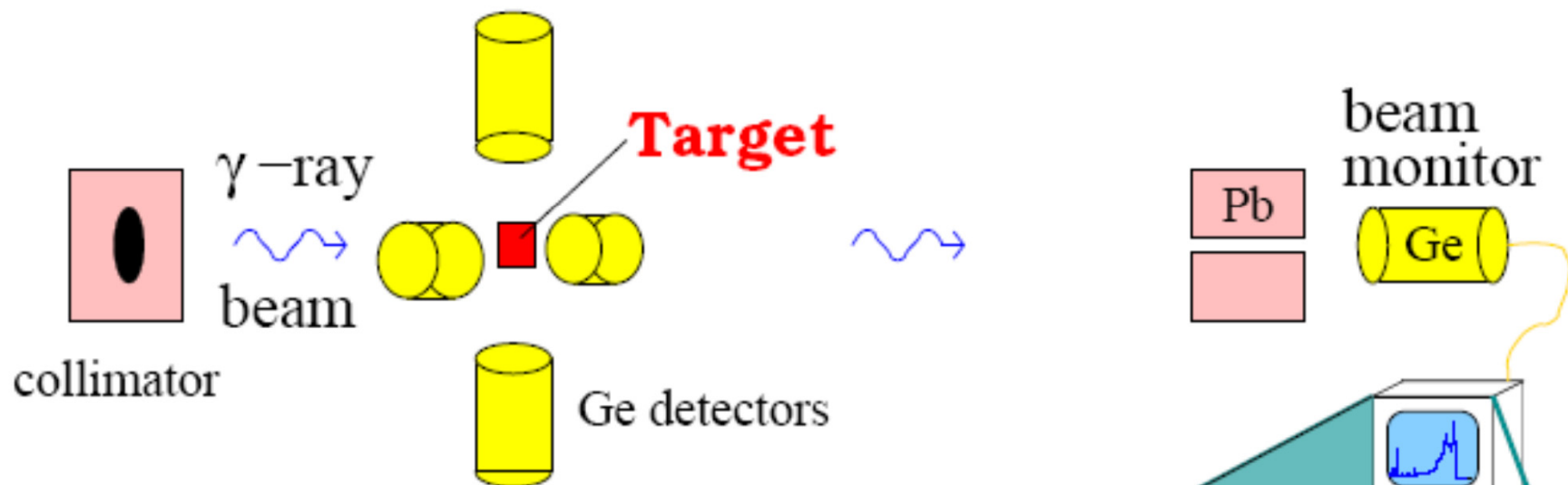
1.7 – 6.4 eV

~ 1000

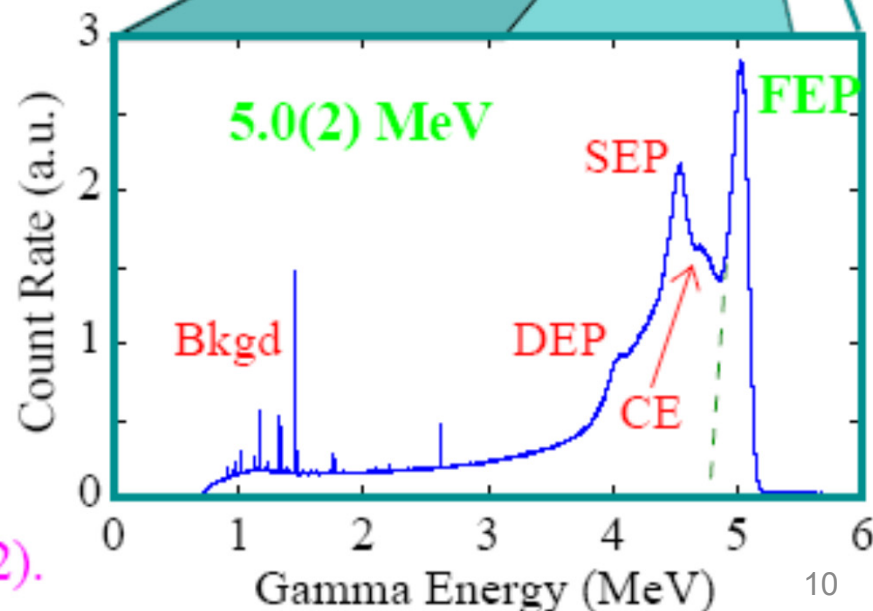
$$E_\gamma = \frac{4\gamma^2 E_{ph}}{(1 + r + \gamma^2 \theta^2)}; \quad r = \frac{4\gamma E_{ph}}{mc^2}; \quad E_{ph} = \frac{2\gamma^2 hc}{\lambda_w (1 + K_w^2 / 2)}; \quad \gamma = \frac{E_e}{mc^2};$$

nearly monochromatic, tunable, completely polarized

Looking at the HIGS Gamma-Ray Beam

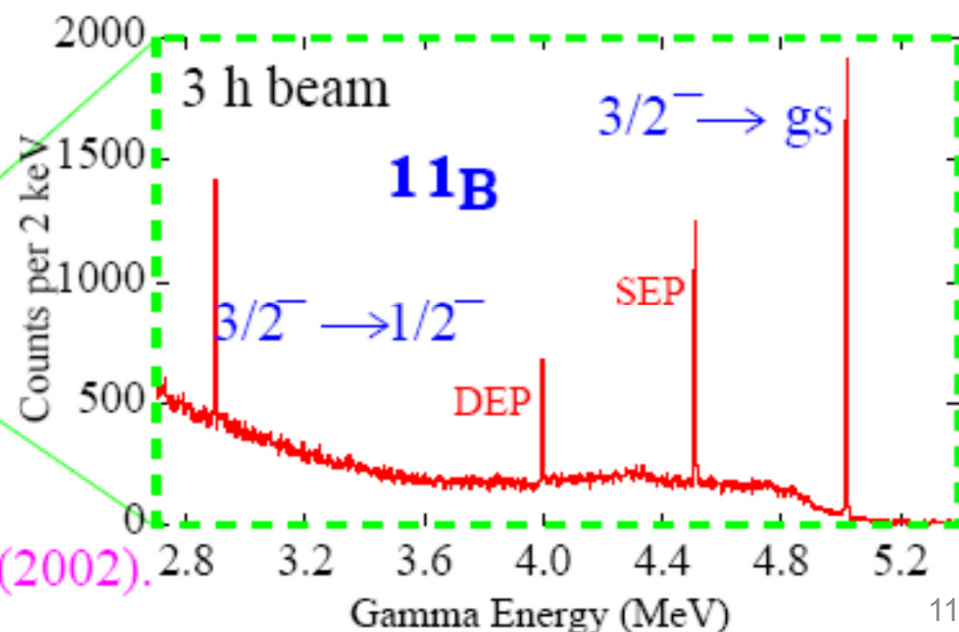
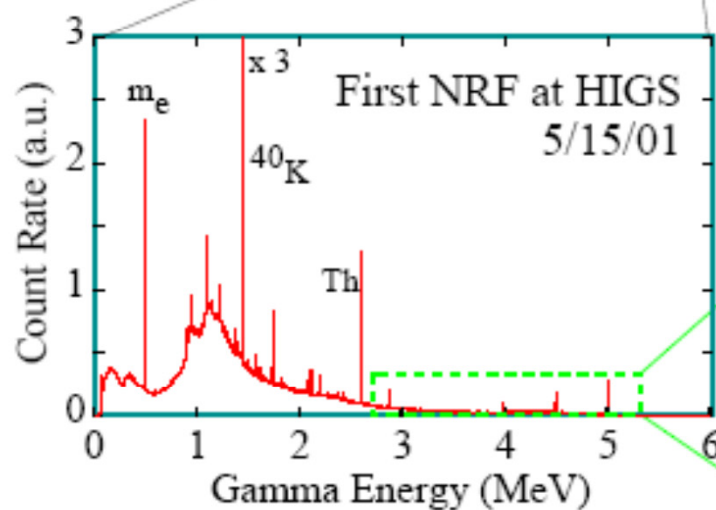
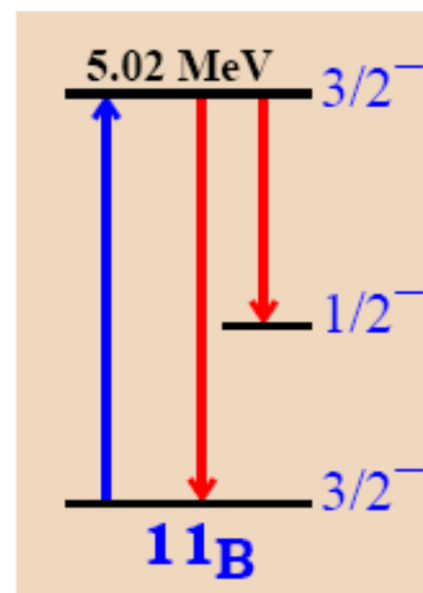
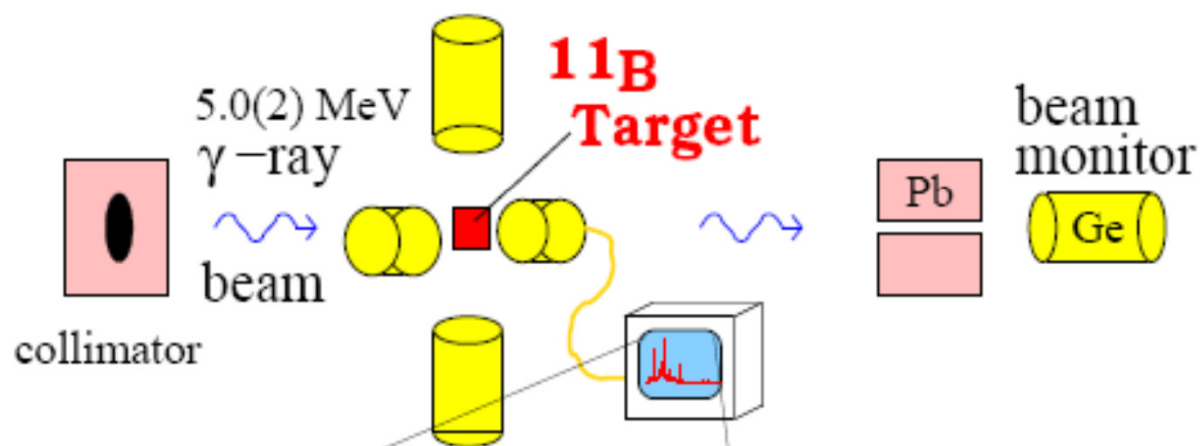


Flux at target: $10^7/\text{s}$
at maximum: $10^5/(\text{s keV})$
Resolution: **3%**
(with 1" collimator)



N.Pietralla et al.
Nucl.Instrum.Methods A 483, 556 (2002).

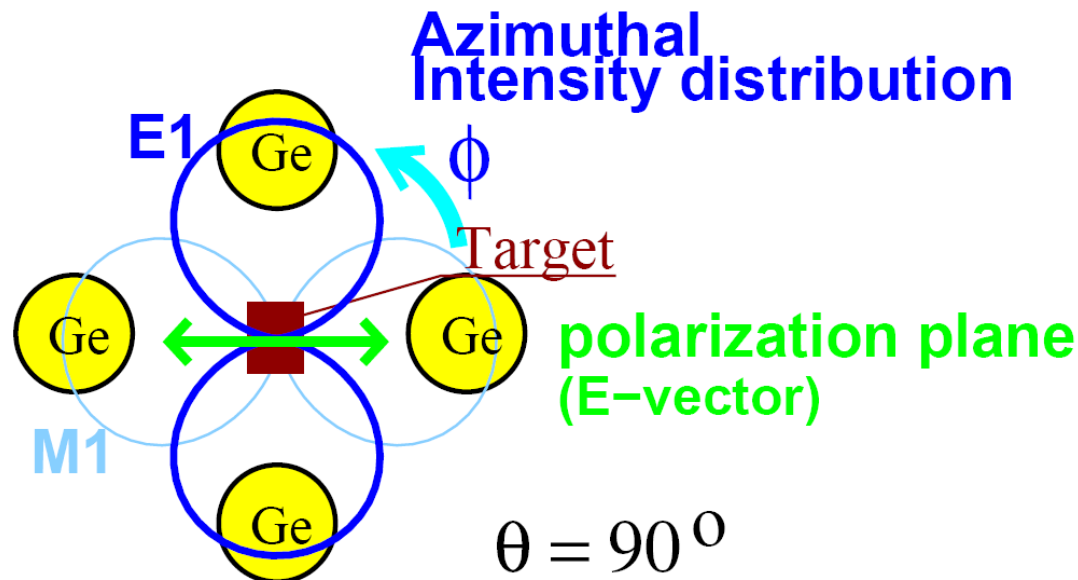
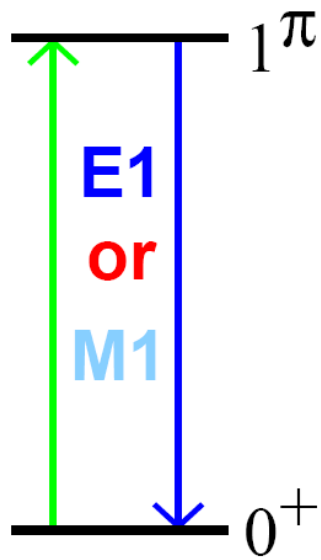
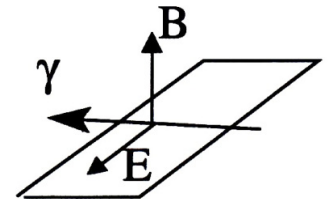
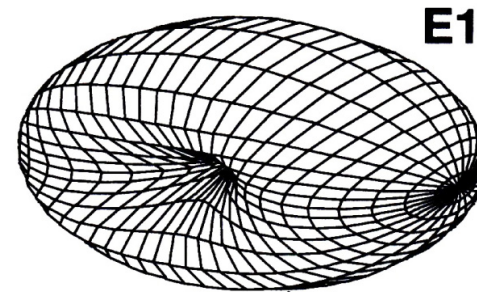
Looking at the Target

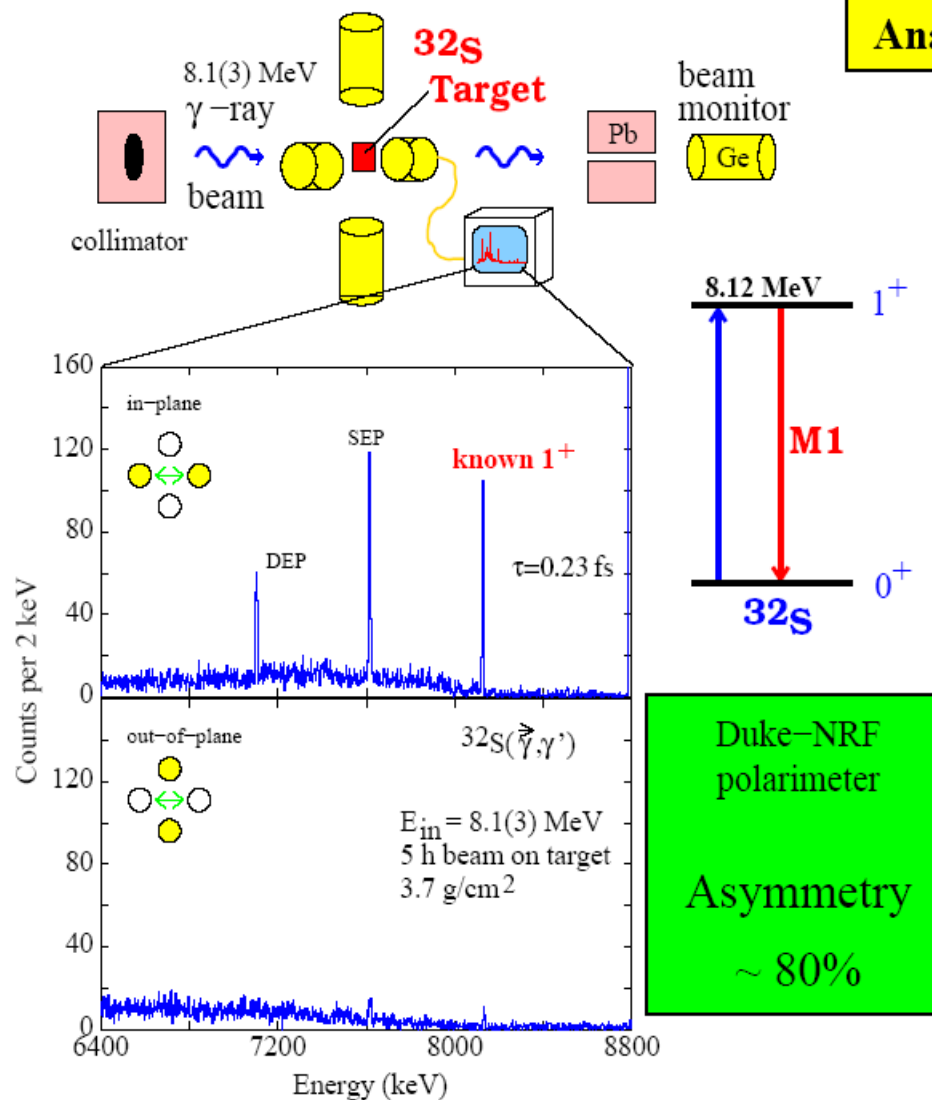


N. Pietralla et al.
Nucl. Instrum. Methods A483, 556 (2002).

Polarimetry with linearly polarized γ -ray beams

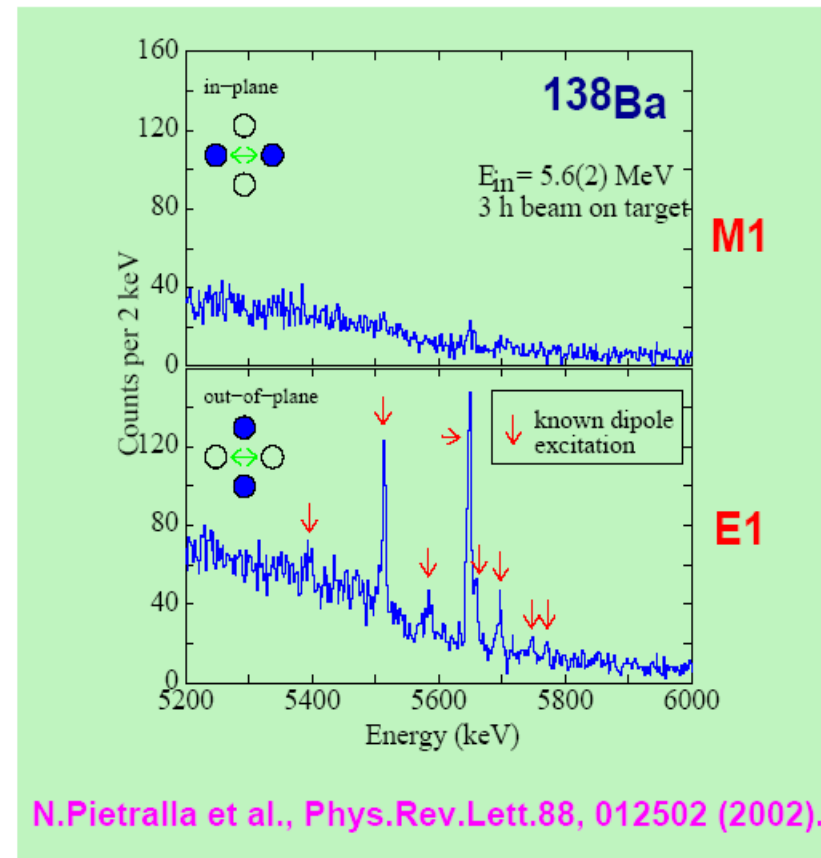
Scattering intensity distribution for linearly polarized photons





Analyzing Power for the Pygmy Resonance

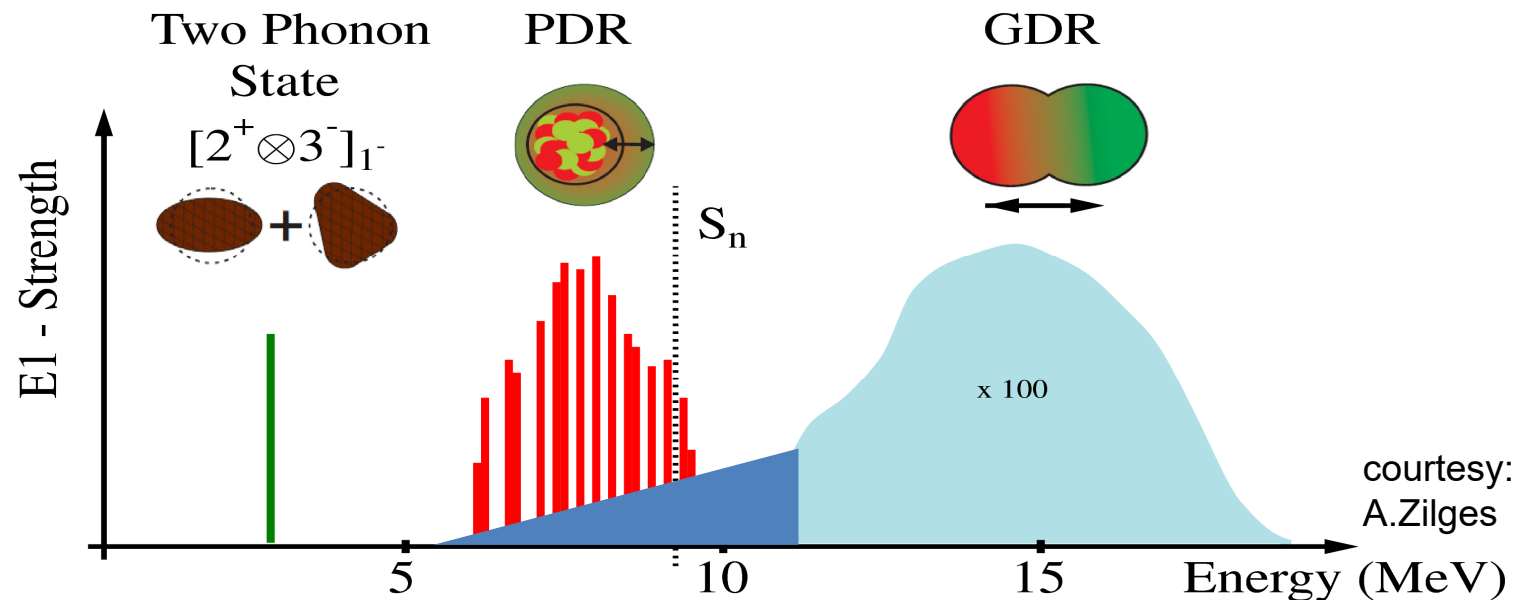
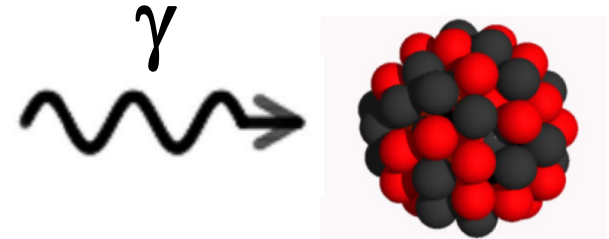
"pygmy resonance": all E1 !



N.Pietralla et al., Nucl.Instrum.Methods A483, 556 (2002).

Electromagnetic dipole-response

Typical E1 strength distribution in heavy atomic nuclei



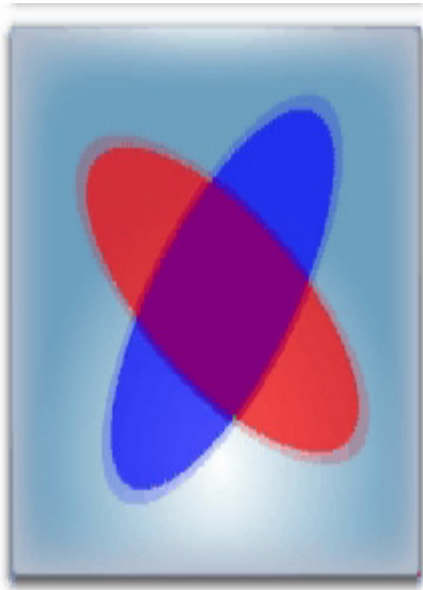
Discovery of Scissors mode: deformed nuclei



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Darmstadt, 1983

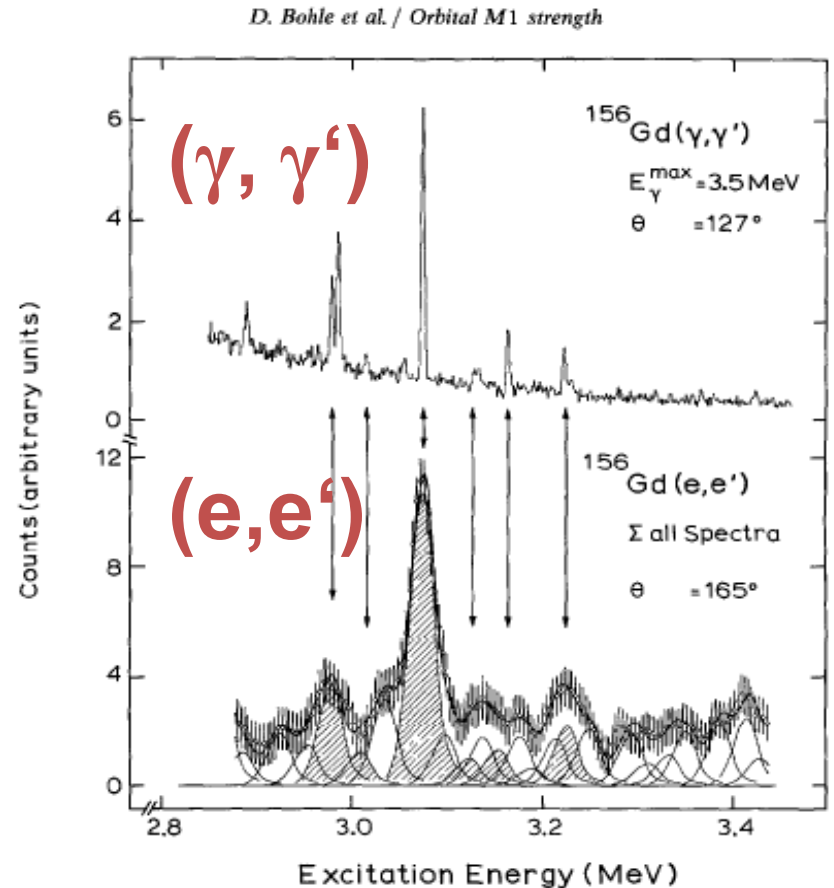
Achim Richter



Predicted by

Lo Iudice & Palumbo, 1978

@ 3 MeV: Iachello, 1981 in IBM-2



Bohle et al., NPA 458, 205 (1986).

DALINAC data

E2 decay strength of the Scissors mode

PRL 118, 212502 (2017)

PHYSICAL REVIEW LETTERS

week ending
26 MAY 2017

E2 decay strength of the M1 scissors mode of ^{156}Gd and its first excited rotational state

T. Beck,^{1,*} J. Beller,¹ N. Pietralla,¹ M. Bhike,² J. Birkhan,¹ V. Derya,³ U. Gayer,¹ A. Hennig,³

J. Isaak,^{4,5,†} B. Löher,^{4,5} V. Yu. Ponomarev,¹ A. Richter,¹ C. Romig,^{1,‡} D. Savran,^{4,5}

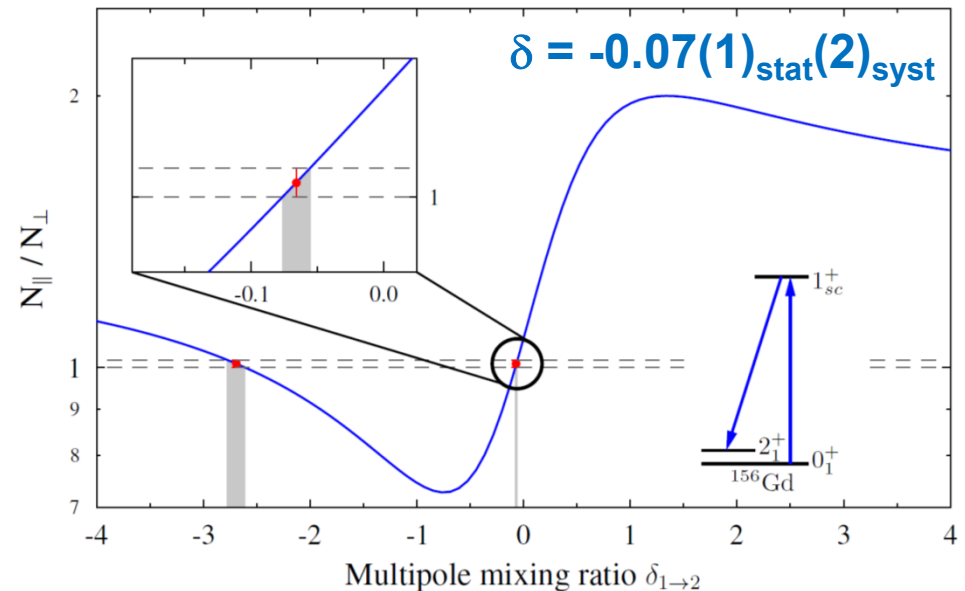
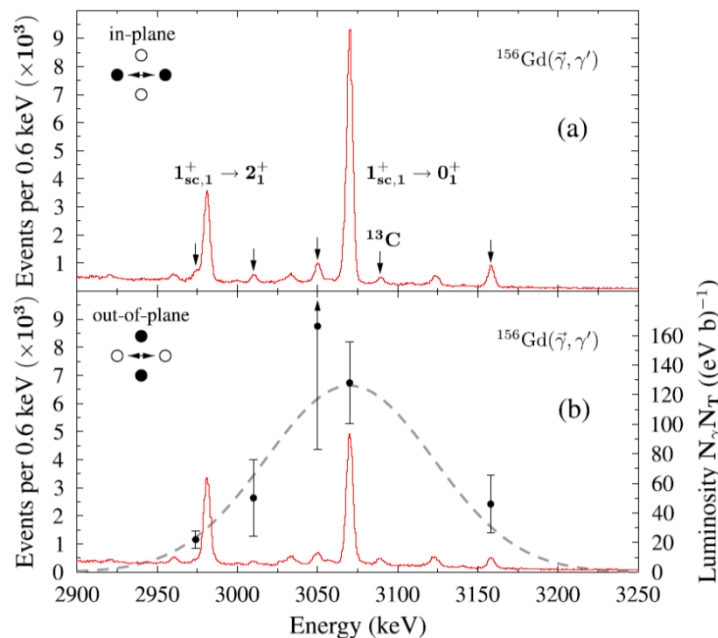
M. Scheck,^{1,6,7} W. Tornow,² V. Werner,¹ A. Zilges,³ and M. Zweidinger¹

¹Institut für Kernphysik, TU Darmstadt, Schlossgartenstr. 9, D-64289 Darmstadt, Germany

²Department of Physics, Duke University and Triangle Universities Nuclear Laboratory, Durham, North Carolina 27708-0308, USA

³Institut für Kernphysik, Universität zu Köln, Zùlpicher Str. 77, D-50937 Köln, Germany

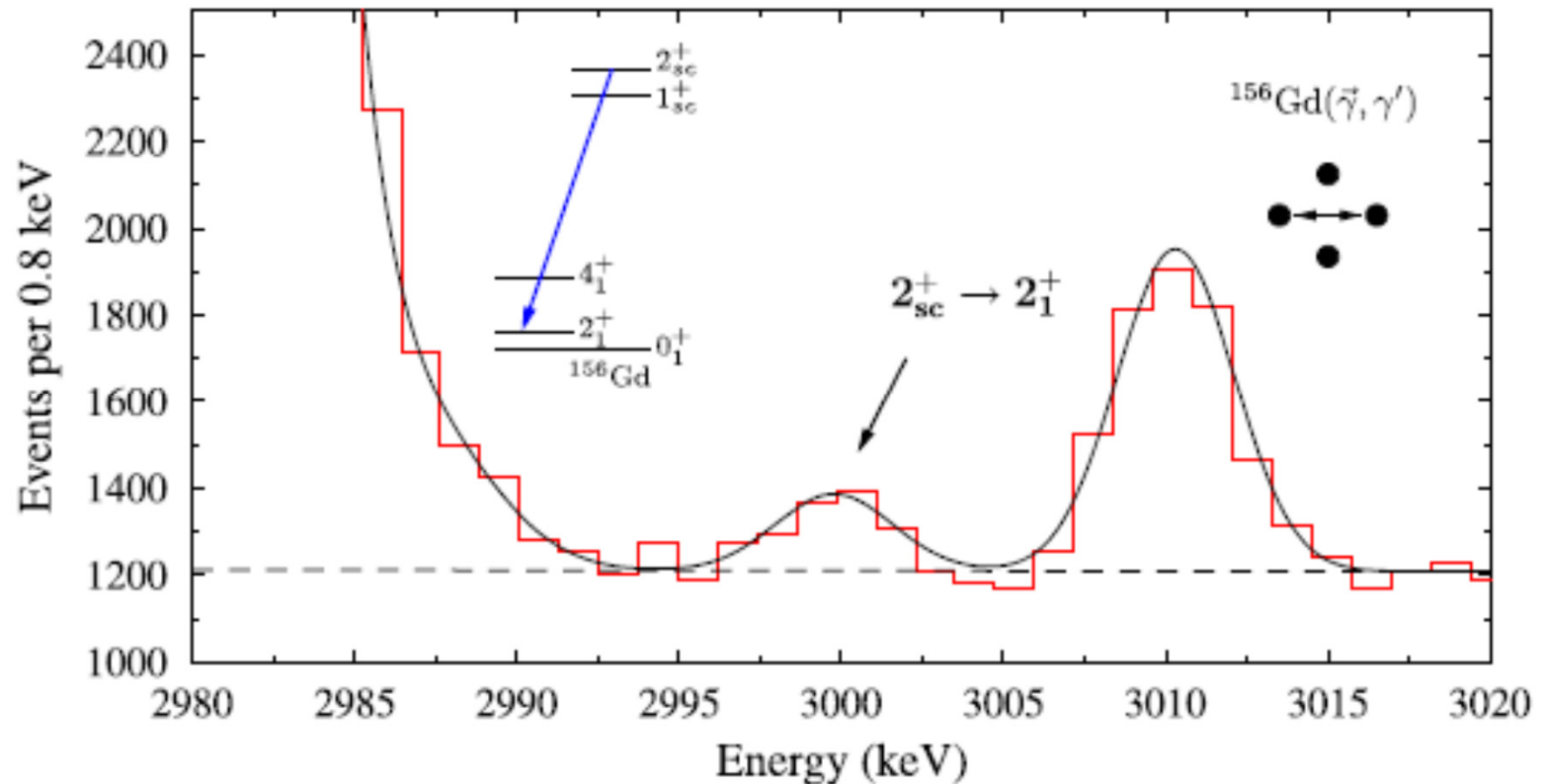
$$B(E2; 1^+_{\text{scM}} \rightarrow 2^+_1) = 0.037(19) \text{ W.u.}$$



S-DALINAC & HIγS

Evidence for rotational excitation of the Scissors Mode

S-DALINAC & HI γ S data



T. Beck et al., Phys. Rev. Lett. 118 212502 (2017).

Supernova 1987A

22.02.1987

Sanduleak-69 202

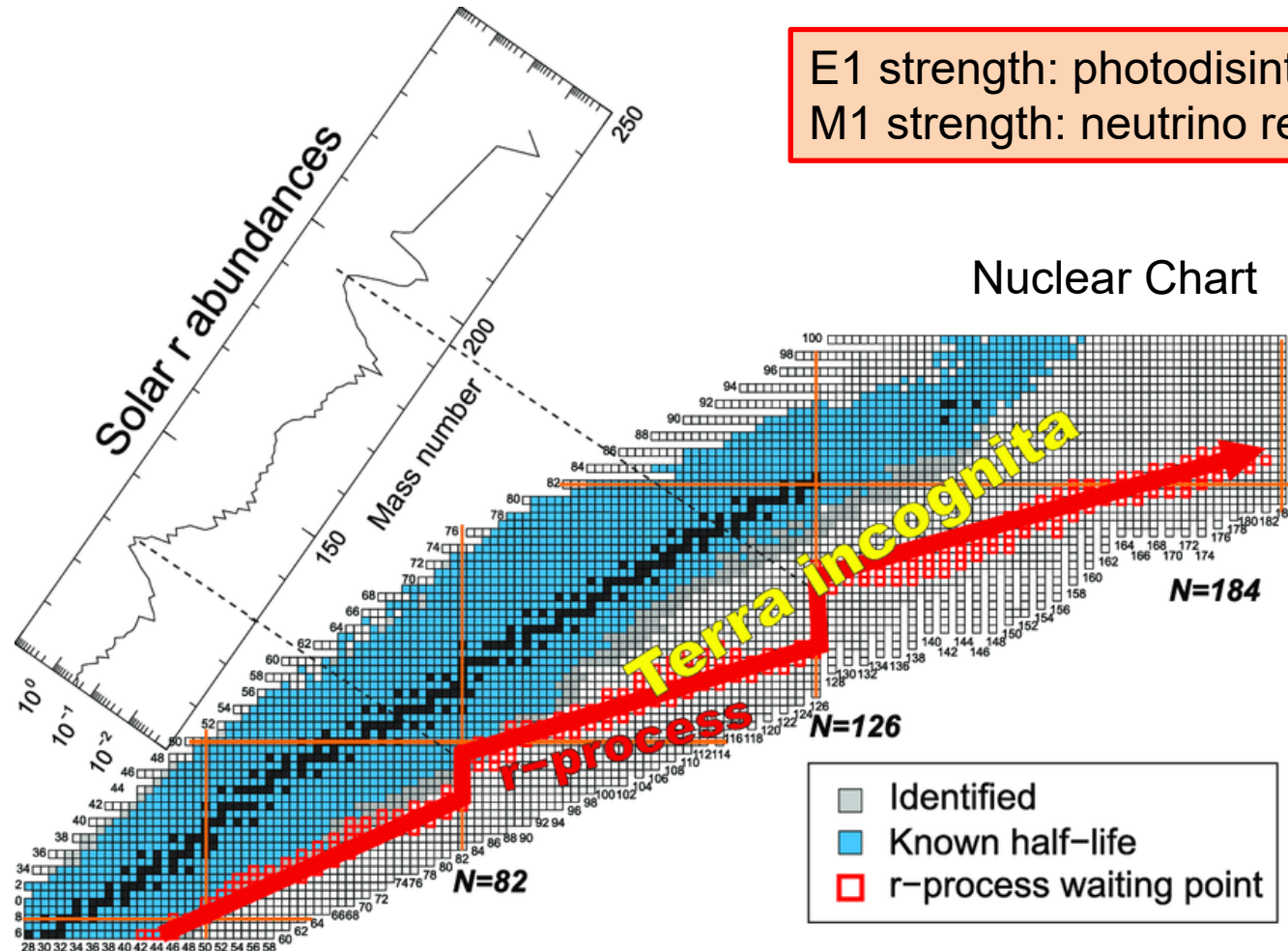


**Great Magellanean Cloud
160.000 lightyears away**

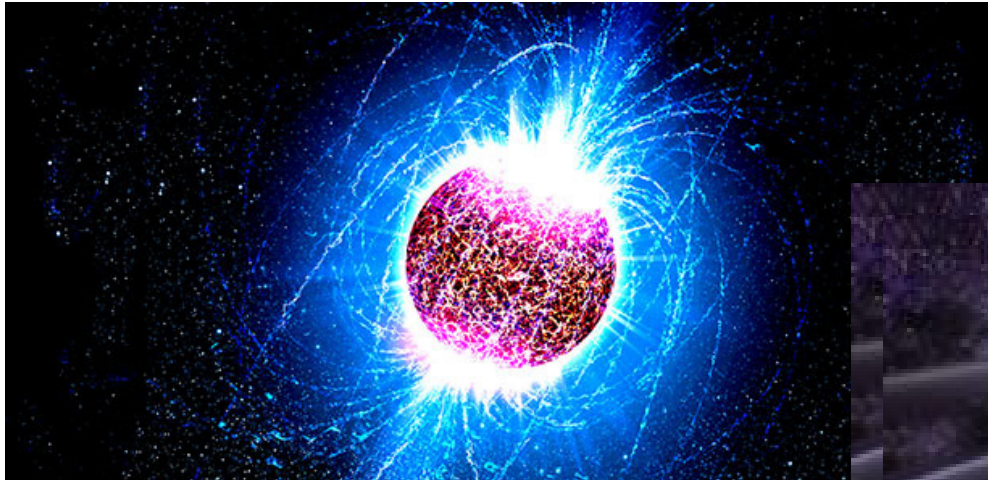
23.02.1987



Cosmic Nucleosynthesis



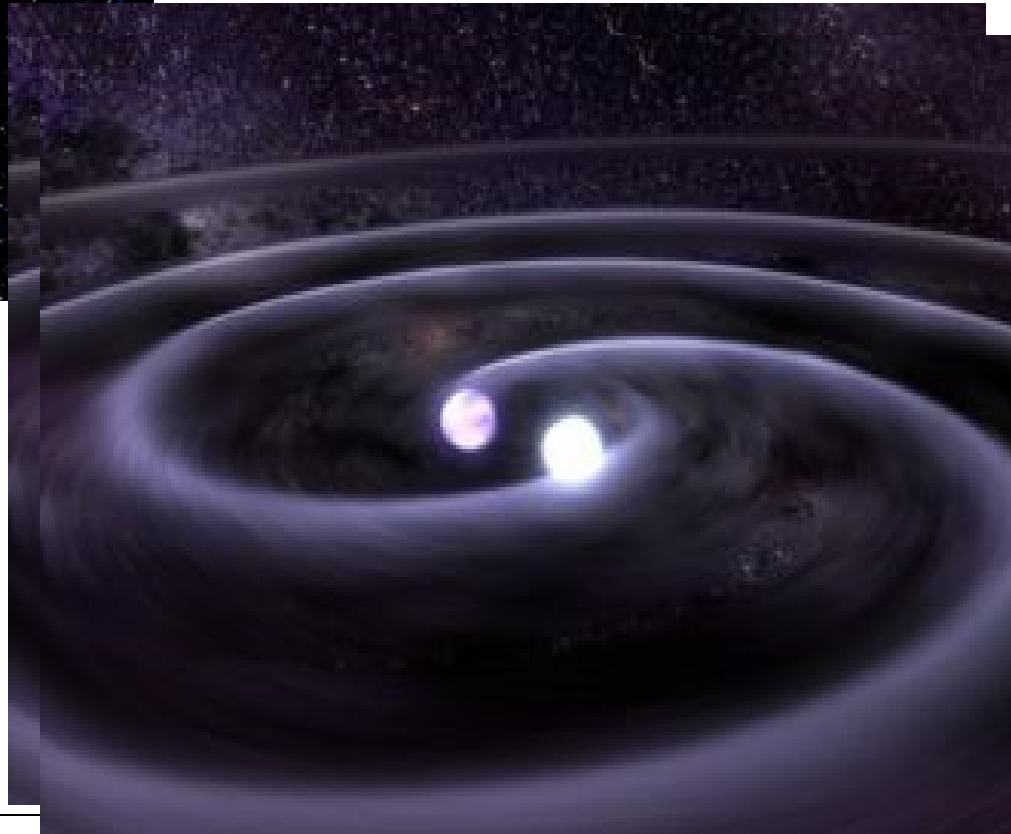
Neutron-Star Mergers



Artist's views !

Abundant source of neutrons
for nucleosynthesis

Neutron-star merger:
Explosive ejection of neutron-
rich matter into space

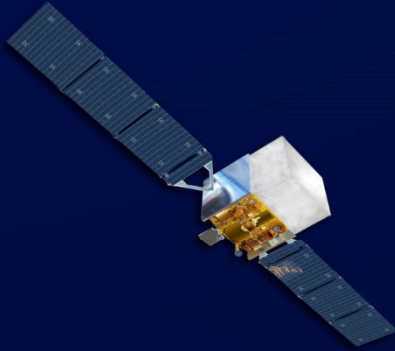


Multi-Messenger Observation of N-Star Merger



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Fermi



Gamma rays, 50 to 300 keV

GRB 170817A



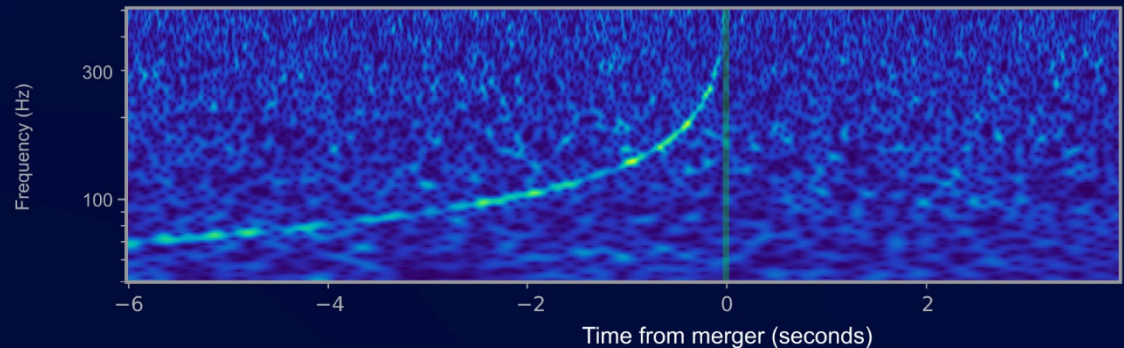
LIGO



Gravitational-wave strain

GW 170817

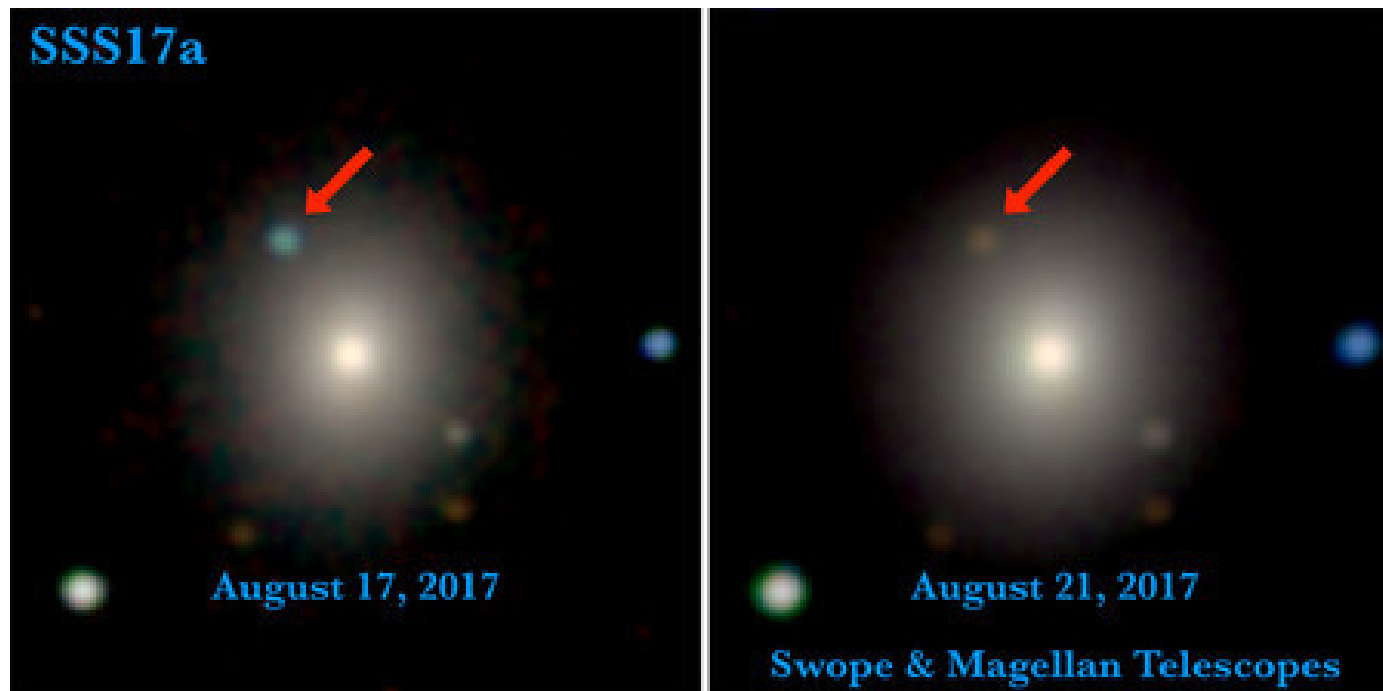
17.8.2017



LIGO/VIRGO – collaboration, Fermi Satellite

Neutron-Star Merger in the Visible Spectrum

A star changes colour in 4 days ! „Kilonova“

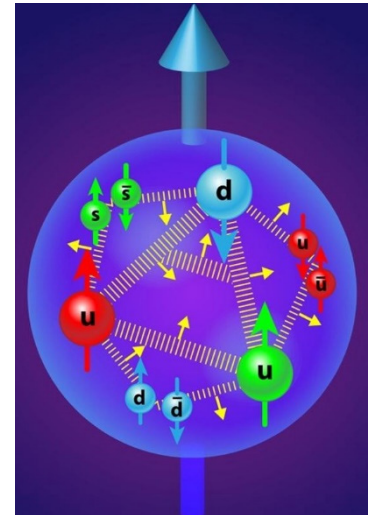
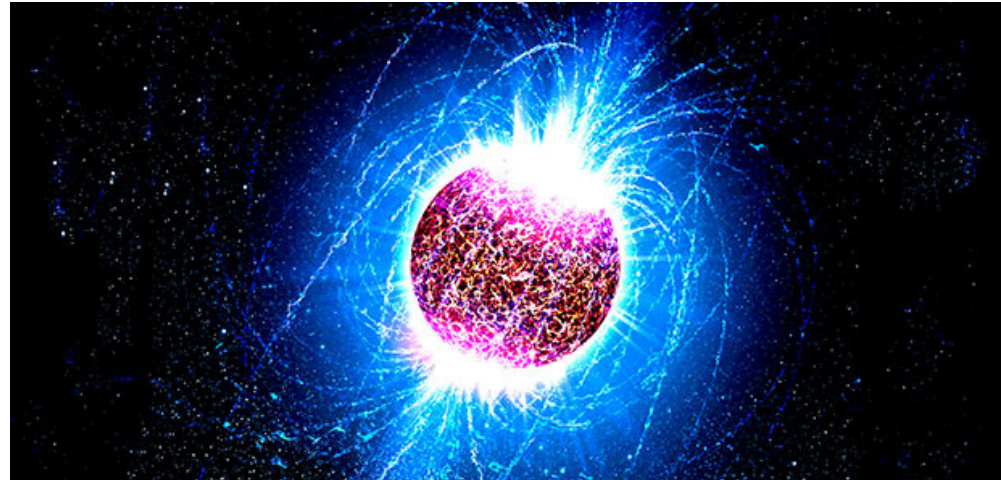


Light curve of neutron-star mergers due to synthesized rare-earth nuclei

B. D. Metzger et al., Mon. Not. Roy. Ast. Soc. 406, 2650 (2010).

Motivation for Nuclear Photonics

- Photons provide a sensitive probe for nuclear structure
- Properties of nuclear isotopes as a function of mass (neutron number)
- Low-energy frontier of Quantum Chromo-Dynamics
- Photonuclear reactions impact the structure and dynamics of stars
- Formation of chemical elements in the Universe
- Nuclei as detector material for neutrino experiments and searches beyond the Standard Model
- → **Nuclear photonics**



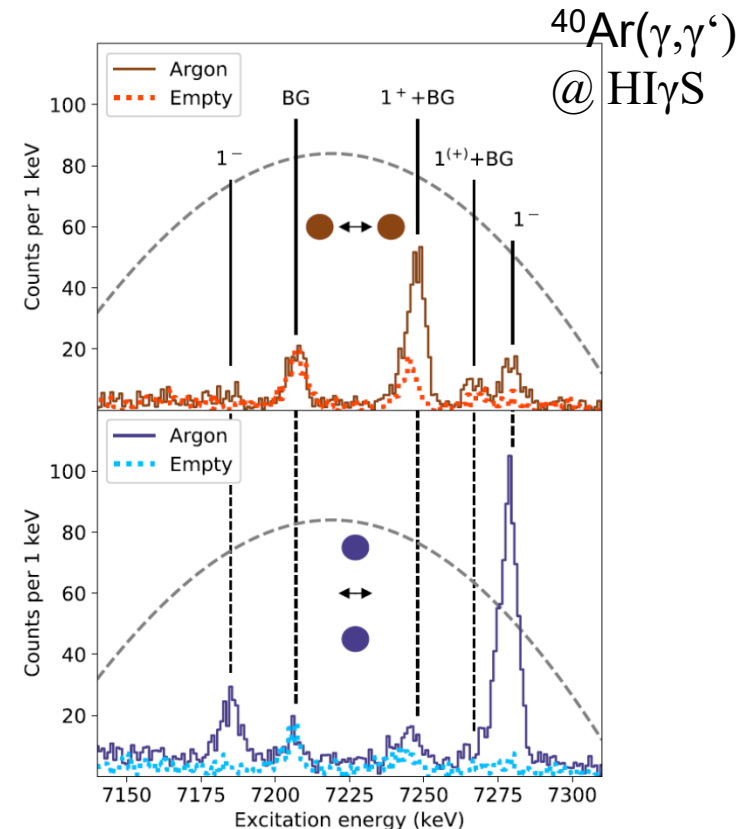
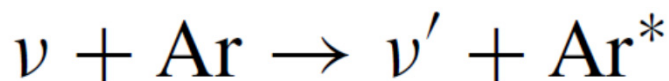
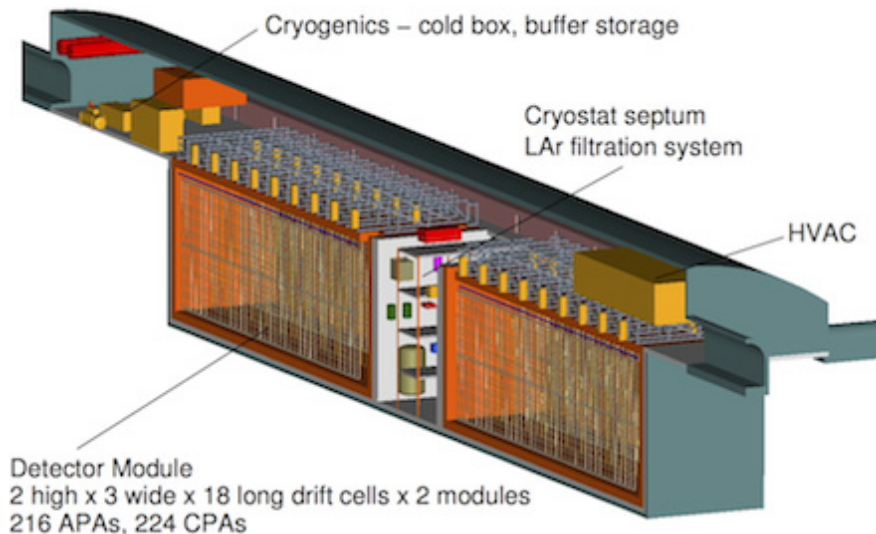
Nuclear Structure for Neutrino Studies



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Deep Underground Neutrino Experiment

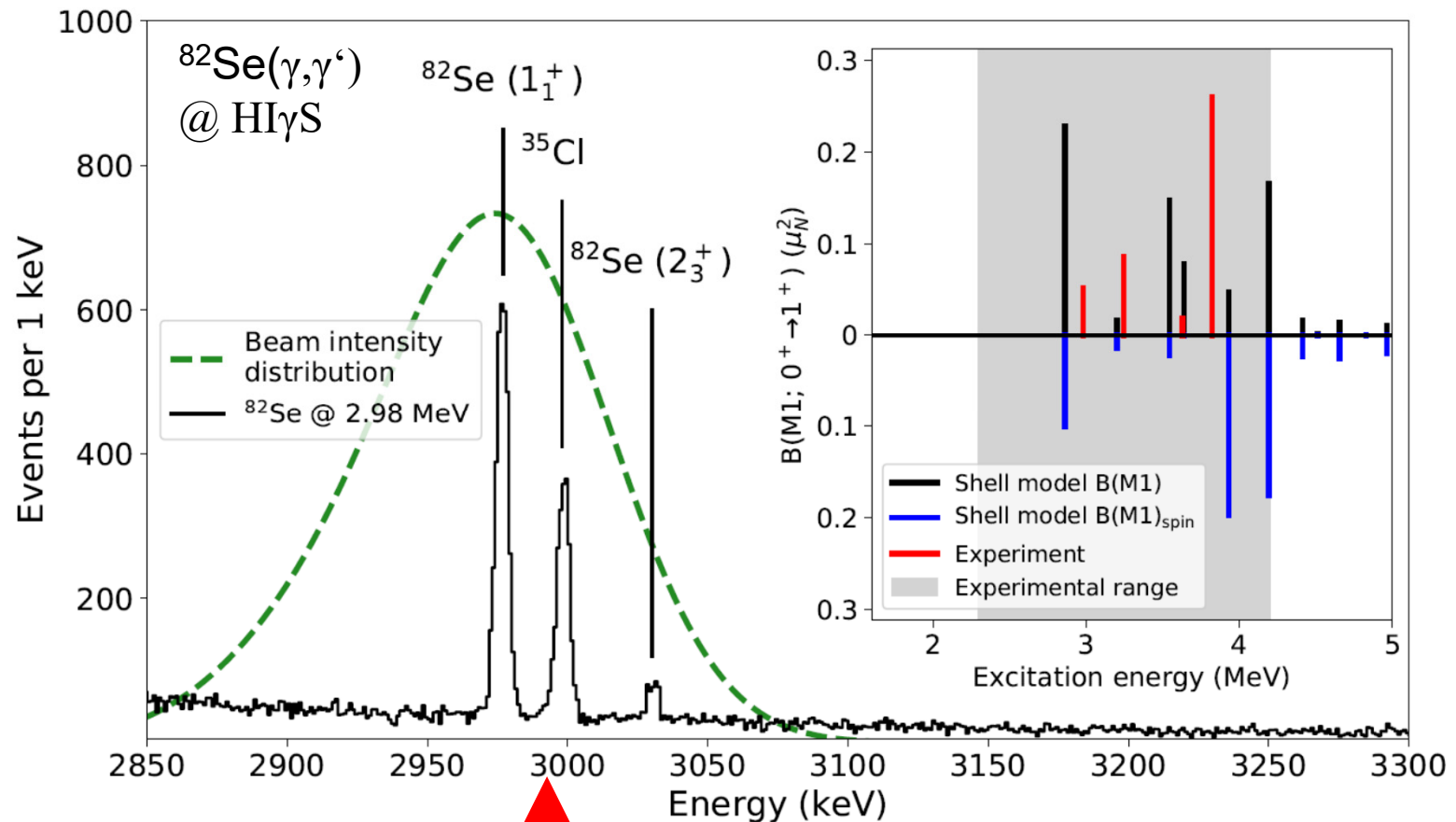
DUNE LAr – detector for cosmic ν 's



U.Gayer et al., Phys. Rev. C 100, 034305 (2019)

→ total low-energy neutrino cross section

Nuclear Structure for $0\nu\beta\beta$ searches



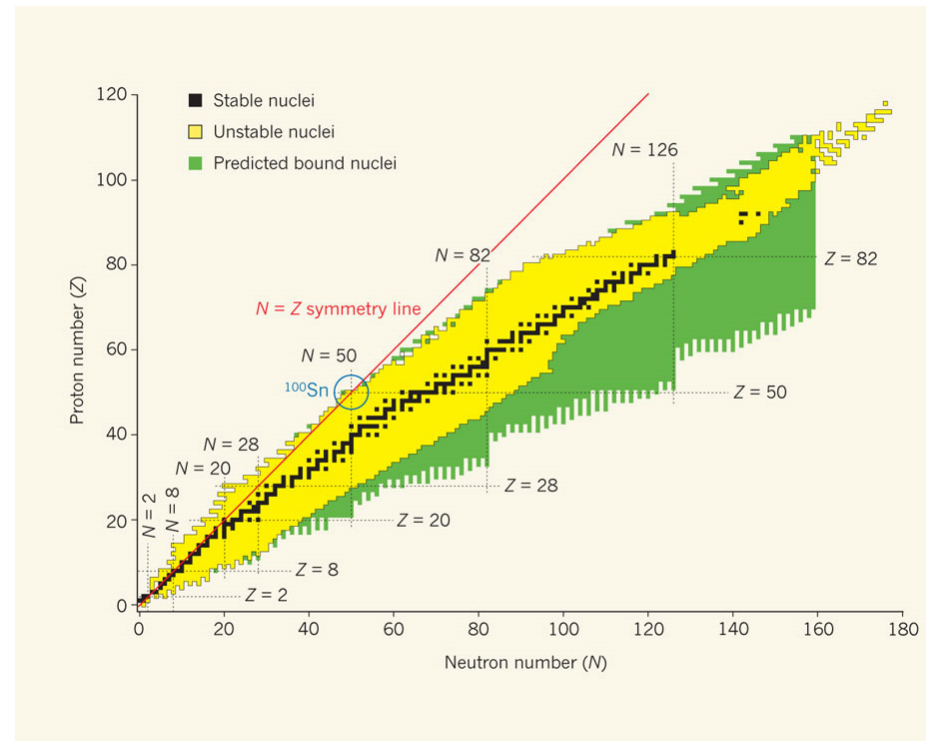
SuperNEMO

$$Q_{\{0\nu\beta\beta\}} = 2995 \text{ keV}$$

V.Werner et al. (TU Darmstadt, 2019)

Physics with Photon Beams

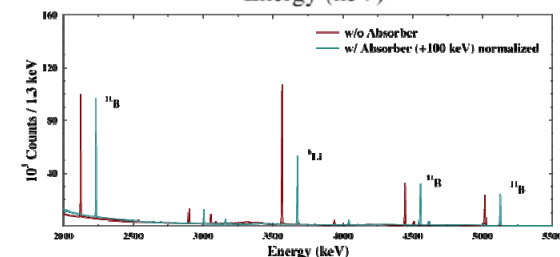
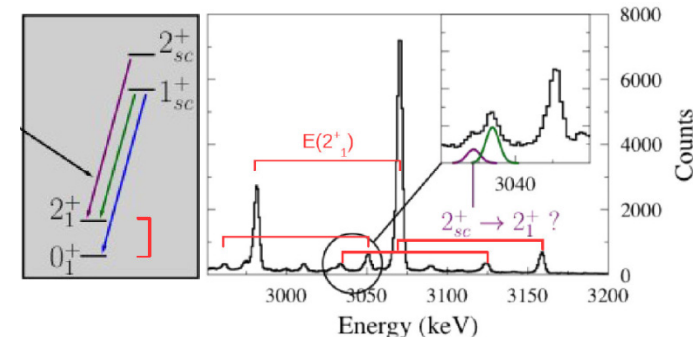
- Nuclear Structure Physics
 - Nuclear single particle structure
 - Collective nuclear structures
 - Photofission
- Particle-Physics Metrology
 - Neutrino detectors
 - Nuclear matrix elements for $\beta\beta$ -decay
- Nuclear Astrophysics
 - Capture / desintegration reactions
 - Nuclear synthesis
- Applications
 - Radiotomography of fuel rods
 - Cultural heritage, etc.



„Discovery Frontier“ for Nuclear Photonics at 4th Generation γ Source

High-Intensity Frontier = „Discovery Frontier“ (scientific opportunities)

- „Availability Frontier“
Access to rare isotopes,
→ broader „nuclear gene pool“
- „Sensitivity Frontier“
weak channels: strong physics
- „Precision Frontier“
high count rates, new methods



Thank you very much!